

# Faulting processes during early-stage rifting: seismic and geodetic analysis of the 2009–2010 Northern Malawi earthquake sequence

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## SUMMARY

In December, 2009, a rare sequence of earthquakes initiated within the weakly extended Western Rift of the East African Rift system in the Karonga province of northern Malawi, providing a unique opportunity to characterize active deformation associated with intrabasinal faults in an early-stage rift. We combine teleseismic and regional seismic recordings of the largest events, InSAR imagery of the primary sequence, and recordings of aftershocks from a temporary (4-month) local network of six seismometers to delineate the extent and geometry of faulting. The locations of ~1900 aftershocks recorded between January and May 2010 are largely consistent with a west-dipping normal fault directly beneath Karonga as constrained by InSAR and CMT fault solutions. However, a substantial number of epicentres cluster in an east-dipping geometry in the central part of the study area, and additional west-dipping clusters can be discerned near the shore of Lake Malawi, particularly in the southern part of the study area. Given the extensive network of hanging wall faults mapped in the Karonga region on the surface and in seismic reflection images, the distribution of events is strongly suggestive of multiple faults interacting to produce the observed deformation, and the InSAR data permit this but do not require it. We propose that fault interaction contributed to the seismic moment release as a series of  $M_w$  5-to-6 events instead of a normal main shock–aftershock sequence. We find the depth of fault slip during the main shocks constrained by InSAR peaks at less than 6 km, while the majority of recorded aftershocks are deeper than 6 km. This depth discrepancy appears to be robust and may be explained by fault interaction. Structural complexities associated with fault interaction may have limited the extent of coseismic slip during the main shocks, which increased stress deeper than the coseismic slip zone on the primary fault and synthetic faults to the east, causing the energetic aftershock series. There is no evidence of deformation at the Rungwe volcanic province ~50 km north of the earthquake sequence between 2007 and 2010, consistent with previous interpretations of no significant magmatic contribution during the sequence.

**Key words:** Radar interferometry; Africa; Earthquake source observations; Seismicity and tectonics; Continental tectonics: extensional.

## 1 INTRODUCTION

An unusual sequence of earthquakes occurred in December 2009 near the town of Karonga, Malawi. Only fourteen earthquakes with a magnitude  $>5$  have occurred around the northern Lake Malawi rift between 1973 and 2017 (NEIC catalogue), and nine of those took place over a 2-week period between 6 and 19 December, 2009 (Fig. 1). More than 200 000 people were affected by these earthquakes, including four deaths and almost 200 injuries (Office of the United Nations Residence Coordinator 2009). Moreover, the sequence caused the evacuation of thousands of households for months, given the prolonged duration of the earthquake activity (Office of the United Nations Residence Coordinator 2009).

The earthquakes are scientifically interesting for several reasons. First, the earthquakes did not occur on one of the  $>100$ -km-long border faults that bound the  $\sim 50$ -km-wide sedimentary basins and elongate lakes (e.g. Lake Albert, Lake Tanganyika and Lake Malawi) of the Western Branch of the East African rift system (EARS, e.g. Ebinger *et al.* 1991, 1999). The primary border fault in this segment of the EARS, the Livingstone Fault, is on the east side of the lake, 40–50 km northeast of the events, and thus the locations, mechanisms and event depths during 2009–2010 imply that the earthquakes are occurring due to slip on normal faults within the hanging wall (Fig. 1). As nearly all moderate-size earthquakes within the Western Rift of the EARS appear to be on major border faults (e.g. Jackson & Blenkinsop 1993; Nyblade & Langston 1995), the 2009 events offer a rare opportunity to evaluate how rifting is accommodated within the hanging wall of the system. A second unusual characteristic of this earthquake sequence is that in contrast to a classical main shock–aftershock sequence, the Karonga earthquake series comprises three large events with similar magnitudes of 5.8–6.0, with the largest event being the last one on 19 December 2009. Including two foreshocks, there were 14 other events with magnitudes in excess of 4.5, with all events in this primary sequence occurring in a 27-d period. Subsequent aftershock activity continued after 19 December 2009, but at a much lower magnitude level.

There are several possible explanations that can account for these earthquakes in the hanging wall and their seismological character: (1) a single immature fault in the hanging wall (e.g. Biggs *et al.* 2010) with a heterogeneous distribution of stress and/or frictional properties (e.g. King & Nábělek 1985; Hillers *et al.* 2007; Lohman & McGuire 2007; Hamiel *et al.* 2012); (2) the interaction of multiple hanging wall faults with each other and/or with the border fault [for example, Hamiel *et al.* (2012) modelled the fault to have seven segments] or (3) a magmatic or fluid-driven swarm. The primary goal of this work is to determine which of these three models, or what combination of these three, is most plausible.

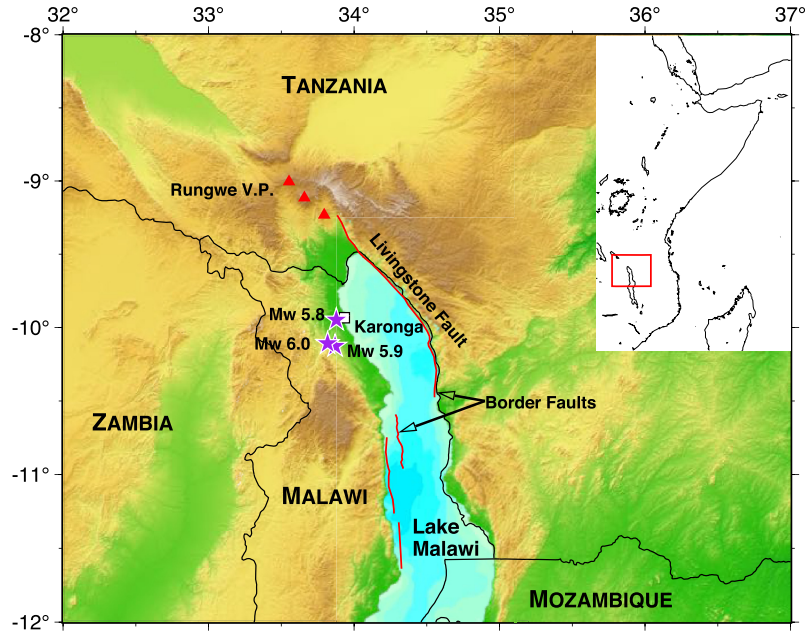
Numerical and analogue models suggest that once faults link to form favourably oriented, laterally continuous structures, they are capable of accommodating a large portion of deformation (e.g. Cowie *et al.* 2000). For this reason, understanding the relative maturity of hanging-wall fault(s) associated with the Karonga earthquake is significant. Quantifying the maturity and behaviour of intrabasin faults is also important for characterizing the accommodation of extension, either through distributed extension, or through hanging wall flexure.

Biggs *et al.* (2010) propose that the 2009 earthquakes occur on a single, immature, previously unmapped, west-dipping normal fault located near Karonga, which likely connects to a series of surface ruptures (Macheyeki *et al.* 2015) identified by the Malawi Geological Survey as the St Mary's Fault or St Mary Fault by

Kolawole *et al.* (2018). The surface location of this structure, and its westward dip, are both well constrained by the surface deformation observed in the Interferometric Synthetic Aperture Radar (InSAR) analysis (Biggs *et al.* 2010; Hamiel *et al.* 2012). A single immature fault within the hanging wall could generate a sequence of events with similar magnitudes because it might have a more irregular large-scale geometry, such that ruptures are modulated by fault bends (e.g. King & Nábělek 1985). Furthermore, since young faults have accommodated less total slip, properties along the fault are likely to be more variable, including the presence of more asperities, which could discourage rupture of a large portion of the fault in a single event (e.g. Hillers *et al.* 2007). Biggs *et al.* (2010) suggest that slip on a single fault can explain the pattern of surface deformation captured in InSAR data, but with slip during the final  $M_w$  6.0 event occurring on a different portion of the fault than slip during the preceding events, which implies a rough, immature fault.

In the second model, a series of secondary faults within the hanging wall could interact with one another and/or the border fault to cause a sequence of events with similar magnitudes. Slip on one fault can cause loading and subsequent failure on another fault (e.g. Stein 1999). In the Karonga basin, this interaction could involve either a number of secondary faults or the border fault. Existing constraints on the geometry of secondary faults imply that fault interaction here could be important. Seismic reflection data collected in the lake to the east of the epicentral region image a series of closely spaced ( $\sim 5$  km), west-dipping intrabasin faults (Flannery & Rosendahl 1990; Mortimer *et al.* 2007). West of the epicentral region, the east-dipping Karonga Fault has a clear topographic expression and appears to be one of the most prominent faults within the hinge zone of the flexing hanging wall (e.g. Laó-Dávila *et al.* 2015). The close spacing and opposing dips of the Karonga and St Mary's Faults implies that they intersect or otherwise interact at a fairly shallow depth within the crust (upper 10 km, e.g. Biggs *et al.* 2010). Finally, pre-existing faults and basement fabrics have been recognized within the study area (e.g. the Precambrian Mughese Shear Zone) and may influence and interact with modern rift faults (e.g. Kolawole *et al.* 2018). The possible participation of other fault structures in the deformation are not well determined from existing studies (e.g. any slip on segments beneath the lake will be invisible to InSAR).

For the third model, although there is no surface volcanism directly within the epicentral region, the potential role of deep magmatic fluids in the area is debated. There are hot springs in northern Malawi (Dulanya 2006), and the active Rungwe volcanic province is 50–70 km to the north (Fig. 1). Rungwe is the most southerly surface expression of volcanism in the EARS and one of a handful of volcanic provinces in the Western Rift, which is less magmatic than other parts of the EARS (e.g. Ebinger *et al.* 1989). The only previous earthquake sequences in the southern EARS that exhibit similar swarm characteristics are demonstratively volcanic. These include eight events with  $M_w > 5$  (the largest event occurring third in the sequence) associated with volcanic diking and subsequent eruption in northern Tanzania (e.g. Calais *et al.* 2008; Fischer *et al.* 2009), and three  $M_w$  5.1–5.2 events over a 3-d period associated with a major eruption at Nyiragongo volcano eruption near Lake Kivu (Shuler & Ekström 2009). Previous InSAR results of the Karonga sequence (Biggs *et al.* 2010; Hamiel *et al.* 2012) imply that the majority of the moment release is on shallow tectonic faults, and there is no evidence for dyke or fluid involvement in the sequence. We have processed InSAR data over a larger spatial area including the Rungwe volcanic province to further constrain the potential for volcano–tectonic interactions during this earthquake sequence.



**Figure 1.** (a) Tectonic framework of the 2009 Karonga earthquakes on SRTM3 topography, with the Rungwe Volcanic Province shown as red triangles (Ebinger *et al.* 1987, 1989), and red lines showing trace of primary border faults (Mortimer *et al.* 2007; Accardo *et al.* 2018). Stars show the locations of the largest earthquakes of the 2009 sequence from the NEIC catalogue; CMT locations for these events are shown in Fig. 2. White box indicates town of Karonga. (Inset) Red box shows study area in context of the East Africa rift system. Lake bathymetry is estimated from a surface fit to legacy MCS data (Lyons *et al.* 2011).

To distinguish between these models, we present an analysis of data recorded with a small network of seismometers deployed for four months following the earthquake swarm (January–May 2010) to characterize the spatial and temporal distribution of seismicity. Further, we compare these results with new models of the 2009 main shocks including ground deformation from InSAR and analysis of teleseismic and regional seismic waves. Our combined seismic and InSAR analysis in this paper will evaluate the three models for why the 2009 earthquakes near Karonga did not follow a normal main shock–aftershock sequence.

## 2 DATA AND METHODS

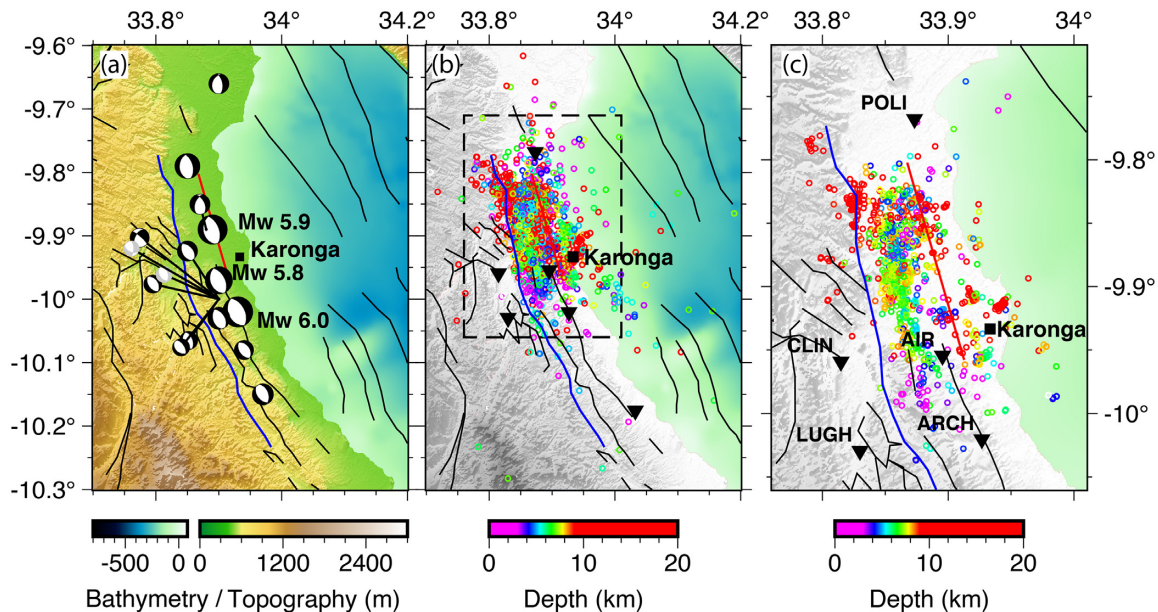
### 2.1 Local aftershock array

At the time of the 2009 earthquake sequence, the nearest seismometers (from AfricaArray, AA) were separated by several hundred kilometers and only one of these stations was working in Malawi, in Zomba over 600 km to the south. The Geological Survey of Malawi requested assistance in monitoring the ongoing sequence, and we were able to secure five instruments for rapid deployment, including three RAMP (Rapid Array Mobilization Program) instruments from IRIS ([www.iris.edu](http://www.iris.edu)) and two additional instruments from Cynthia Ebinger. The five instruments were rushed to the Karonga region and deployed for a four-month period between 6 January to 6 May 2010. Combined with a relocated AA station in Karonga, this network provides up to six stations within an approximately  $20 \times 50$  km region surrounding Karonga (Fig. 2). The five portable stations are intermediate-period, three-component seismometers that recorded at 100 Hz, while the AA station is a broadband sensor that recorded at 20 Hz. Continuous seismic data are available under network code YL2010 at the IRIS Data Management Center (Gaherty & Shillington 2010).

All collected seismic data were run through an automated detection and association algorithm to provide an initial set of possible events. Each event was visually inspected and  $P$  and  $S$  arrival times manually repicked by analysts. For all events with a minimum of four local stations providing traveltimes, the events were located using the hypoinverse location algorithm (Klein 2002) and several possible seismic velocity models. 1-D  $P$ -velocity models were derived for the Karonga region from a published model for an analogous portion of the southern EARS (Kim *et al.* 2009), and preliminary analyses based on active-source and local earthquake data collected during the SEGMENT experiment (Oliva *et al.* 2016; Shillington *et al.* 2016; Accardo *et al.* 2018). We established  $V_p/V_s$  of 1.70 based on a trial-and-error analysis of residual behaviour. The resulting catalogues were evaluated based on overall mean and median traveltime residuals and estimated error ellipses, and the stability and number of events that had good solutions. The results shown in the figures are based on the preferred velocity model (dubbed SEG4), but some of the depth and statistical behaviour of the other models are presented in the results below. In general, the great bulk of the seismicity is located within the footprint of the six-station local array (Fig. 2b). Approximate local magnitudes ( $M_l$ ) were determined using an automated estimate of displacement amplitudes in the direct  $P$  and  $S$  arrival windows. The preferred hypoinverse catalogue contains 1911 events with  $M_l$  ranging from  $-1$  to 4.6, a mean RMS residual of 0.06 s, and a median location error (ellipse long axis) of 2 km.

Using this catalogue as a starting point, we performed double-difference relocation of the events using the hypoDD algorithm (Waldhauser & Ellsworth 2000) (Figs 2c, 3, Fig. S1). We utilized station-specific velocity models based on SEG4 with local correction for station elevation and sediment thickness. For events with  $<5$  km separation difference, differential  $P$  and  $S$  times were calculated from the hand-picked traveltimes. We also incorporated cross-correlation differential times for event pairs with  $<3$  km separation





**Figure 2.** (a) CMT solutions for the 16 largest events that occurred in November–December 2009 (Table 2). Events are scaled by magnitude ( $M_w$  4.5–6.0), with the largest three events labelled. Nine events from the Global CMT catalogue are plotted at their catalogue locations, with one shifted for clarity. The seven new solutions are indicated by the black lines pointing to the common centroid location used for these events; their relative locations are accurately represented as shown in Fig. S4. The two foreshocks are shown with grey mechanisms. Background topography relative to sea level is derived from the SRTM3 database, and lake bathymetry is estimated from a surface fit to legacy MCS data (Lyons *et al.* 2011) relative to the lake surface (elevation  $\sim 478$  m). (b) Hypoinverse locations for approximately 1900 aftershocks recorded in Jan–May 2010 recorded using the local array (black triangles), with event symbol colours scaled by depth. Topography has been converted to grey scale for clarity. Black box shows zoom region of panel (c). (c) HypoDD relative relocations for  $\sim 1050$  events, coloured by depth. Note change in map scale relative to panels (a) and (b). For all panels, lines show mapped faults (Geological Survey of Malawi 1966; Mortimer *et al.* 2007; Biggs *et al.* 2010), with the Karonga (blue) and St Mary’s (red) Faults highlighted.

difference and correlation coefficient greater than 0.7. Requiring a minimum of eight constraints per pair, the resulting preferred differential-location catalogue contains 1049 linked events within a single cluster. We tested a variety of choices for data weighting and damping, and the final solution weighs the hand-picked and cross-correlation times approximately equally, and weighs the hand-picked  $S$  times half of the hand-picked  $P$  and cross-correlation times. The final locations produce RMS misfit of the hand-picked and cross-correlation differential times of 0.06 and 0.03 s, respectively.

The complete hypoinverse and hypoDD catalogues are available as direct downloads in the supplement.

## 2.2 Geodetic data

### 2.2.1 2009 earthquakes

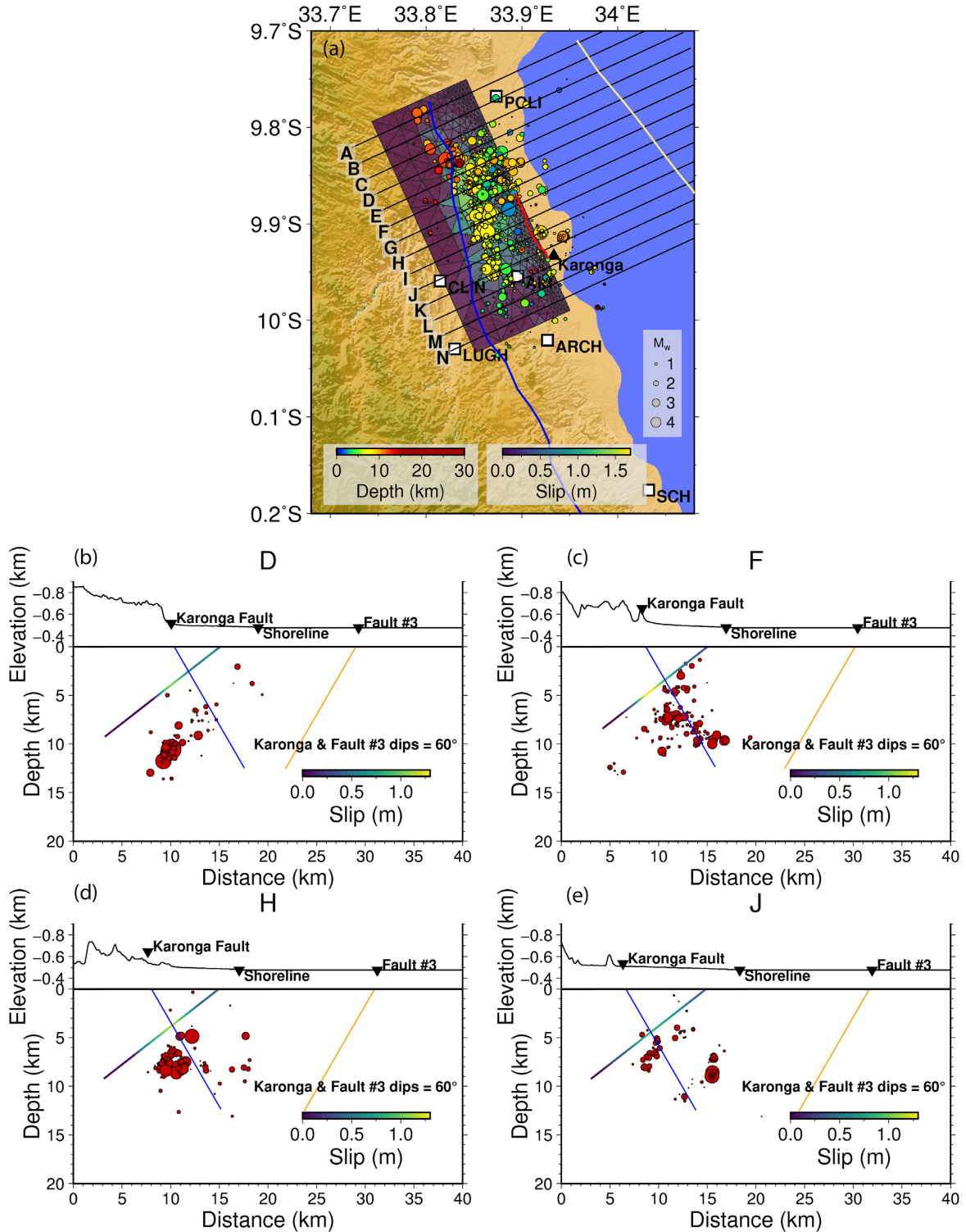
We used InSAR data recorded by the Advanced Land Observation Satellite (ALOS-1) of the Japan Aerospace Exploration Agency and from multiple beams of the Envisat satellite of the European Space Agency spanning the earthquake sequence (Figs 4a–c and Fig. S2). We used the same data as Biggs *et al.* (2010), but we processed them independently, used a different slip inversion strategy, and considered the role of multiple faults. ALOS data are processed using the ISCE software version 170806 (Rosen *et al.* 2012) with topographic effects removed using the Shuttle Radar Topography Mission topography (Farr *et al.* 2007), and all interferograms were unwrapped using SNAPHU (Chen & Zebker 2002). Envisat data are processed using the Caltech/JPL-developed software ROI-PAC (Rosen *et al.* 2004) with the same procedure for topography removal

and unwrapping, which can generate similar interferograms to the ISCE-derived products.

We processed one ALOS-1 pair and five Envisat pairs during the coseismic time span, but only the ALOS-1 interferogram and two Envisat interferograms, which remain coherent around the centre of surface deformation, were used for fault-slip inversion (Table 1). Three other coseismic interferograms from Envisat are not included as the input of inversion, but we compare them with the predictions of the forward model as an independent test (Fig. S2).

All three interferograms used in the inversion are downsampled using a resolution-based quad-tree decomposition algorithm (Lohman & Simons 2005) for achieving computational efficiency. A total of 1476 points (731 from the ALOS interferogram, 397 from the Envisat interferogram 3A and 348 from the Envisat interferogram 4A) are collected to reflect coseismic surface deformation (line-of-sight changes) as the input (Fig. 4d). To determine an estimate of fault location and geometry, we adopt the direct search algorithm that uses neighbourhood approximation in parameter space (Sambridge 1999), assuming a uniform fault dislocation and a homogeneous elastic half-space (Okada 1985). The initial model is set to the best-fitting fault geometry from Biggs *et al.* (2010). Once the optimal estimate of fault plane is found under these assumptions, we set up an automated fault model discretization with distributed slip on triangular dislocations (Barnhart & Lohman 2010) with a depth-dependent cell size that is based on checkerboard tests done by Hamiel *et al.* (2012). A linear growing factor is set up so that the size of a fault patch is 0.5 km when depth is less than 1.5 km and is 3 km when depth is 9 km, which is within resolution capabilities of models as seen from the checkerboard tests. We perform the inversion using all the input points of surface deformation and

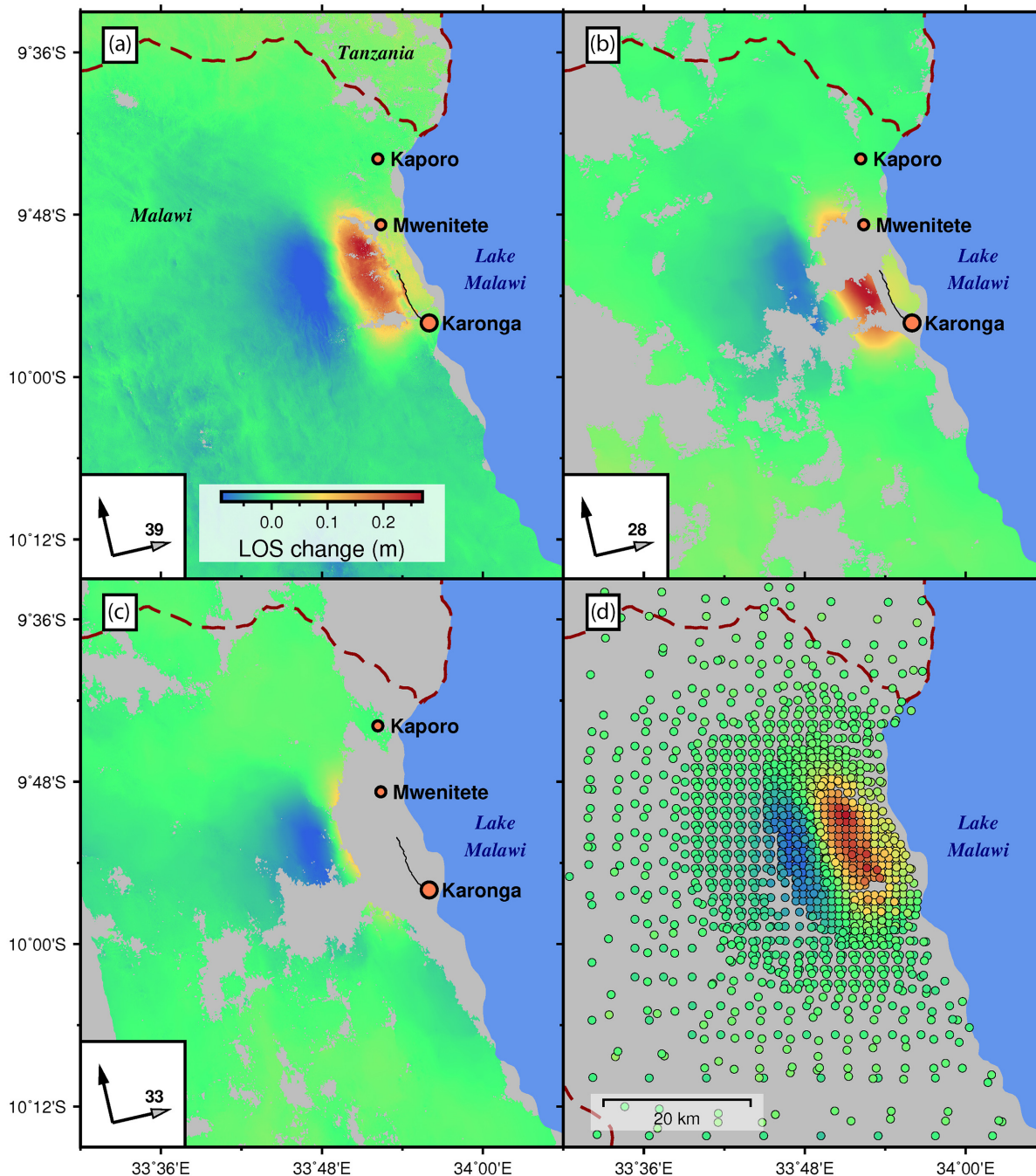




**Figure 3.** A map (subpanel a) and the four NE–SW cross sections D, F, H and J (subpanels b–e) showing inferred St Mary’s slip distribution (coloured by slip) and hypoDD aftershock locations during January–May 2010 (scaled by local magnitude), with offshore faults in yellow (Mortimer *et al.* 2007), the Karonga Fault in blue (Geological Survey of Malawi 1966), and the surface rupture (in red) from the field survey conducted in October 2010 (Macheyeki *et al.* 2015). Six local seismic stations are marked as white squares. All cross sections in (a) are shown in Fig S3.

the estimated fault geometry, assuming a homogeneous elastic half-space (Fig. 5). We use second-order Tikhonov Regularization in our smoothing matrix and determine the weighting parameter using the

$j\mathcal{R}_i$  approach, and evaluate the trade-off between smoothing and model fit using an L-curve (Barnhart & Lohman 2010).



**Figure 4.** The unwrapped interferograms used in the fault slip inversions, and the stacked, downsampled measurements of LOS change. The black line in (a–c) indicates the locations of fault rupture from Macheyeke *et al.* (2015). (a) ALOS-1 between 23 September and 24 December, also showing the locations of major towns and the surface rupture (black line) from Macheyeke *et al.* (2015). Spacecraft heading shown as black arrow and LOS (with angle from vertical) shown as grey arrow. (b) Envisat between 1 December and 5 January. (c) Envisat between 6 September and 24 January. (d) Example of downsampled ALOS-1 interferogram (from a) using resolution-based quad-tree decomposition algorithm.

We also compare the optimal fault geometry to mapped surface ruptures (Macheyeke *et al.* 2015), existing fault locations (Geological Survey of Malawi 1966), and seismic profiles in Lake Malawi (Mortimer *et al.* 2007). Motivated by these additional datasets as well as the aftershock relocations, we test alternative fault geometries based on these constraints, and perform inversions to assess how the slip distribution changes.

To determine the effects on the slip distribution caused by assuming a layered elastic model instead of a homogeneous half-space,

we perform the distributed slip inversion using the preferred seismic velocity model SEG4 as a layered half-space (Fig. S3). While the homogeneous half-space model was calculated using triangular dislocations, for the layered half-space model, we use rectangular dislocations in the software EDGRN/EDCMP (Wang *et al.* 2003). To convert from triangular to rectangular dislocations, we manually discretized the preferred fault plane into 248 rectangular patches, with the size similar to the best triangularly discretized fault model. Parameters for  $S$ -wave velocity and density in EDGRN are derived

**Table 1.** Interferograms used in this study. The first one is from ALOS PALSAR and the rest are from ENVISAT ASAR.

No.	Dates (yymmdd)	Used in fault inversion	Beam number	Orbit	Baseline (m)	Figures
1	090923-091224	Yes	ALOS	Ascending	100	Fig. 4a
2	091201-100105	Yes	3A	Ascending	150	Fig. 4b
3	090906-100124	Yes	4A	Ascending	-70	Fig. 4c
4	090518-100118	No	2D	Descending	20	Fig. S2
5	091217-100121	No	2A	Ascending	230	Fig. S2
6	090531-091227	No	5D	Descending	-60	Fig. S2

from SEG4  $P$ -wave velocity, using the empirical relationships given by Castagna *et al.* (1985) and Quijada & Stewart (2007), respectively. We apply the same smoothing criteria to the slip inversion, and compare the results to the slip distribution from the homogeneous elastic half-space.

Since the aftershock hypocentres suggest the involvement of multiple faults (Fig. 3; see Section 3.1 for details), we also carry out slip inversions using multiple fault planes. The first fault plane is the optimal estimate from the uniform-slip search algorithm, and the geometry of the second fault plane is inferred from the east-dipping Karonga Fault or a west-dipping fault following an alignment/cluster of aftershocks. We manually discretize both faults with a cell size similar to the final discretization of the single fault inversion, and invert for distributed slip on triangular dislocations using the same algorithm from the single fault inversion. Once the slip distribution is calculated, we compare the fit with the results from the single fault inversion, and discuss if the addition of the second fault significantly alters or improves the model.

### 2.2.2 Time-series including Rungwe volcanic province

We used 150 ALOS-1 interferograms of various timespans to construct an InSAR time-series covering the 2007–2010 time period for the northern Lake Malawi/Rungwe Volcanic Province region using the method of Henderson & Pritchard (2013). We estimate the error on the time-series by quantifying the variance of the time-series in areas with no deformation. Based on analysis of pixels located in the north of our coverage area (far from the sources of deformation), scatter for the time-series is about  $\pm 2$  cm, which is similar to error bounds from other ALOS-1 time-series (e.g. Chaussard *et al.* 2013; Ebmeier *et al.* 2013).

## 2.3 Focal mechanisms from telesismic and regional seismograms

The focal mechanisms of the largest earthquakes of the December 2009 sequence can be characterized using moment–tensor analysis. The global centroid-moment tensor (GCMT) catalogue (Ekström *et al.* 2012) lists nine events in the period December 6–19 (Table 2), with the smallest event having  $M_w = 4.9$  and the largest  $M_w = 6.0$ . We complement the GCMT catalogue by applying the standard GCMT analysis to several smaller events that were well recorded by stations of the Global Seismographic Network and AfricaArray (Nyblade 2007) at regional (up to  $\sim 30^\circ$ ) epicentral distance (Table 2). We successfully modelled seven additional events that span the time window of 22 November through 19 December (Fig. 2a). These events include two foreshocks (events 1 and 2 in Table 2) to the first damaging event on December 6 (event 3). Due to the small size of these events, we were not able to simultaneously constrain

their locations and moment–tensor solutions. We first utilized a relative relocation algorithm (Howe 2019) similar to that of Cleveland & Ammon (2013) and confirmed that the events fall in a tightly clustered region (Fig. S4). Based on this result, the hypocentres for the small events were fixed at a location of  $10^\circ\text{S}$ ,  $33.9^\circ\text{S}$  and 12-km depth, and the fault-plane parameters were determined.

## 3 RESULTS

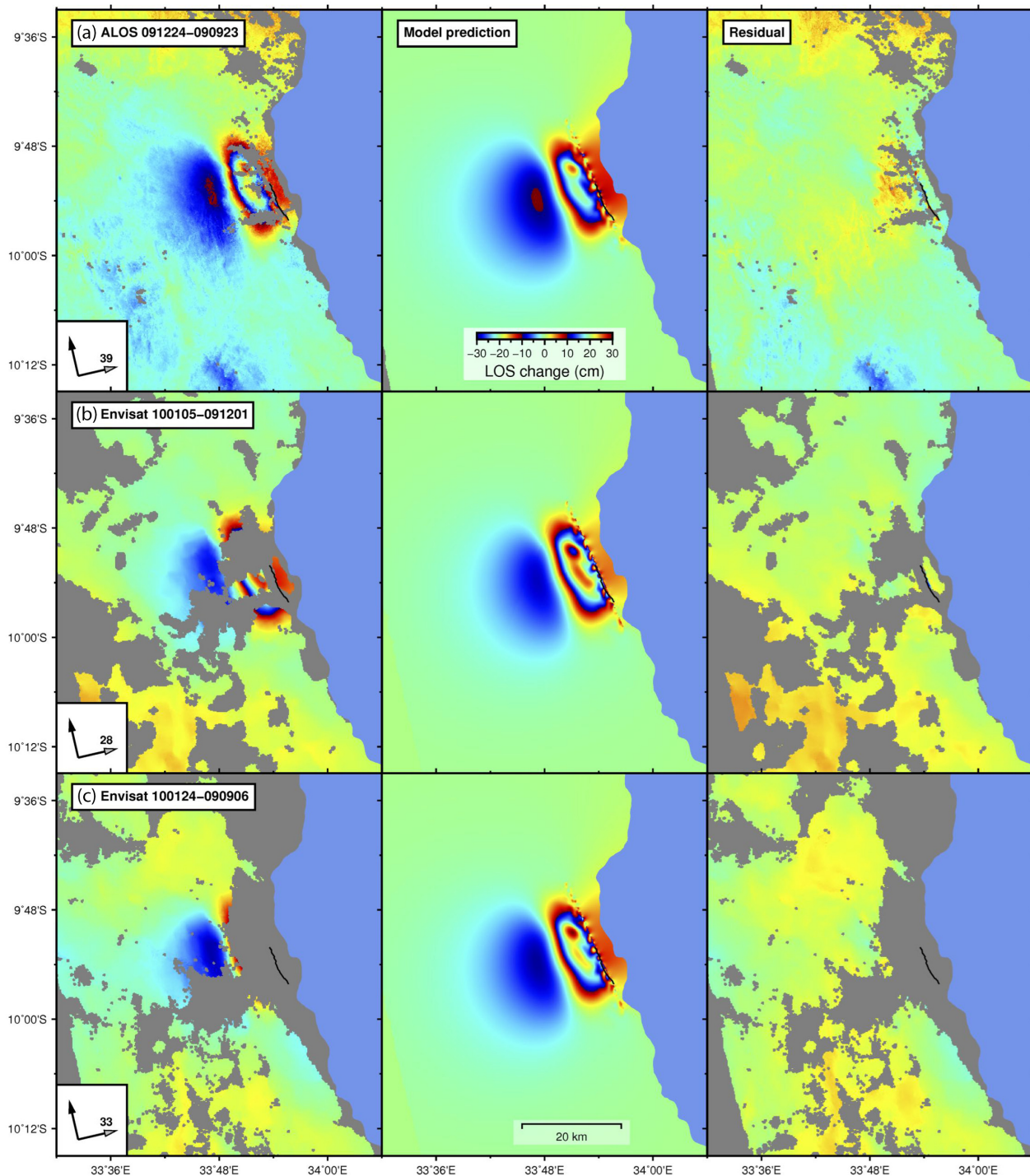
### 3.1 Aftershock sequence

The complete set of  $\sim 1900$  aftershock hypocentres forms a concentrated rectangular cloud trending NW–SE (Fig. 2b), coinciding with the bulk of the CMT locations, as well as the primary fault identified by Biggs *et al.* (2010). The spatial distribution of earthquakes terminates abruptly to the west, with very few events located west of the local array. A larger number of events are located beneath Lake Malawi to the east of the array, including a fairly concentrated cluster just southeast of Karonga town that forms part of a quasi-linear band paralleling the dominant cloud onshore. Hypocentral depths range from the surface to over 15 km depth, with the deepest events clustered in the far northwest region, as well as in the southeast corner directly beneath Karonga.

The more precise relative locations of the  $\sim 1050$  hypoDD events illuminate more features within the cloud of the seismicity than seen in the complete catalogue, providing greater clarity on likely fault activity (Fig. 2c). Overall, the spatial distribution of the events is similar to the absolute locations, with earthquakes clustered in the region of faulting inferred by Biggs *et al.* (2010), with a sharp boundary to the west, and several events distributed to the east. The faulting is further illuminated by plotting hypocentres in cross-section perpendicular to the strike of the St Mary’s Fault (Figs 3b–e and Fig. S1).

In the north, the events cluster in a confined, west-dipping body at 6–13 km depth that projects back to the surface location of the St Mary’s Fault (Fig. 3b). Stepping south into the middle of the cluster (Fig. 3c), the events shallow to  $< 10$  km depth, and appear to bifurcate into two trends: one dipping to the west from 2 to 10 km depth with a surface projection at the location of the St Mary’s Fault; and one dipping to the east from  $\sim 5$  to 10 km depth that may represent a deep extension of the Karonga Fault. This interpretation implies that the Karonga and St Mary’s Faults intersect near 5 to 8 km depth. Continuing south (Fig. 3d), the west-dipping St Mary’s Fault persists, while the east-dipping feature is no longer apparent in the aftershock patterns. Finally, in the southern portion of the sequence near Karonga village (Fig. 3e), the seismicity breaks into two to three subparallel west-dipping features indicative of synthetic faults with approximately 5-km spacing. One of the lineations at 2 to 10 km depth is consistent with St Mary’s Fault, suggesting that this fault is active over the entire length of the aftershock sequence.





**Figure 5.** The data (left-hand column), model prediction (centre column), and residual after data minus model (right-hand column) for three interferograms that were unwrapped and then rewrapped using the same colour scale. (a) ALOS-1 between 23 September and 24 December, same as Fig. 4(a). (b) Envisat between 1 December and 5 January, same as Fig. 4(b). (c) Envisat between 6 September and 24 January, same as Fig. 4(c). The modelled interferograms are generated using the best-fitting single-fault slip distribution. The black line indicates the locations of fault rupture from Macheyeke *et al.* (2015).

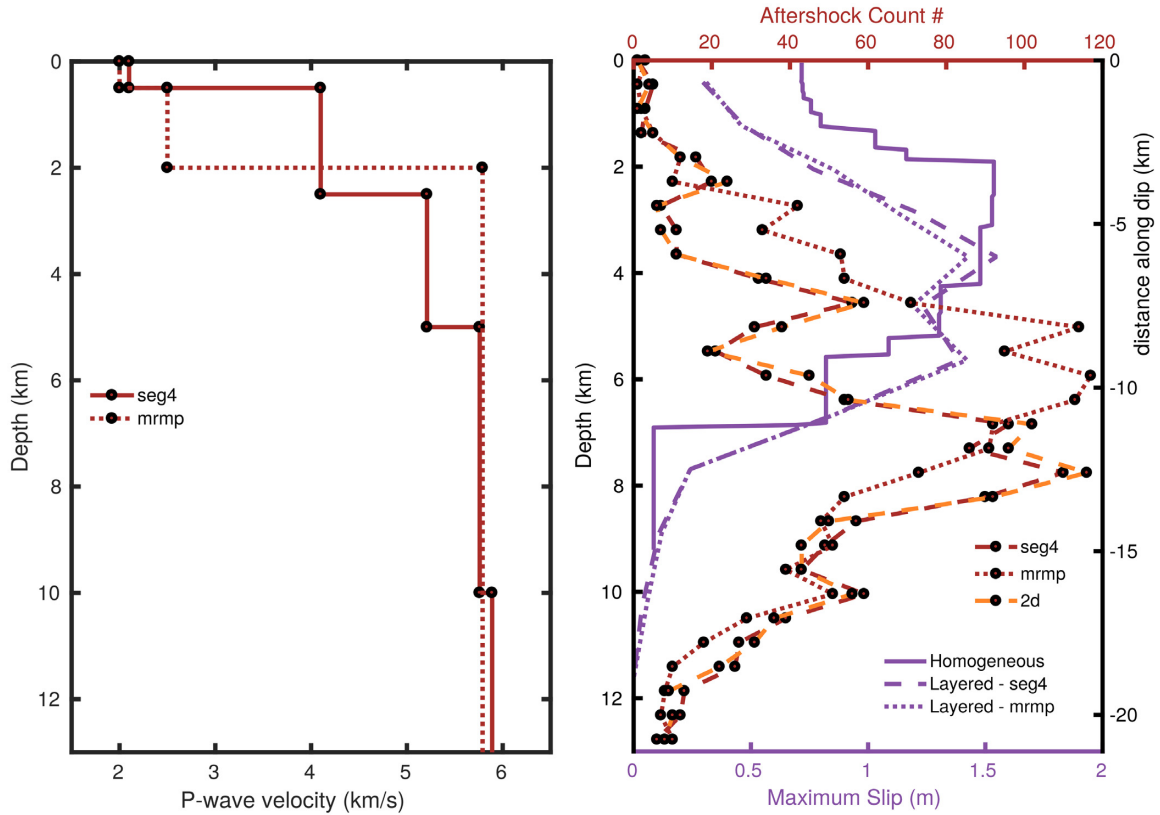
The additional features are active at larger depth (6 to 12 km), and they suggest west-dipping faults that would project to the surface beneath Lake Malawi.

There is a significant proportion of the aftershocks lying at or below the downdip limit of the primary slip surface inferred from InSAR analysis (<6 km; Fig. 6). While many of the aftershock lineations described above extend upward to depths <5 km, the

mean depths of the catalogues are 7.1 and 7.5 km for the complete and hypoDD catalogues, respectively. The densest clusters of events in the hypoDD catalogue locate to the 5-10 km depth range. In addition, the approximate dips of the west-dipping lineations inferred from the cross-sections (Fig. 3) is of order  $\sim 60^\circ$ , steeper than the west-dipping focal planes inferred for the CMT events (generally  $35\text{--}50^\circ$ ). Absolute depths are quite difficult to constrain

**Table 2.** CMT analysis of primary event sequence. Events in global CMT catalogue (Ekström *et al.* 2012) are marked with a “\*”. All other events have fixed location. All events are from the year 2009.

No	Mon.	Day	Hour	Min.	Sec.	Lat.	Lon.	Depth (km)	Scalar Moment (N-m)	Mag (NEIC)	M <sub>w</sub>	Strike	Dip	Rake
1	11	22	20	39	09.20	-10.00	33.90	12	1.01E+16	4.6	4.6	145	43	-105
2	11	25	21	14	33.50	-10.00	33.90	12	8.39E+15	4.5	4.5	113	50	-147
3*	12	06	17	36	39.83	-9.97	33.90	12	4.90E+17	5.8	5.7	162	36	-81
4	12	06	17	58	15.40	-10.00	33.90	12	4.18E+16	5.3	5.0	157	44	-74
5*	12	06	18	00	05.83	-10.03	33.90	12	8.18E+16	5.3	5.2	165	43	-75
6*	12	06	18	29	17.85	-10.15	33.97	14	6.48E+16	5.2	5.1	141	45	-103
7	12	06	19	36	43.50	-10.00	33.90	12	2.36E+16	5.1	4.8	162	36	-87
8	12	07	03	35	43.80	-10.00	33.90	12	1.51E+16	5.0	4.7	126	34	-109
9*	12	07	09	31	47.98	-9.66	33.90	20	3.69E+16	5.0	5.0	202	35	-62
10	12	07	18	09	40.30	-10.00	33.90	12	2.43E+16	4.6	4.9	210	52	-16
11*	12	07	18	16	34.99	-9.89	33.88	18	3.49E+16	4.9	5.0	202	49	-54
12*	12	08	03	09	01.44	-9.89	33.89	12	1.01E+18	5.9	5.9	172	36	-77
13*	12	11	04	49	10.88	-10.08	33.94	17	2.73E+16	5.0	4.9	149	49	-84
14	12	11	20	06	28.40	-10.00	33.90	12	1.85E+16	4.6	4.8	146	48	-89
15*	12	12	02	27	06.69	-9.79	33.85	12	2.17E+17	5.5	5.5	182	37	-75
16*	12	19	23	19	20.43	-10.02	33.93	12	1.09E+18	6.0	6.0	155	38	-88

**Figure 6.** Left-hand panel: two  $P$ -wave velocity models tested for aftershock locations (MRMP and SEG4). Right-hand panel: histogram of fault slip (purple lines) from our inversions in homogeneous half-space and layered models compared to the distribution of hypoDD relocated aftershock depths using the two 1-D seismic velocity models (red lines), and a 2-D model approximation using station-specific sediment thickness (orange line). All models suggest that most geodetic coseismic slip is shallower than the majority of aftershocks.

using a limited distribution of stations directly above the events, as is the case here. Tests with alternative velocity models suggest that the spatial patterns interpreted above are robust, but the absolute depths can differ by an average of up to 2 km. Even with this uncertainty, most of the aftershocks still occur at greater depths than the primary slip zone inferred from InSAR.

The temporal and magnitude statistics of the complete hypoinverse catalogue are broadly consistent with a main shock–aftershock

sequences on shallow crustal faults. Above a magnitude of completeness of 0.8, the sequence has a Gutenberg–Richter  $b$ -value of  $\sim 0.8$ , and a modified Omori’s law decay exponent of 0.73. These parameters represent the aftershock behaviour in the time window of the local deployment, which begins approximately 18 d after the main shock, which we assume to be the largest event (19 December) of the primary sequence. As such, it does not describe the full complexity of the extended 27-d primary sequence, nor does it

include the likely energetic 18-d period following the  $M_w$  6 event (for which we have no events). To the extent that these statistics are representative, they imply a relatively energetic, slowly decaying aftershock sequence, with an anomalously low  $b$ -value (deficit of smaller events) relative to most normal-faulting sequences (Schorlemmer *et al.* 2005). Low  $b$ -values are suggestive of high differential stress (Scholz 2015). Such a stress state is atypical of extensional systems, although similarly low  $b$ -values have been observed in other regions of the western rift (Lavayssiere *et al.* 2019). One interpretation is that the relatively high stress state suggests that the faults are relatively immature and occur in strong, intact crust, consistent with inferences from previous studies (Biggs *et al.* 2010; Fagereng 2013).

### 3.2 Focal mechanisms

In map view, the CMT events fall along a NW–SE lineation (Fig. 2) that correlates well with the strike of the primary fault inferred by Biggs *et al.* (2010). Uncertainties in global velocity models result in absolute CMT locations determined to within only 10–20 km, and the north–south distribution of epicentres and precise correlation of the largest events with the region of highest slip is difficult to ascertain. The relative relocation analysis (Fig. S4) suggests that all of the events are tightly clustered within a  $\sim 20$ -km-long linear trend, roughly consistent with the length of the primary faulting inferred from InSAR. The final, largest event in the sequence lies at the southern end of the cluster, which also agrees with the InSAR modelling (Biggs *et al.* 2010).

The combined observed seismic moment for the 16 events (Table 2) is equal to  $3.19 \times 10^{18}$  N m, corresponding to a moment magnitude of  $M_w = 6.3$ . Most focal mechanisms suggest normal faulting on NW–SE striking fault(s), while two (including one foreshock) show evidence of a significant strike-slip component (events 2, 10). All events have moderate fault dips of 34–56°, depending on choice of fault plane. For most events, the west-dipping plane has a shallower dip. The two mechanisms with a strike-slip component fall at the northern end of the event cluster (Fig. S4). For one of these (event 10) there are sufficient data to carefully evaluate the robustness of the mechanism, and it is clear that the strike-slip component of this event is required.

For all events, depths appear to be shallow: most have a fixed GCMT value of 12 km, although four events give best solutions with depths of 14–20 km. For the largest events in the 2009 sequence, Biggs *et al.* (2010) determined shallow (<8 km) depths using teleseismic  $P$  waveform modelling.

### 3.3 InSAR slip distribution of the 2009 earthquake sequence

The best-fitting fault geometry from the uniform-slip inversion agrees with the locations of St Mary’s Fault and the surface rupture from the 2009 earthquakes, and dips at 38–45° to the west (Fig. S5); this result agrees with previous analysis of Biggs *et al.* (2010). The result indicates a fault plane that is 16.8 km long and 5.8 km wide, with an average slip amount of  $\sim 0.83$  m. The geodetic moment corresponds to a moment magnitude of  $M_w$  6.2 similar to that found in previous work ( $3.22$ – $3.5 \times 10^{18}$  N m, Biggs *et al.* 2010; Hamiel *et al.* 2012) and close to the total seismic moment of  $3.2 \times 10^{18}$  N m ( $M_w$  6.3) for the sequence. Since the result is not sensitive to the variation of rake between  $-70^\circ$  and  $-110^\circ$ , we select a rake of  $-90^\circ$  assuming a pure normal faulting event. We utilized a dip

of  $38^\circ$ , which is suggested by GCMT and Biggs *et al.* (2010), as the final dip value being used in the distributed slip inversion. The complete parameters used in the distributed slip inversion are listed in Table 3.

The result from the distributed slip inversion shows that the slip zone projects to a location at depth between the surface traces of St Mary’s Fault and Karonga Fault, where most aftershocks are located (Fig. 3a). The maximum slip is between 1.4 and 1.6 m depending on the smoothing level (Fig. S6), but always occurred at 2–5 km depth (Figs 3b–e and 6). The RMS misfit is 0.014 m ( $R^2 = 0.95$ ), similar to the expected data noise suggesting a very good fit to the geodetic observations at the determined smoothing level. The model resolution has been assessed in Hamiel *et al.* (2012) using a checkerboard test and we use the same subfault size–depth relationship for each fault patch. The inversion result from the layered velocity model shows a similar slip distribution as a function of depth as the homogeneous half-space (Fig. 6), suggesting that the geodetic observations are robustly in favour of much shallower fault movement (<6 km) than the nominal GCMT centroid depths (12–20 km). This agrees well with the depth of the largest hypocentres (<8 km) determined using teleseismic  $P$  waveforms (Biggs *et al.* 2010), and it is slightly shallower than the mean depth of the aftershock distribution ( $\sim 7$  km).

We select two cases from the slip inversions using multiple fault planes to illustrate the impact of multiple faults. The first fault in both case is the St Mary’s Fault plane. The second fault of the first case shares the same geometry (strike, dip, location and dimension) as the optimal fault plane, but is 4.5 km deeper. In the second case, the second fault resembles the geometry of the east-dipping Karonga Fault with a dip of  $60^\circ$ . All the cells of the second faults also follow the same depth–size relationship which applies to the St Mary’s Fault plane. The results (Figs 8b, c, e and f) show that the models are similar in terms of both slip distribution and fit to the data (not shown). We do not think it is possible to argue that one model fits all of the data better than another given the poor interferogram quality—there are problems with coherence in the Envisat data and with the small number of interferograms overall, it is difficult to reduce the atmospheric noise. In both cases, slip along the second fault is confined within a small region, with a maximum slip of only  $\sim 0.4$  m. The primary optimal fault plane still accounts for over 90 per cent of the seismic moment, thus it is not likely that deeper west-dipping fault(s) or the Karonga Fault accommodated significant slip in the 2009 November–December earthquakes.

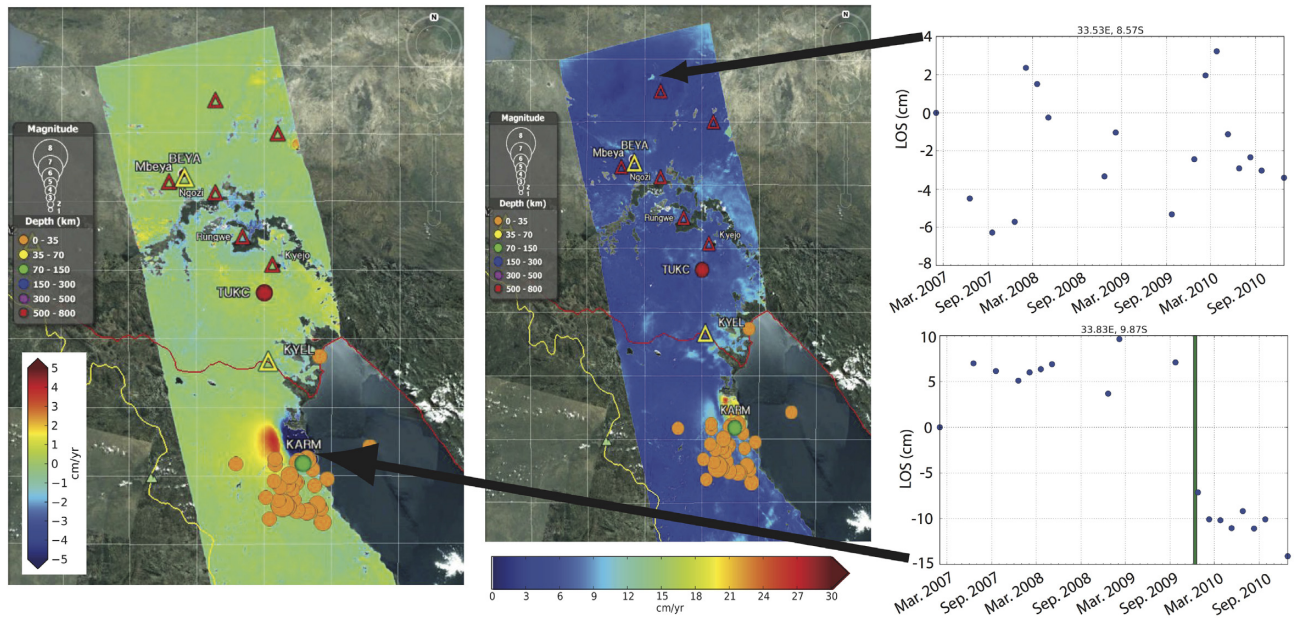
### 3.4 Regional InSAR time-series

The ALOS-1 time-series between 2007 and 2010 (Fig. 7) shows no significant deformation ( $\pm 2$  cm) at volcanoes in the Rungwe volcanic province. No deformation associated with the Karonga earthquake sequence was observed within the volcanic province. Individual interferograms suggest the presence of a seasonal signal (amplitude of 9 cm) in the northern part of the region (the signal is also apparent in the L2 residual plot for the time-series). The pattern shows maximum uplift in February–June and subsidence during the remaining months of the year. This uplift lags the typical rainy season (December–March) by approximately two months, and the correlation between the deformation and the wet/dry seasons suggests that the signal may be groundwater response. Surface deformation from the 2009 to 2010 Karonga earthquake sequence is clearly visible in the InSAR timeseries. The deformation pattern is located just north of the main cluster of located earthquakes and shows



**Table 3.** The best-fitting fault geometry used in the inversion of slip distribution.

Parameter	Value	Comments
Strike	156°	Constrained by the uniform slip inversion and surface rupture
Dip	38°	From gCMT and Biggs <i>et al.</i> (2010)
Rake	-90°	Assuming a pure normal fault (e.g. Biggs <i>et al.</i> 2010)
Depth of the Fault Top	0 km	Constrained by surface rupture
Width	15 km	Uniform slip inversion suggests at least 12 km
Length	29 km	Uniform slip inversion suggests at least 17 km
Fault northern end	(33.84190°E, 9.75042°S)	From uniform slip inversion
Fault southern end	(33.95012°E, 9.98973°S)	From uniform slip inversion



**Figure 7.** ALOS-1 time-series of InSAR phase change between 2007 and 2010. (a) Best-fitting linear ALOS time-series. Note that the numerical estimate for the linear rate very close to the earthquake location is not reliable. (b) L2 residual plot from the time-series show locations not fit well by a linear rate. For a and b: Red triangles are volcanoes, yellow are campaign GPS, labelled red, green and orange dots are cGPS, unlabelled orange circles are earthquakes during 2009–2010. No GPS stations were close enough to measure the ground deformation from the 2009 to 2010 sequence. (c) Seasonal signal in a rural part of Tanzania. Subsidence in June–Dec., uplift in February–May. (d) Deformation during the 2009–2010 earthquake sequence. Post-seismic deformation is clearer in individual interferograms (e.g. Hamiel *et al.* 2012).

+15 cm and –10 cm motion in radar line-of-sight coordinates. Although not overly noticeable in the time-series (Fig. 7), postseismic deformation is evident in individual interferograms (Hamiel *et al.* 2012).

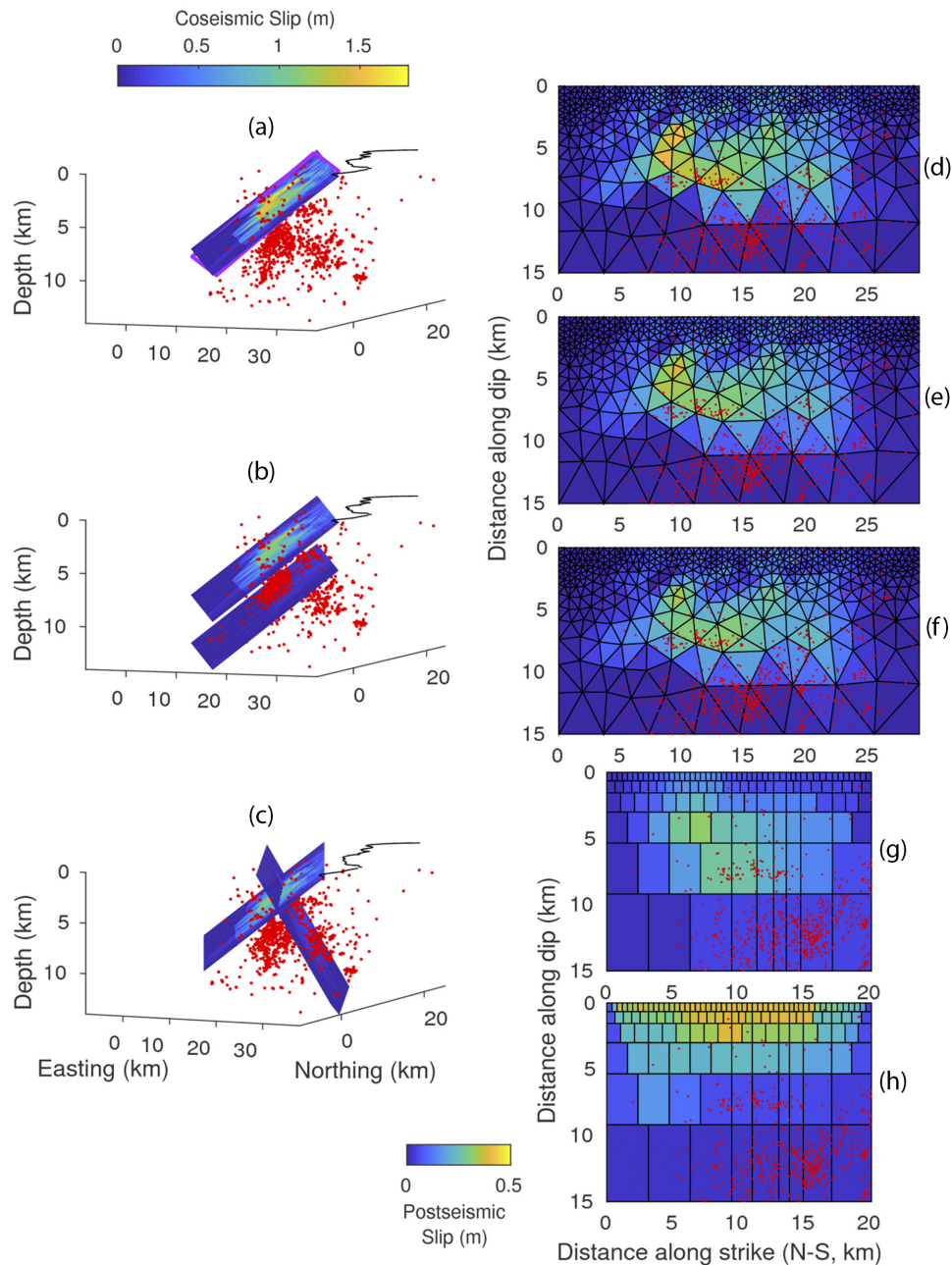
## 4 DISCUSSION

### 4.1 Were multiple faults involved in the earthquake sequence?

Based on the relocated aftershock distribution, it is clear that more than a single fault was involved in producing aftershocks between January and May 2010. The January and May 2010 sequence shows an east–west distribution spanning 30 km (Fig. 3). It appears to involve multiple west-dipping faults including the St Mary’s Fault, as well as a possible east-dipping deep extension of the Karonga Fault (Figs 3c–e and Fig. S1). The evidence for multiple faults is particularly compelling just west of the lakeshore between 9.90°S and 9.95°S, where there are several deeper clusters (>5 km depth) east of the trace of the St Mary’s Fault (Figs 2c and 3e). These

events are well recorded by the local array, and are consistent with plausible west-dipping faults that breach the surface beneath Lake Malawi (Mortimer *et al.* 2007), and are clearly inconsistent with the fault plane associated with the St Mary’s Fault. Coulomb stress calculations suggest that faults of this orientation and location (deeper and to the east of the primary slip zone) would be pushed closer to failure by slip on a St Mary’s type feature (Fagereng 2013).

The activation of an east-dipping fault beneath the centre of the sequence is less clear. This region is the most energetic portion of the sequence (in terms of number of events), and the events remain somewhat cloud-like, even following relocation with hypoDD. A planar east-dipping feature can be discerned in some of the cross sections (e.g. Fig. 3c), and the coincidence of this feature with a projection of the Karonga Fault provides a reasonable interpretation. However, a simple analysis of Coulomb stress change on the Karonga Fault due to the coseismic slip on the St Mary’s Fault (similar to Fagereng 2013) suggests a decrease of stress along Karonga Fault. Nevertheless, it is possible that the stress heterogeneity introduced by the fault intersection contributes to the large number of events, as well as the cloud-like seismicity pattern (e.g. Sumy *et al.*



**Figure 8.** The inverted slip distributions during the 2009 earthquakes, and the comparison between different models in terms of their slip patterns. (a) Slip distribution using the best-fitting fault geometry (St Mary's Fault) only. Purple line sketches the fault extent used by Hamiel *et al.* (2012). The shoreline of Lake Malawi is a black line, and aftershocks during January–May 2010 (using the velocity model seg4) are shown as red dots. (b) Slip distribution using the best-fitting fault geometry and another deep west-dipping fault, which shares the same strike and dip with the best-fitting fault. (c) Slip distribution using the best-fitting fault geometry and an east-dipping fault approximating the Karonga Fault. (d) Slip distributions using only the best-fitting fault geometry. (e) Slip on the St Mary's Fault, using the two-fault model from (b). (f) Slip on the St Mary's fault, using the two-fault model from (c). (g) Coseismic slip on the St Mary's fault from Hamiel *et al.* (2012). (h) Postseismic slip on the St Mary's fault from Hamiel *et al.* (2012). Note that (h) is plotted using a different colour scale. (d)–(h) are viewed at an angle normal to the fault plane.

2013). Aftershock activation on intersecting antithetic faults have been observed in other extensional settings (e.g. Karasözen *et al.* 2018; Lavecchia *et al.* 2018; Walters *et al.* 2018).

Whether multiple faults were involved during the November–December 2009 main shocks is harder to address. The InSAR data are permissive of multiple faults being active during the November–December 2009 sequence (Figs 8a–f), but we agree with Biggs *et al.* (2010) that the majority of slip occurred on the west-dipping

St Mary's Fault. Little slip is preferred on the alternate fault geometries suggested by the aftershocks, and the slip inferred from the InSAR data matches the total seismic moment of the November–December sequence quite well as found in previous work (Biggs *et al.* 2010; Hamiel *et al.* 2012). These results suggest that primary slip on the immature St Mary's Fault excited subsequent slip on synthetic west-dipping hanging-wall faults beneath Lake Malawi and possibly on the east-dipping Karonga Fault.

The teleseismic and regional locations of the November–December 2009 events are also generally consistent with the sequence primarily being located on the St Mary’s Fault. While the absolute locations of the largest events in the sequence (events 3, 12, 15 and 16) that dominate the moment release are located systematically south of the geodetically inferred slip patch, the offset ( $\sim 5\text{--}10\text{ km}$ ) is within that expected for location error for earthquake catalogues in Africa (e.g. Weinstein *et al.* 2017) and globally (e.g. Weston *et al.* 2014). The relative locations of these events follow a NW–SE trend parallel to the St Mary’s Fault, with a north–south extent that is quite similar to the length of the St Mary’s Fault in the geodetic models. Furthermore, the largest event in the sequence is to the south, consistent with InSAR modelling of this event (Biggs *et al.* 2010). The events with a significant strike-slip component suggest activity on additional fault(s) in the northern portion of the sequence, perhaps associated with a change of basement crustal fabric inferred from potential field data (Kolawole *et al.* 2018).

In summary, although the majority of slip and seismic moment release appear to be associated with the St Mary’s Fault, our results suggest that part of the main December earthquake series and the resulting aftershock series occurred on multiple interacting faults. These fault interactions may have contributed to the complexity of the Karonga earthquake series.

#### 4.2 Depth of coseismic slip and aftershocks

We find that there is an offset in depth between the projected location of coseismic slip on the St Mary’s Fault and the distribution of aftershocks (Figs 3b–e and 8). This may be due in part to uncertainty in fault-plane geometry and/or event depths, but tests of both variations in fault dip and velocity models suggest that these uncertainties cannot explain the difference (Fig. 6 and Fig. S3). The InSAR observations prefer shallow slip ( $< 6\text{ km}$  depth) on a fault dipping between  $38^\circ$  and  $45^\circ$ , shallower than the dip suggested by the aftershocks ( $\sim 60^\circ$ ). As discussed above, the absolute depth and apparent dip of the aftershock distribution could be shallowed by alternative velocity models, but it is clear that even in this case the bulk of the seismicity is deeper than 5 km, deeper than the peak slip (Fig. 6). Such a difference between the location of coseismic slip and subsequent aftershocks has been seen in a number of tectonic settings (e.g. Nissen *et al.* 2010; Roustaei *et al.* 2010; Elliott *et al.* 2011; Wei *et al.* 2015; Karasözen *et al.* 2016, 2018; Lavecchia *et al.* 2018). It is often interpreted as resulting from fault complexity that results in the arrest of coseismic slip, either due to changes in fault geometry (e.g. Elliott *et al.* 2011; Karasözen *et al.* 2018; Lavecchia *et al.* 2018) or changes in fault strength and/or rate-state behaviour (e.g. Semmane *et al.* 2005; Roustaei *et al.* 2010; Wei *et al.* 2015).

The total geodetic moment is similar to the total seismic moment as well as previous geodetic estimates (Biggs *et al.* 2010; Hamiel *et al.* 2012), so there is no geodetic evidence for the involvement of aseismic slip or fluids in driving the earthquake sequence, including fluids associated with the Rungwe volcanic province. Postseismic deformation is most likely caused by fault afterslip and not poroelastic relaxation or fault healing (Hamiel *et al.* 2012). Afterslip is shallow (upper 2 km) and of relatively low magnitude, with a maximum slip of 40 cm compared to 120 cm of coseismic slip and a total moment between December 2009 and August 2010 that is about 40 per cent of the coseismic moment (Figs 8g and h, Hamiel *et al.* 2012). The afterslip does not appear to have been important in producing aftershocks in the 2010 January–May period regardless of whether they are associated with St Mary’s fault or other nearby

west-dipping faults, since few of our detected events occur in the upper 2 km (Figs 3 and 8). Further, we do not detect any continued deformation near the 2009–2010 afterslip location between 2014 and 2017 with Sentinel-1 InSAR analysis (Henderson *et al.* 2017).

#### 4.3 Regional and global comparison

The suggestion of intersecting faults in the 2010 aftershock data provides further context for understanding the sequence of main events in December 2009. With several  $M_w$  5+ events that do not follow typical main shock–aftershock behaviour, the 2009 events suggest some basic characteristics of an earthquake swarm (Holtkamp & Brudzinski 2011), although there is insufficient local station coverage during the December 2009 period to adequately assess whether the events are truly swarm-like. Swarms are often associated with aseismic deformation and/or fluid migration (e.g. Holtkamp & Brudzinski 2011), neither of which is inferred here. In addition, the catalogue presented here suggests that by January 2010, the process can be characterized as an energetic aftershock sequence with magnitude and decay characteristics within the range of normal earthquakes. Rather than a swarm, Biggs *et al.* (2010) characterize the 2009 events as a sequence of moderate events, which they interpret as resulting from structural complexity on an immature St Mary’s Fault. We concur with the characterization as a sequence, but suggest that fault intersection and interaction also contribute to the complexity of the sequence.

While most of the cumulative slip during the main shock sequence appears to have occurred on the St Mary’s Fault, other nearby structures could have modulated failure on the St Mary’s Fault and explain the complexity of the sequence. An intersecting structure can form a geometrical barrier to fault propagation (e.g. King 1986), and the presence of such barriers could contribute to the occurrence of a complex series of events rather than one larger event (e.g. Elliott *et al.* 2011; Karasözen *et al.* 2016; Lavecchia *et al.* 2018; Walters *et al.* 2018). The intersection of the Karonga Fault with the St Mary’s Fault helps explain the depth distribution of the aftershock sequence described above, and it may also explain the along-strike complexity of the main December sequence. Variation in the interaction between these faults along strike could produce barriers that limit slip in any individual event. Additionally, magnetic data and geologic mapping constrain the tectonic grain in the Proterozoic basement in the region of the Karonga sequence (Kolawole *et al.* 2018). A change in the orientation of this fabric is observed near Karonga town; it is oriented WNW to the north and NW to the south. The intersection of these ancient structures with the St Mary’s Fault may also contribute to the complex earthquake series, as suggested by Kolawole *et al.* (2018).

## 5 CONCLUSIONS

We combine teleseismic and regional seismic recordings, InSAR imagery, and recordings of aftershocks from a temporary (4-month) local network of six seismometers to examine the unusual 2009 Karonga earthquake series in the weakly extended Malawi Rift. Consistent with previous studies (e.g. Biggs *et al.* 2010; Hamiel *et al.* 2012), new analysis of InSAR data and global/regional CMT solutions from this study suggest that a west-dipping fault (the St Mary’s Fault) in the hanging wall is responsible for the majority of slip during the main shocks. However, aftershocks also delineate several other closely spaced west-dipping faults to the east of the St Mary’s Fault in the southern part of the study area and one



possible east-dipping fault to the west in the central part of the study area; the latter appears to project to the surface trace of the Karonga Fault. These results suggest that the interactions of closed spaced and sometimes intersecting faults may have contributed to the complexity of the December 2009 Karonga earthquake series. InSAR data permit the involvement of multiple faults, but do not require it—we have modelled up to two, but the seismicity suggests that more may have been active. However, the clear majority of coseismic slip occurred on the St Mary's Fault in agreement with previous work (Biggs *et al.* 2010; Hamiel *et al.* 2012). We suggest that structural complexities on the St Mary's Fault resulting from the intersection of the St Mary's with other faults at depth may have limited coseismic slip during the main shocks to shallow depths. Structural heterogeneity on the St Mary's Fault caused by its youthfulness (Biggs *et al.* 2010) and the interaction of multiple faults at depth may also explain the occurrence of a series of similar sized earthquakes instead of a normal main shock–aftershock series during the first month of activity.

The majority of aftershocks occur at greater depths (>6 km) than the slip during main shocks based on InSAR data (<6 km); this depth discrepancy appears to be robust. We suggest that shallow rupture during the December 2009 main shock series increased stress on deeper faults and nearby favourably oriented faults (e.g. Fagereng 2013), causing the energetic aftershock series. The aftershock series is associated with unusually low *b*-values for extensional faults (Schorlemmer *et al.* 2005), which indicates high differential stress (Scholz 2015), consistent with rupture on immature faults in strong crust (e.g. Biggs *et al.* 2010; Fagereng 2013).

We calculated InSAR interferograms over a broader area than previous studies, and do not observe evidence for ground deformation from fluids/magma or aseismic slip near Karonga or the Rungwe Volcanic Province.

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## SUPPORTING INFORMATION

Supplementary data are available at *GJI* online.

**karonga2010\_hypo2k.loc.tex.** ASCII text file containing complete Hypoinverse aftershock catalogue.

**karonga2010\_hypoDDv2.loc.tex.** ASCII text file containing hypoDD aftershock catalogue.

**Figure S1.** The complete 14 E–W cross sections from Fig. 3 a covering the faulting region, each labelled A to N. See Fig. 3 caption for further details of the symbols and colours.

**Figure S2.** Envisat interferograms that were too noisy to use in the inversion and the predicted interferograms generated from the preferred slip distribution using the three interferograms in Fig. 5 as described in the main text. The modelled surface deformation is wrapped using a radar wavelength of 5.6 cm (the wavelength of Envisat ASAR sensor).

**Figure S3.** Top panel: modelled slip distribution using the layered velocity model seg4 and surface deformation from the three interferograms used in the half-space inversion (Fig. 5). The fault is discretized into 248 rectangular cells with a varying size, resembling the size of cells used in the half-space inversion (Fig. 8). Bottom panel: shows the ALOS-1 data (left, Figs 4a and 5a), modelled interferogram (centre) and residual for the resampled data.

**Figure S4.** Surface-wave relative location analysis (Howe 2019) applied to CMT events from Table 2 in manuscript. (A) Initial PDE locations (white squares) and final relative event locations (red circles). Location of the cluster centroid is not constrained by the analysis. (B) Relative-location error ellipses for each event. Errors are generally <2 km and NW–SE trend is robust. (C) Final CMT mechanisms plotted at accurate relative locations. Entire cluster has been shifted ~15 km north and ~3 km east to correspond to the dominant faulting inferred from InSAR analysis. Event(s) with strike-slip component discussed in the text are located at the northern end of the cluster.

**Figure S5.** The relationship between the model InSAR RMS misfit and the dip of the fault plane used in the model, when other fault parameters remain fixed. We select dip = 38° (doubled circle) as our preferred fault plane and describe the model results in the main manuscript.

**Figure S6.** L-curve showing InSAR coseismic slip model roughness versus model minus data difference) and corresponding slip distributions at different points along the L-curve. We select the results from the third row as our preferred, best-fitting model in the main text.

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## Key words

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Composition and structure of the continental crust	Space geodetic surveys	Interferometry
Composition and structure of the core	Tides and planetary waves	Inverse theory
Composition and structure of the mantle	Time variable gravity	Joint inversion
Composition and structure of the oceanic crust	Transient deformation	Neural networks, fuzzy logic
Composition of the planets		Non-linear differential equations
Creep and deformation		Numerical approximations and analysis
Defects	GEOGRAPHIC LOCATION	Numerical modelling
Elasticity and anelasticity	Africa	Numerical solutions
Electrical properties	Antarctica	Persistence, memory, correlations, clustering
Equations of state	Arctic region	Probabilistic forecasting
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Friction	Australia	Spatial analysis
High-pressure behaviour	Europe	Statistical methods
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Microstructure	Japan	Time-series analysis
Permeability and porosity	New Zealand	Tomography
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	South America	
GENERAL SUBJECTS	GEOMAGNETISM and ELECTROMAGNETISM	PLANETS
Core	Archaeomagnetism	Planetary interiors
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Geomechanics	Controlled source electromagnetics (CSEM)	
Geomorphology	Dynamo: theories and simulations	SEISMOLOGY
Glaciology	Electrical anisotropy	Acoustic properties
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Hydrogeophysics	Electromagnetic theory	Coda waves
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Acoustic-gravity waves	Palaeomagnetic secular variation	Rheology and friction of fault zones
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Gravity anomalies and Earth structure	Rock and mineral magnetism	Seismic interferometry
Loading of the Earth	Satellite magnetism	Seismicity and tectonics
Lunar and planetary geodesy and gravity		Seismic noise
Plate motions	GEOPHYSICAL METHODS	Seismic tomography
Radar interferometry	Downhole methods	Site effects
Reference systems	Fourier analysis	Statistical seismology
Satellite geodesy	Fractals and multifractals	Surface waves and free oscillations
Satellite gravity	Image processing	Theoretical seismology
Sea level change		

Tsunami warning  
 Volcano seismology  
 Wave propagation  
 Wave scattering and diffraction

#### TECTONOPHYSICS

Backarc basin processes  
 Continental margins: convergent  
 Continental margins: divergent  
 Continental margins: transform  
 Continental neotectonics  
 Continental tectonics: compressional  
 Continental tectonics: extensional  
 Continental tectonics: strike-slip and transform  
 Cratons  
 Crustal structure  
 Diapirism  
 Dynamics: convection currents, and mantle plumes  
 Dynamics: gravity and tectonics  
 Dynamics: seismotectonics  
 Dynamics and mechanics of faulting  
 Dynamics of lithosphere and mantle  
 Folds and folding  
 Fractures, faults, and high strain deformation zones  
 Heat generation and transport

Hotspots  
 Impact phenomena  
 Intra-plate processes  
 Kinematics of crustal and mantle deformation  
 Large igneous provinces  
 Lithospheric flexure  
 Mechanics, theory, and modelling  
 Microstructures  
 Mid-ocean ridge processes  
 Neotectonics  
 Obduction tectonics  
 Oceanic hotspots and intraplate volcanism  
 Oceanic plateaus and microcontinents  
 Oceanic transform and fracture zone processes  
 Paleoseismology  
 Planetary tectonics  
 Rheology: crust and lithosphere  
 Rheology: mantle  
 Rheology and friction of fault zones  
 Sedimentary basin processes  
 Subduction zone processes  
 Submarine landslides  
 Submarine tectonics and volcanism  
 Tectonics and climatic interactions  
 Tectonics and landscape evolution  
 Transform faults  
 Volcanic arc processes

#### VOLCANOLOGY

Atmospheric effects (volcano)  
 Calderas  
 Effusive volcanism  
 Eruption mechanisms and flow emplacement  
 Experimental volcanism  
 Explosive volcanism  
 Lava rheology and morphology  
 Magma chamber processes  
 Magma genesis and partial melting  
 Magma migration and fragmentation  
 Mud volcanism  
 Physics and chemistry of magma bodies  
 Physics of magma and magma bodies  
 Planetary volcanism  
 Pluton emplacement  
 Remote sensing of volcanoes  
 Subaqueous volcanism  
 Tephrochronology  
 Volcanic gases  
 Volcanic hazards and risks  
 Volcaniclastic deposits  
 Volcano/climate interactions  
 Volcano monitoring  
 Volcano seismology