#### Lecture 24

## Exemplary Inverse Problems including Vibrational Problems

#### Syllabus

Lecture 01 Describing Inverse Problems Probability and Measurement Error, Part 1 Lecture 02 Probability and Measurement Error, Part 2 Lecture 03 Lecture 04 The L<sub>2</sub> Norm and Simple Least Squares A Priori Information and Weighted Least Squared Lecture 05 **Resolution and Generalized Inverses** Lecture 06 Lecture 07 Backus-Gilbert Inverse and the Trade Off of Resolution and Variance Lecture 08 The Principle of Maximum Likelihood Lecture 09 **Inexact Theories** Lecture 10 Nonuniqueness and Localized Averages Vector Spaces and Singular Value Decomposition Lecture 11 Lecture 12 Equality and Inequality Constraints Lecture 13  $L_1$ ,  $L_\infty$  Norm Problems and Linear Programming Lecture 14 Nonlinear Problems: Grid and Monte Carlo Searches Nonlinear Problems: Newton's Method Lecture 15 Lecture 16 Nonlinear Problems: Simulated Annealing and Bootstrap Confidence Intervals Lecture 17 **Factor Analysis** Varimax Factors, Empircal Orthogonal Functions Lecture 18 Lecture 19 Backus-Gilbert Theory for Continuous Problems; Radon's Problem Lecture 20 Linear Operators and Their Adjoints Lecture 21 Fréchet Derivatives Lecture 22 Exemplary Inverse Problems, incl. Filter Design Lecture 23 Exemplary Inverse Problems, incl. Earthquake Location Lecture 24 **Exemplary Inverse Problems, incl. Vibrational Problems** 

### Purpose of the Lecture

solve a few exemplary inverse problems

tomography vibrational problems determining mean directions

## Part 1

## tomography



#### discretization: model function divided up into M pixels $m_i$

#### data kernel

## $G_{ij}$ = length of ray *i* in pixel *j*

#### data kernel

## $G_{ij}$ = length of ray *i* in pixel *j*

here's an easy, approximate way to calculate it

### start with G set to zero



then consider each ray in sequence

#### divide each ray into segments of arc length $\Delta s$



and step from segment to segment

## determine the pixel index, say *j*, that the *center* of each line segment falls within



You can make this approximation indefinitely accurate simply by decreasing the size of  $\Delta s$ 

(albeit at the expense of increase the computation time)

Suppose that there are  $M = L^2$  voxels

A ray passes through about L voxels

**G** has *NL*<sup>2</sup> elements *NL* of which are non-zero

so the fraction of non-zero elements is 1/L

hence **G** is very sparse

#### In a typical tomographic experiment

#### some pixels will be missed entirely

#### and some groups of pixels will be sampled by only one ray

#### In a typical tomographic experiment

#### some pixels will be missed entirely the value of these pixels is completely undetermined

#### and some groups of pixels will be sampled by only one ray only the average value of these pixels is determined

hence the problem is mixed-determined (and usually *M*>*N* as well)

#### you must introduce some sort of a priori information to achieve a solution say

## a priori information that the solution is small

or

a priori information that the solution is smooth

## Solution Possibilities

 Damped Least Squares (implements smallness): Matrix G is sparse and very large use bicg() with damped least squares function

2. Weighted Least Squares (implements smoothness): Matrix F consists of G plus second derivative smoothing use bicg() with weighted least squares function

## Solution Possibilities

 Damped Least Squares: Matrix G is sparse and very large use bicg() with damped least squares function

test case has very

good ray coverage,

2. Weighted Least Squares: Matrix **F** consists of **G** plus second derivative smoothing use **bicg()** with weighted least squares function





#### Data, plotted in Radon-style coordinates



Lesson from Radon's Problem: Full data coverage need to achieve exact solution

#### True model

#### Estimated model



#### Estimated model

#### Estimated model



streaks due to minor data gaps they disappear if ray density is doubled

# but what if the observational geometry is poor

# so that broads swaths of rays are missing ?

## complete angular coverage (B)







#### incomplete angular coverage









#### Part 2

## vibrational problems

## statement of the problem

Can you determine the structure of an object just knowing the characteristic frequencies at which it vibrates?



the Fréchet derivative of frequency with respect to velocity is usually computed using *perturbation theory* 

hence a quick discussion of what that is ...

## perturbation theory

a technique for computing an approximate solution to a complicated problem, when

1. The complicated problem is related to a simple problem by a small perturbation

2. The solution of the simple problem must be known

simple example

perturbation theory for a quadratic equation with a small first order term

$$x^2 + \varepsilon bx - c^2 = 0$$

assume  $\varepsilon$  is small and write solution as a power series in  $\varepsilon$ 

$$x = x_0 + \varepsilon x_1 + \varepsilon^2 x_2 + \dots$$

where  $x_0$  solves the equation when  $\mathcal{E}=0$ 

$$x_0^2 - c^2 = 0$$

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where  $x_0$  solves the equation when  $\mathcal{E}=0$ 

$$x_0^2 - c^2 = 0$$
  
we know the  
solution to this  
equation:  $x_0 = \pm c$ 

plug the power series into the original equation  $x^{2} + \varepsilon bx - c^{2} = 0$ 

 $(x_0 + \varepsilon x_1 + \cdots)^2 + \varepsilon b(x_0 + \varepsilon x_1 + \cdots) - c^2 = 0$ 

#### multiply out

 $x_0^2 + 2\varepsilon x_1 x_0 + \varepsilon b x_0 - c^2 + O(\varepsilon^2) = 0$ 

Group terms of equal power in  $\varepsilon$ 

 $(x_0^2 - c^2)\varepsilon^0 + (2x_1x_0 + bx_0)\varepsilon + O(\varepsilon^2) = 0$ 

since  $\varepsilon$  is arbitrary, coefficients of each power of  $\varepsilon$  must be individually zero

$$\varepsilon^0$$
:  $(x_0^2 - c^2) = 0 \rightarrow x_0 = \pm c$ 

$$\varepsilon^1$$
:  $(2x_1x_0 + bx_0) = 0 \rightarrow x_1 = -b/2$ 

so solution is approximately

$$x = x_0 + \varepsilon x_1 + O(\varepsilon^2) = \pm c - \frac{\varepsilon b}{2} + O(\varepsilon^2)$$

since  $\varepsilon$  is arbitrary, coefficients of each power of  $\varepsilon$  must be individually zero

$$\varepsilon^0: (x_0^2 - c^2) = 0 \rightarrow x_0 = \pm c$$

$$\varepsilon^1$$
:  $(2x_1x_0 + bx_0) = 0 \rightarrow x_1 = -b/2$ 

so solution is approximately

$$x = x_0 + \varepsilon x_1 + O(\varepsilon^2) = \pm c - \frac{\varepsilon b}{2} + O(\varepsilon^2)$$

note this agrees with the exact result computed from the quadratic formula

$$x = -\frac{\varepsilon b}{2} \pm \frac{\sqrt{\varepsilon^2 b^2 + 4c^2}}{2} = \pm c - \frac{\varepsilon b}{2} + O(\varepsilon^2)$$
#### Here's the actual vibrational problem

acoustic equation with spatially variable sound velocity *v* 

$$-\omega_n^2 p_n(\mathbf{x}) = v^2(\mathbf{x}) \,\nabla^2 p_n(\mathbf{x})$$

# acoustic equation with spatially variable sound velocity *v*

 $-\omega_n^2 p_n(\mathbf{x}) = v^2(\mathbf{x}) \nabla^2 p_n(\mathbf{x})$ patterns of vibration or eigenfunctions frequencies of vibration or or modes eigenfrequencies

# assume velocity can be written as a perturbation around some simple structure $v^{(0)}(\mathbf{x})$

#### $v(\mathbf{x}) = v^{(0)}(\mathbf{x}) + \varepsilon v^{(1)}(\mathbf{x}) + \dots$

#### eigenfunctions known to obey orthonormality relationship

# $\int p_n(\mathbf{x}) p_m(\mathbf{x}) v^{-2}(\mathbf{x}) \mathrm{d}^3 x = \delta_{nm}$

now represent eigenfrequencies and eigenfunctions as power series in  $\varepsilon$ 

$$\omega_n = \omega_n^{(0)} + \varepsilon \omega_n^{(1)} + \varepsilon^2 \omega_n^{(2)} + \cdots$$

$$p_n(\mathbf{x}) = p_n^{(0)}(\mathbf{x}) + \varepsilon p_n^{(1)}(\mathbf{x}) + \varepsilon^2 p_n^{(2)}(\mathbf{x}) + \cdots$$

now represent eigenfrequencies and eigenfunctions as power series in  $\varepsilon$ 

$$\omega_n = \omega_n^{(0)} + \varepsilon \omega_n^{(1)} + \varepsilon^2 \omega_n^{(2)} + \cdots$$

 $p_n(\mathbf{x}) = p_n^{(0)}(\mathbf{x}) + \varepsilon p_n^{(1)}(\mathbf{x}) + \varepsilon^2 p_n^{(2)}(\mathbf{x}) + \cdots$ represent first-order perturbed shapes as sum of unperturbed shapes  $= \sum b_{nm} p_m^{(0)}$  $\omega_m \neq \omega_n$ 

### plug series into original differential equation

group terms of equal power of  $\epsilon$ 

solve for first-order perturbation in eigenfrequencies  $\omega_n^{(1)}$ and eigenfunction coefficients  $b_{nm}$ 

(use orthonormality in process)

#### result

$$\omega_n^{(1)} = \omega_n^{(0)} \int \left[ p_n^{(0)}(\mathbf{x}) \right]^2 \left[ v^{(0)}(\mathbf{x}) \right]^{-3} v^{(1)}(\mathbf{x}) \, \mathrm{d}^3 x$$
$$b_{nm} = \frac{2 \left( \omega_m^{(0)} \right)^2}{\left( \omega_m^{(0)} \right)^2 - \left( \omega_n^{(0)} \right)^2} \int p_n^{(0)}(\mathbf{x}) \, p_m^{(0)}(\mathbf{x}) \left[ v^{(0)}(\mathbf{x}) \right]^{-3} \, v^{(1)}(\mathbf{x}) \, \mathrm{d}^3 x$$

#### result for eigenfrequencies

$$\omega_n^{(1)} = \omega_n^{(0)} \int \left[ p_n^{(0)}(\mathbf{x}) \right]^2 \left[ v^{(0)}(\mathbf{x}) \right]^{-3} v^{(1)}(\mathbf{x}) \, \mathrm{d}^3 x$$

#### write as standard inverse problem

$$\omega_n^{(1)} = \int G_n(\mathbf{x}) \, v^{(1)}(\mathbf{x}) \, \mathrm{d}^3 x \quad \text{with} \quad G_n(\mathbf{x}) = \omega_n^{(0)} \left[ p_n^{(0)}(\mathbf{x}) \right]^2 \left[ v^{(0)}(\mathbf{x}) \right]^{-3}$$

#### standard continuous inverse problem

$$\omega_n^{(1)} = \int G_n(\mathbf{x}) \, v^{(1)}(\mathbf{x}) \, \mathrm{d}^3 x \quad \text{with} \quad G_n(\mathbf{x}) = \omega_n^{(0)} \left[ p_n^{(0)}(\mathbf{x}) \right]^2 \left[ v^{(0)}(\mathbf{x}) \right]^{-3}$$

#### standard continuous inverse problem



#### standard continuous inverse problem

data kernel or Fréchet derivative  $\omega_n^{(1)} = \int G_n(\mathbf{x}) v^{(1)}(\mathbf{x}) d^3x \quad \text{with} \quad G_n(\mathbf{x}) = \omega_n^{(0)} [p_n^{(0)}(\mathbf{x})]^2 [v^{(0)}(\mathbf{x})]^{-3}$ depends upon the unperturbed velocity structure, the unperturbed eigenfrequency and the unperturbed mode

#### 1D organ pipe





#### solution to unperturbed problem

$$p_n^{(0)}(x) = \frac{2[v^{(0)}]^2}{h} \sin\left\{\frac{(n - \frac{1}{2})\pi}{h}x\right\}$$

$$\omega_n^{(0)} = \frac{\pi (n - \frac{1}{2})v^{(0)}}{h} \quad \text{with } n = 1, 2, 3, \cdots$$

#### velocity structure



#### How to discretize the model function?

our choice is very simple

**m** is velocitty function evaluated at sequence of points equally spaced in *x* 

#### the data a list of frequencies of vibration



true, unperturbed true, perturbed observed = true, perturbed + noise

#### the data kernel



 $\omega_i$ 

#### Solution Possibilities

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#### Solution Possibilities

- Damped Least Squares (implements smallness): Matrix G is not sparse use bicg() with damped least squares function
  Our choice
- 2. Weighted Least Squares (implements smoothness): Matrix F consists of G plus second derivative smoothing use bicg() with weighted least squares function





#### the model resolution matrix



#### the model resolution matrix



This problem has a type of nonuniqueness

that arises from its symmetry

a positive velocity anomaly at one end of the organ pipe

trades off with a negative anomaly at the other end

this behavior is very common and is why eigenfrequency data are usually supplemented with other data

e.g. travel times along rays

that are not subject to this nonuniqueness

#### Part 3

#### determining mean directions

#### statement of the problem

you measure a bunch of directions (unit vectors) what's their mean?



what's a reasonable probability density function for directional data?

#### Gaussian doesn't quite work because its defined on the wrong interval $(-\infty, +\infty)$

#### coordinate system



#### distribution should be symmetric in $\phi$

### Fisher distribution similar in shape to a Gaussian but on a sphere





### Fisher distribution similar in shape to a Gaussian but on a sphere



#### solve by

#### direct application of

#### principle of maximum likelihood

#### maximize joint p.d.f. of data

$$p(\theta, \phi) = \left[\frac{\kappa}{4\pi \sinh(\kappa)}\right]^N \exp\left[\kappa \sum_{i=1}^N \cos(\theta_i)\right] \prod_{i=1}^N \sin(\theta_i)$$

### with respect to $\kappa$ and $\cos(\theta)$

# **x**: Cartesian components of observed unit vectors

**m**: Cartesian components of central unit vector; must constrain  $|\mathbf{m}|=1$ 

 $\cos(\theta_i) = \mathbf{x}^{\mathrm{T}} \mathbf{m} = [x_i m_1 + y_i m_2 + z_i m_3]$
### likelihood function

 $L = \log(p) = N \log(\kappa) - N \log(4\pi) - N \log[\sinh(\kappa)]$ 

$$+\kappa \sum_{i=1}^{N} [x_i m_1 + y_i m_2 + z_i m_3] + \sum_{i=1}^{N} \log [\sin(\theta_i)]$$
  
constraint  
$$C = \sum_i m_i^2 - 1 = 0$$

#### unknowns

**т**, к

## Lagrange multiplier equations

$$\kappa \sum_{i} x_{i} - 2\lambda m_{1} = 0$$

$$\kappa \sum_{i} y_i - 2\lambda m_2 = 0$$

$$\kappa \sum_{i} z_{i} - 2\lambda m_{3} = 0$$

$$\frac{N}{\kappa} - N\frac{\cosh(\kappa)}{\sinh(\kappa)} + \sum_{i=1}^{N} [x_i m_1 + y_i m_2 + z_i m_3] = 0$$

### Results

$$[m_1, m_2, m_3]^{\mathrm{T}} = \frac{[\sum_i x_i, \sum_i y_i, \sum_i z_i]^{\mathrm{T}}}{\{(\sum_i x_i)^2 + (\sum_i y_i)^2 + (\sum_i z_i)^2\}^{\frac{1}{2}}}$$

$$\kappa \approx \frac{N}{N - \sum_{i} \cos(\theta_i)}$$
 va

valid when  $\kappa > 5$ 

### Results

$$[m_1, m_2, m_3]^{\mathrm{T}} = \frac{[\sum_i x_i, \sum_i y_i, \sum_i z_i]^{\mathrm{T}}}{\{(\sum_i x_i)^2 + (\sum_i y_i)^2 + (\sum_i z_i)^2\}^{\frac{1}{2}}}$$

central vector is parallel to the vector that you get by putting all the observed unit vectors end-to-end

# Solution Possibilities

Determine **m** by evaluating simple formula

1. Determine  $\kappa$  using simple but approximate formula only valid when  $\kappa > 5$ 

our choice

2. Determine  $\kappa$  using bootstrap method

#### Application to Subduction Zone Stresses

#### Determine the mean direction of P-axes of deep (300-600 km) earthquakes in the Kurile-Kamchatka subduction zone



data
central direction
bootstrap