

EESC 9945

Geodesy with the Global Positioning
System

Class 6: *Point Positioning using Pseudorange*

GPS Positioning Solutions

- **Point Positioning:** Determination of the coordinates of a point with respect to an implied coordinate system
- **Relative Positioning:** Determination of the location of one point with respect to another
- Pseudorange solutions are point positions
- What is the “implied coordinate system?”

Least-square solutions

1. Data y and measurement errors $\epsilon \rightarrow \Lambda = \langle \epsilon \epsilon^T \rangle$
2. Model $y = f(x; t, \dots) + \epsilon$
3. Parameters x to be estimated
4. *a priori* parameter estimates x_o
5. Pre-fit residuals $\Delta y = y - y_o = y - f(x_o)$
6. Partial derivatives $\partial f(x) / \partial x|_{x=x_o} \rightarrow A$
7. Solve $\hat{x} = x_o + \Delta \hat{x} = x_o + (A^T \Lambda^{-1} A)^{-1} A^T \Lambda^{-1} \Delta y$
8. Iterate if non-linear (use \hat{x} as new *a priori* and go to step 5)
9. **NOTE: Need matrix multiplication/inversion routines**

RINEX data

- RINEX naming convention `ssssdddf.yyt`
- `ssss`: 4-character site ID
- `ddd`: Day of year (001–366) of first record
- `f`: File sequence number (`f = 0` for daily file)
- `yy`: two-digit year
- `t`: File type (= “o” for observation file)

RINEX observation header

```
2.11          OBSERVATION DATA    G (GPS)          RINEX VERSION / TYPE
teqc 2008Feb15  UNAVCO Archive Ops  20080410 14:43:33UTCPGM / RUN BY / DATE
Solaris 5.10|UltraSparc IIIi|cc -xarch=v9 SC5.8|=+|*Sparc  COMMENT
BIT 2 OF LLI FLAGS DATA COLLECTED UNDER A/S CONDITION  COMMENT
XXXX                                                MARKER NAME
                                                MARKER NUMBER
Davis, James          EESC 9945          OBSERVER / AGENCY
3841A24273           TRIMBLE 4000SSI      7.19B          REC # / TYPE / VERS
0220136818           TRM29659.00         SCIS           ANT # / TYPE
          0.0000          0.0000          0.0000         APPROX POSITION XYZ
          0.0000          0.0000          0.0000         ANTENNA: DELTA H/E/N
          1      1                                     WAVELENGTH FACT L1/2
          7      L1      L2      C1      P2      P1      S1      S2      # / TYPES OF OBSERV
          30.0000                                       INTERVAL
RINEX file created by UNAVCO GPS Archive.          COMMENT
For more information contact archive@unavco.org    COMMENT
                                                COMMENT
End of DB comments                                COMMENT
SNR is mapped to RINEX snr flag value [1-9]      COMMENT
L1: 3 -> 1; 8 -> 5; 40 -> 9                       COMMENT
L2: 1 -> 1; 5 -> 5; 60 -> 9                       COMMENT
2008      4      9      0      0      30.0000000    GPS          TIME OF FIRST OBS
                                                END OF HEADER
```

Types of observations

7 L1 L2 C1 P2 P1 S1 S2 # / TYPES OF OBSERV

- Lists all observations in file, in order
- L1 and L2: Carrier-beat phase in cycles
- C1: C/A code pseudorange in meters
- P1 and P2: P-code pseudorange in meters
- S1 and S2: SNR
- Most receivers do not have L1 P-code pseudorange, so to calculate ion-free pseudorange we combine C1 and P2

RINEX observation block

- First line of block: Epoch, flag, and satellite list

```
08 4 9 1 47 0.000000 0 11G31G 7G 6G25G16G13G 4G 3G20G23G27
```

- For each satellite, **five** observation types per line in order defined in the “observation type” header line
- Thus, seven observation types require two lines
- Each observation type has three values (F14.3,I1,I1): Observation, LLI flag, and signal strength flag

```
-12295400.45442 -9731348.60754 25175212.6024 25175215.5784  
4.7504 4.5004
```

```
L1-->          L2-->          C1-->          P2-->          P1-->  
S1-->          S2-->
```

Ionosphere-free observation and model

- As per last week, we first form the ionosphere-free ionospheric observable from the C1 and P2 pseudoranges for each satellite s at each epoch t_k

$$R_{\text{LC}}^s(t_k) = \left[1 - \left(\frac{f_2}{f_1} \right)^2 \right]^{-1} \left[R_1^s(t_k) - \left(\frac{f_2}{f_1} \right)^2 R_2^s(t_k) \right]$$

- The model we had developed for the pseudorange (in meters) that neglects ionosphere is

$$R_{\text{LC}}^s(t_k) = \tilde{\rho}_r^s(t_k) + \tilde{\delta}_r(t_k) - \tilde{\delta}^s(t_k) + \tau^z m(\epsilon^s(t_k))$$

- $\tilde{\rho} = \rho_o / \left[1 + \vec{\beta} \cdot \hat{\rho}_o \right]$ is the site-satellite distance corrected for the satellite motion
- Topocentric vector $\vec{\rho}_o = \vec{x}^s - \vec{x}_r$

Ionosphere-free observations

- $\tilde{\delta}_r(t_j)$ is the unknown receiver clock error in meters
(= $c\delta_r$)
- $\tilde{\delta}^s(t_j)$ is the satellite clock error in meters
- τ^z is the zenith (neutral) atmospheric propagation delay
- $m(\epsilon)$ is the mapping function depending on elevation angle ϵ

Pseudorange solution: Parameters

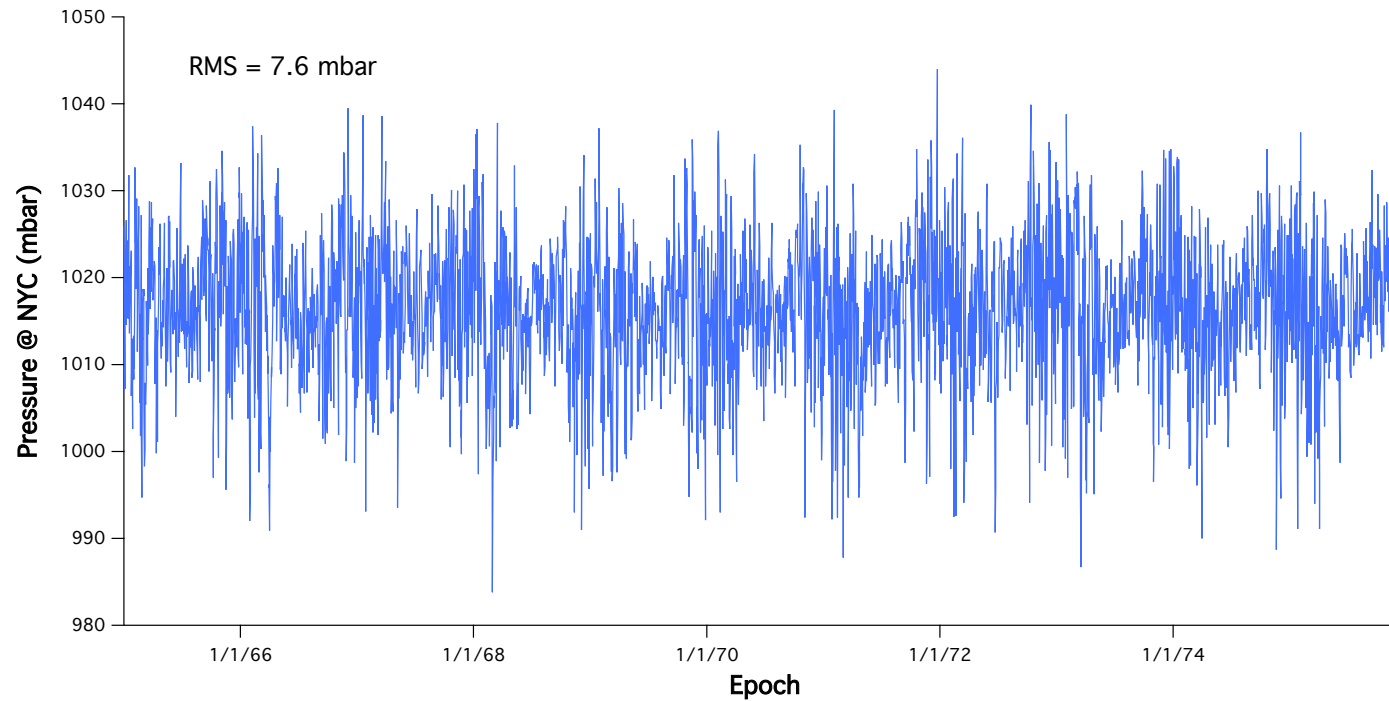
\vec{x}_r	Site vector	Unknown	m	3
\vec{x}_s	Satellite vector	Calc from nav file	m	
$\vec{\beta}$	Satellite velocity	Calc from nav file	1	
$\tilde{\delta}^s$	Satellite clock error	Calc from nav file	m	
δ_r	Receiver clock error	Unknown	m	N_{epoch}
τ^z	Zenith delay	Known at $\sim 0.1\text{--}0.2$ m	m	
$m(\epsilon)$	Mapping function	Formula	1	
ϵ	Elevation angle	Calc from \vec{x}_r, \vec{x}^s	rad	

Total number of parameters: $N_{\text{par}} = 3 + N_{\text{epoch}}$

Zenith delay

- The zenith delay is too large to ignore; the data errors (1–10 m) make it hard to estimate accurately
- We know $\tau^z \simeq (2.2 \text{ mm/mbar}) \times P_{\text{surf}} + \tau_{\text{wet}}^z$

Pressure at NYC



7.6 mbar \rightarrow 17 mm hydrostatic delay

Hydrostatic Zenith Delay

- Typical RMS pressure 3–13 mbar, 7–30 mm
- Sometimes seasonal
- **Mean** pressures can vary site-to-site by > 200 mbar (~ 0.5 m), depending on site altitude
- Reasonable expression for Earth is $P(H) \simeq P_0 e^{-H/H_0}$ with $P_0 \simeq 1013$ mbar and $H_0 \simeq 7.5$ km
- Large pressure drops can occur in storms, hurricanes, etc.

Wet Zenith Delay

- Mean varies $\sim 5\text{--}30$ cm
- RMS $\sim 4\text{--}10$ cm
- Highly dependent on climate
- Strong seasonal variability
- Scale height of ~ 2 km

Zenith Delay for Solutions

- For pseudorange solution, we'd like $\sigma_{\tau^z} \lesssim 0.1\text{--}0.3$ m
- Could use model and no estimation
- Simple model: $\tau^z \simeq (2.31 \text{ m}) e^{-H/H_h} + (0.15 \text{ m}) e^{-H/H_w}$
with $H_h = 7.5$ km and $H_w = 2$ km
- Probably $\sigma_{\tau^z} \lesssim 20$ cm; OK for pseudorange solutions
- For phase solutions, we'd like $\sigma_{\tau^z} \lesssim 5$ mm

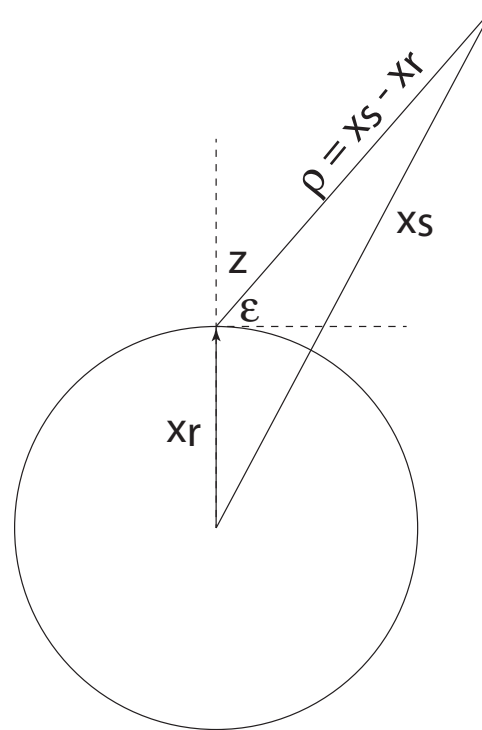
Mapping Function

- Want something slightly better than cosecant law
- One of simplest to use is *Davis et al.* [1985]:

$$m(\epsilon) = \frac{1}{\sin \epsilon + \frac{a}{\tan \epsilon + \frac{b}{\sin \epsilon + c}}} \quad (1)$$

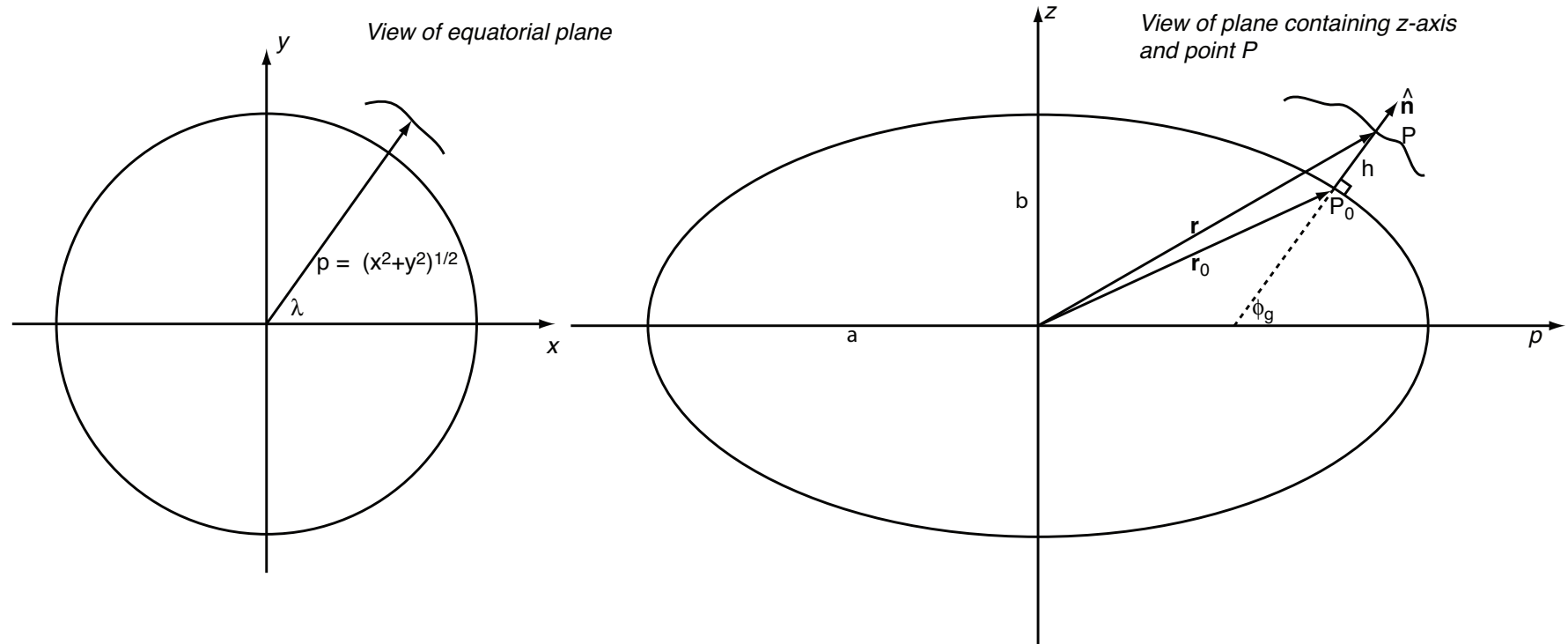
- Nominal constants: $a = 0.001185$; $b = 0.001144$;
 $c = -0.0090$

Calculation of Elevation Angle

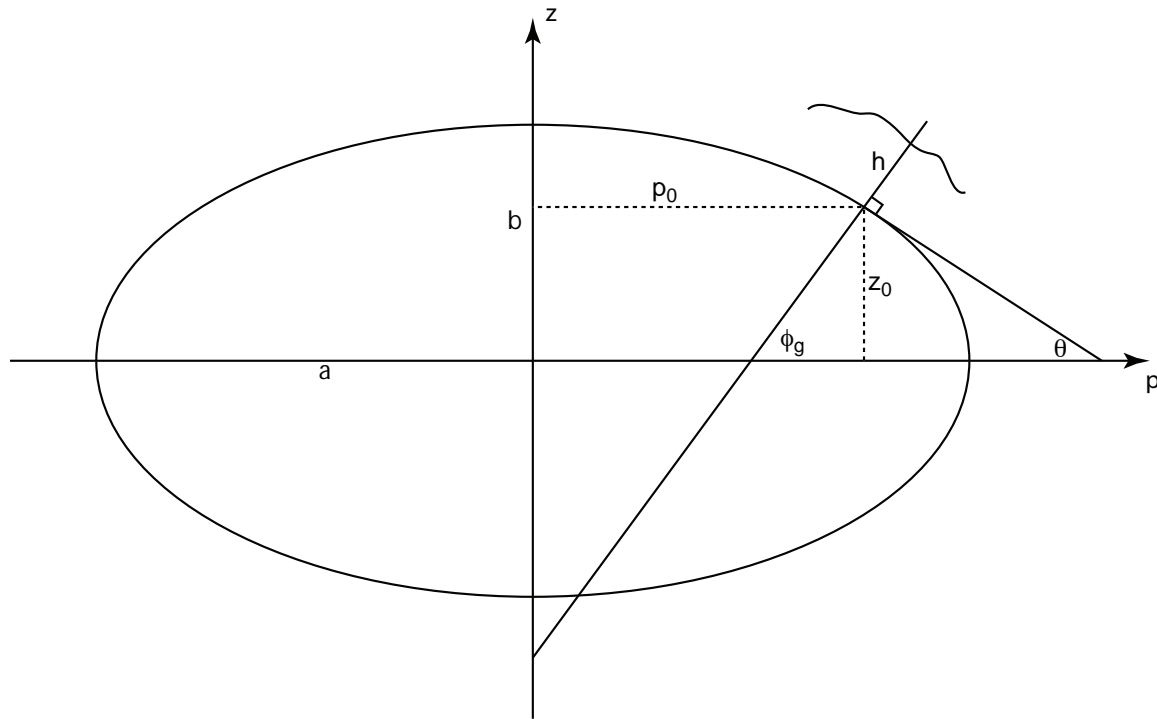


Spherical Earth approximation: $\hat{x}_r \cdot \hat{\rho} = \cos z = \sin \epsilon$

Calculations on a Reference Ellipsoid



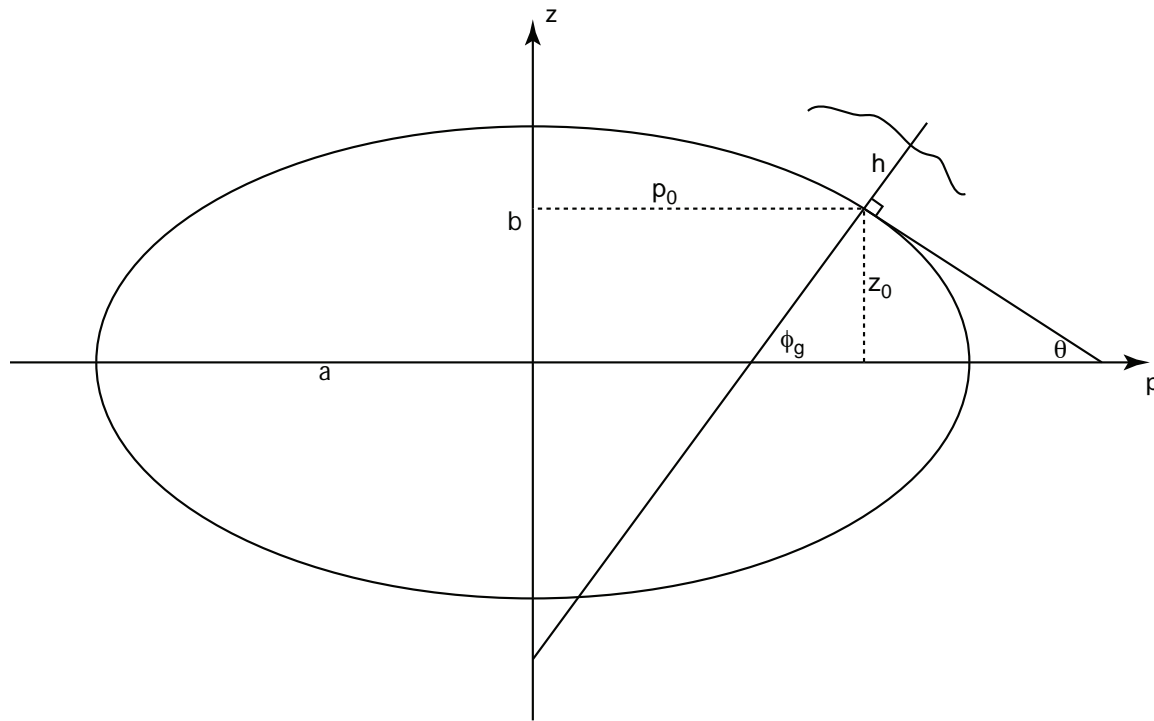
Coordinates are longitude λ , geodetic latitude ϕ_g , and ellipsoidal height h .



The projection along the local normal onto the ellipsoid is $p = p_0$
and $z = z_0$

p_0 and z_0 are related by

$$\frac{p_0^2}{a^2} + \frac{z_0^2}{b^2} = 1$$

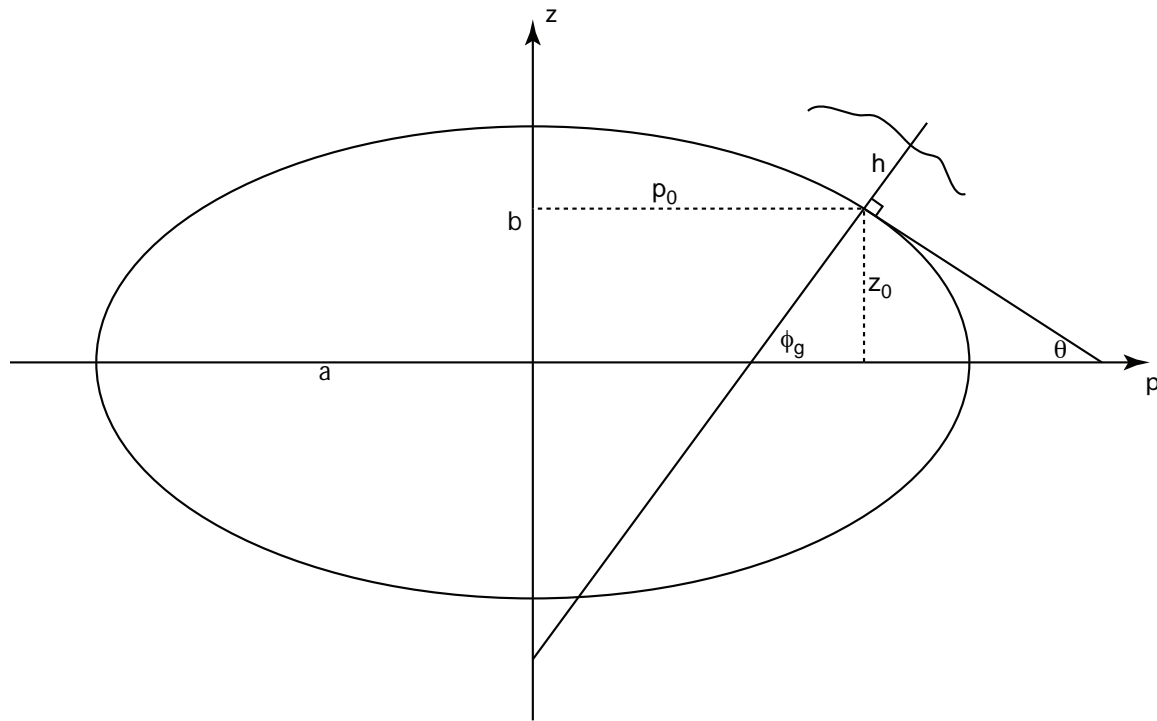


The relation to the geodetic latitude (angle of orthogonal projection of point onto ellipsoid with z -axis) is

$$-\frac{dz_0}{dp_0} = \tan \theta = \tan \left(\frac{\pi}{2} - \phi_g \right) = \cot \phi_g$$

or

$$\frac{dp_0}{dz_0} = -\tan \phi_g$$

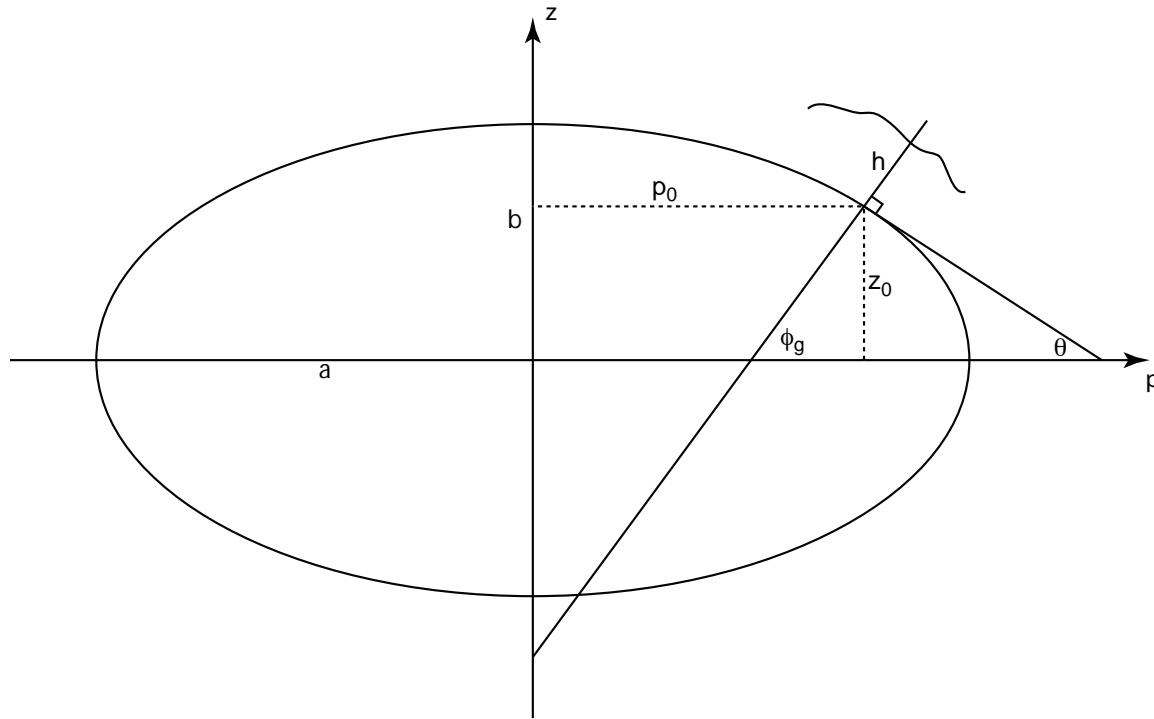


Differentiating eq. of ellipsoid ($\frac{p_0^2}{a^2} + \frac{z_0^2}{b^2} = 1$) w.r.t. z_0 , we get

$$2\frac{p_0}{a^2}(-\tan \phi_g) + 2\frac{z_0}{b^2} = 0$$

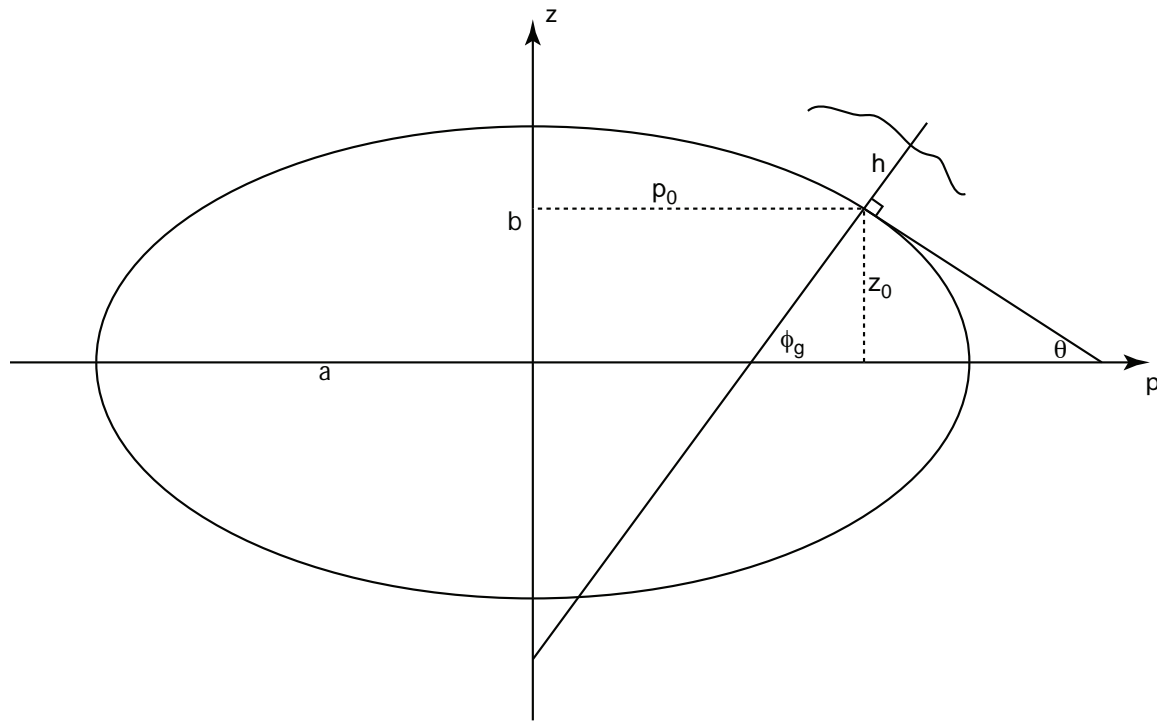
or

$$z_0 = \left(\frac{b}{a}\right)^2 p_0 \tan \phi_g$$



Substituting back into eq. and solving for p_0 yields

$$p_0 = \frac{a^2 \cos \phi_g}{\sqrt{a^2 \cos^2 \phi_g + b^2 \sin^2 \phi_g}}$$

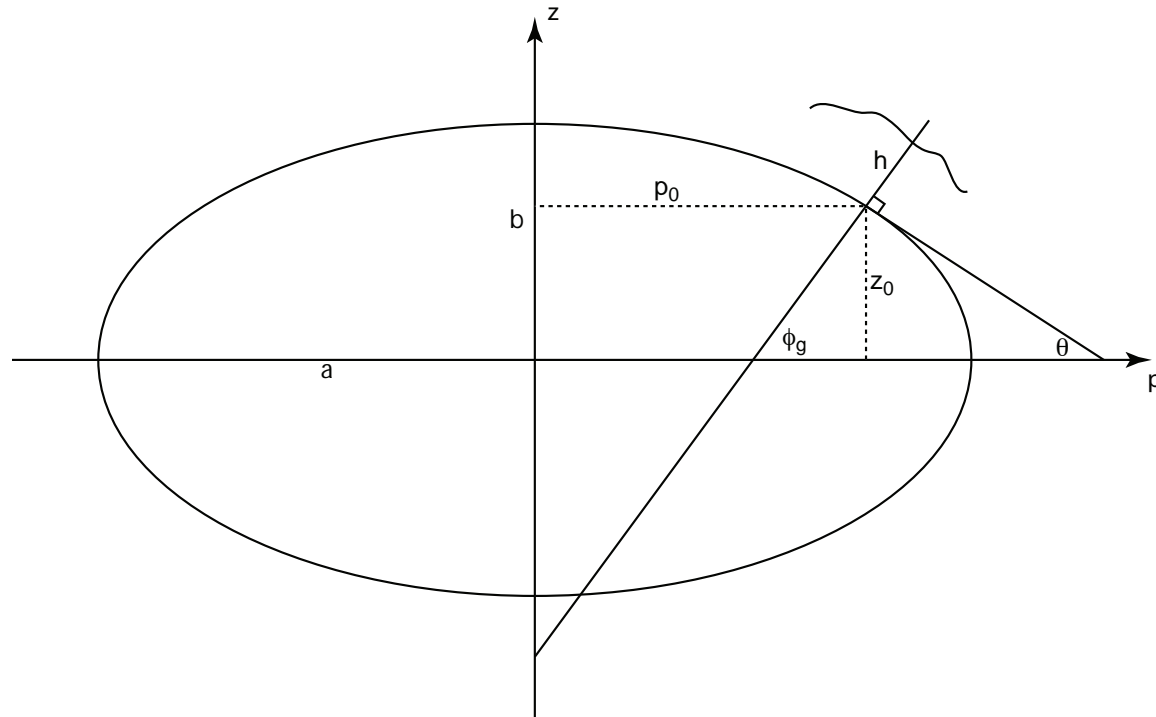


Denoting

$$N_o = \frac{a^2}{\sqrt{a^2 \cos^2 \phi_g + b^2 \sin^2 \phi_g}}$$

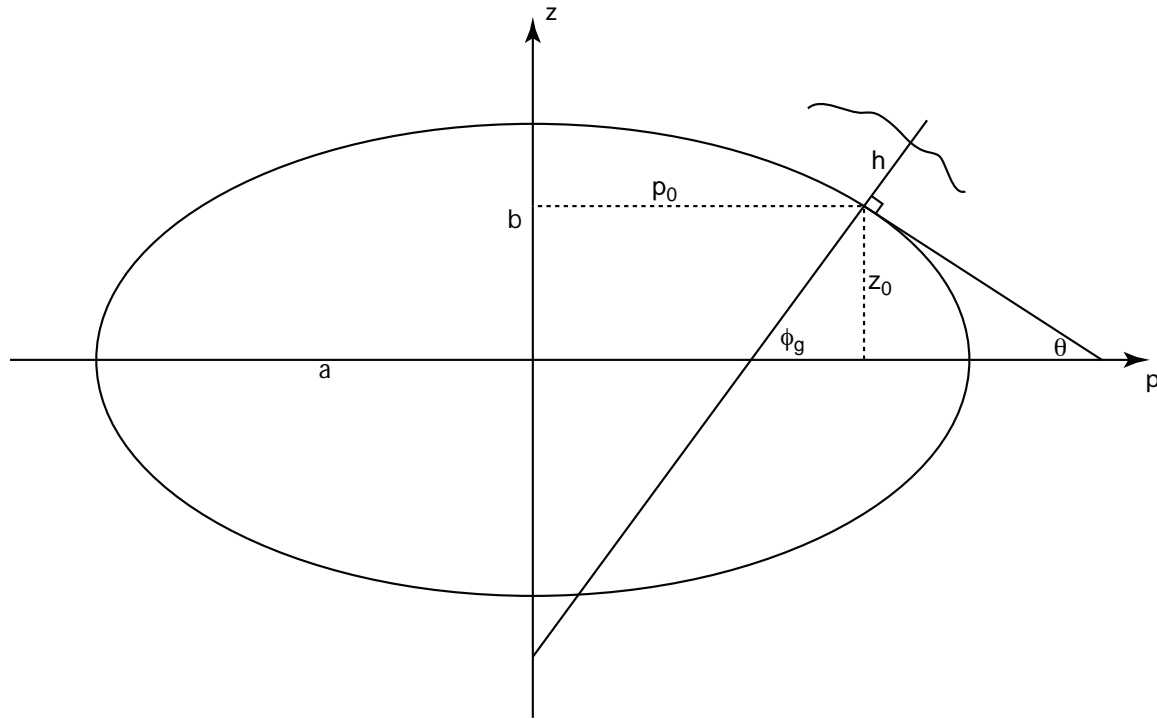
(the prime radius of curvature), then

$$p_o = N_o \cos \phi_g \quad \text{and} \quad z_o = \left(\frac{b}{a}\right)^2 N_o \sin \phi_g$$



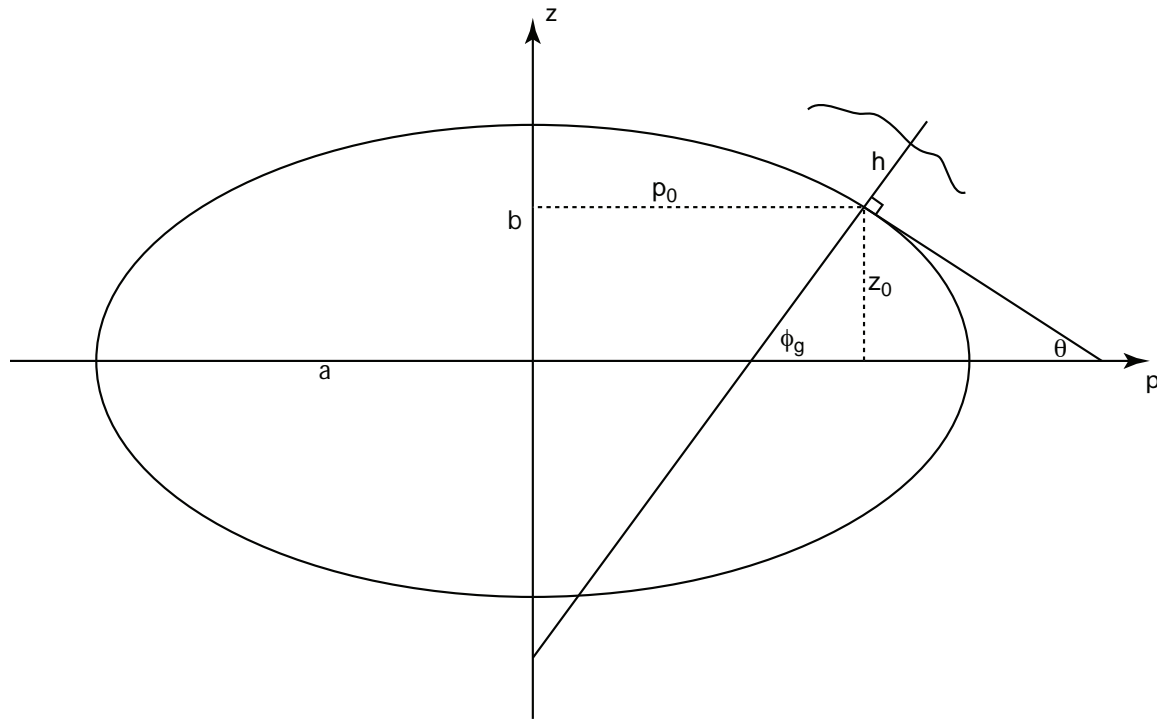
Now we must add the piece between the ellipsoid and the point on the terrain.

The unit normal to the ellipsoid is $\hat{n} = \hat{p} \cos \phi_g + \hat{z} \sin \phi_g$,



The terrain point is thus

$$\vec{x} = p_0 \hat{p} + z_0 \hat{z} + h \hat{n} = \hat{p} (N_o + h) \cos \phi_g + \hat{z} \left(\frac{b^2}{a^2} N_o + h \right) \sin \phi_g$$



So the 3-D position is

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} (N_o + h) \cos \phi_g \cos \lambda \\ (N_o + h) \cos \phi_g \sin \lambda \\ \left[\frac{b^2}{a^2} N_o + h \right] \sin \phi_g \end{pmatrix}$$

Calculation of Ellipsoidal Coordinates

- Previous gives $(\lambda, \phi_g, h) \rightarrow (x, y, z)$
- No closed-form expression for ϕ_g unless $h = 0$
- To perform the reverse use iterative approach for ϕ_g and h :
 1. $h^{(0)} = 0$; $\tan \phi_g^{(0)} = z / \sqrt{x^2 + y^2}$; calculate $N_o^{(0)}$ from $\phi_g^{(0)}$
 2. $\tan \phi_g^{(k+1)} = \left(N_o^{(k)} + h^{(k)} \right) z / \left(\frac{b^2}{a^2} N_o^{(k)} + h^{(k)} \right) \sqrt{x^2 + y^2}$
 3. $N_o^{(k+1)}$ from $\phi_g^{(k+1)}$
 4. $h^{(k+1)} = \sqrt{x^2 + y^2} / \cos \phi_g^{(k+1)} - N_o^{(k+1)}$
 5. Go to Step 2 until converged
- Solution for λ is exact

WGS 84 Ellipsoid

- World Geodetic System: Conventional Terrestrial Reference System
- Adopted by GPS
- Definition of coordinate axes:
 - Origin: Earth's center of mass
 - Z-Axis: Direction of IERS Reference Pole
 - X-Axis: Intersection of IERS Reference Meridian and plane passing through origin and normal to Z-axis.
 - Y-Axis: Completes right-handed, Earth-centered Earth-fixed orthogonal coordinate system
- Uses $a = 6378137.0$ m
- Flattening $1/f = 298.257223563 = a/(a - b)$
- Therefore $b = 6356752.3142$ m

A priori parameter estimates

- Needed for:
 1. Calculation of $\Delta y = y - y_0$
 - (a) Topocentric vector
 - (b) Elevation angle ϵ for mapping function
 - (c) Station height for zenith delay
 2. Calculation of design matrix A (partial derivatives w.r.t. site coordinates)
- Need: A priori site coordinates (x - y - z) and a priori site clock errors (N_{epoch} values)
- What should we use?