Simple Zero-D thermal model of a radioactive, convecting earth. Bill Menke, February 2012.

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Heat in mantle Q = V c_p \rho T = CT
(assumes that mantle is convectively mixed so it is isothermal)
        temperature T
        Volume of mantle V = 4 \pi R^3 / 3
        heat capacity per unit mass = cp
        density p
        summary constant C = V c_p \rho
        also define To as initial temperature of mantle
Heat production of mantle H = V h \rho \exp(-ct) = H_0 \exp(-ct)
        (assume heat production decays exponentially with time)
        Initial heat production per unit mass h
         decay rate c
        Summary constant H_0 = V h \rho
Heat loss through conductive lithosphere q = k A T / L = K T
(assumes lithosphere thin enough that equilibration time is short)
        thermal conductivity k
        thickness of conductive lithosphere = L
        area of surface of earth A = 4 \pi R^2
        summary constant K = k A / L
Conservation of energy
        rate of increase in heat with time t = heat produced - heat lost through surface
        dQ/dt = H(t) - q
        dT/dt = (H_0/C) \exp(-ct) - (K/C) T
        dT/dT + aT = b \exp(-ct) with a = (K/C) and b = (H_0/C)
Solution of ODE (see http://www.sosmath.com/diffeq/first/lineareq/lineareq.html)
        p(t) = a and r(t) = b exp(-ct)
       u(t) = \exp(\int p \, dt) = \exp(at)
        \int u(t) r(t) dt = \int exp(at) b exp(-ct) dt = \int b exp\{(a-c)t\} dt = (b/(a-c))exp((a-c)t\}
        T(t) = (\int u(t) r(t) dt + C) / u(t) = [\{b/(a-c)\} exp\{(a-c)t\} + C] exp(-at)
        T(t=0) = \{b/(a-c)\} + C = T_0
        C = T_0 - \{b/(a-c)\}
        T(t) = [{b/(a-c)} exp{(a-c)} + T_0 - {b/(a-c)}] exp(-at)
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## Analysis

Suppose no heat production, then b=0 and  $T(t) = T_0 \exp(-at)$ , so temperature declines exponentially with time at a rate a=(K/C).

Suppose non-zero heat production (b>0) with infinite half life (c=0), then at long time T(t) = b/a; that is, an equilibrium temperature is reached where the mantle heat production is in equilibrium with the conductive loss through the lithosphere.

## **Numerical Values**

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R = 64000000 \text{ m}
V = 4 \pi R^3 / 3 = 1.04686E + 21 m^3
co= 950 J/kg-K
\rho = 5000 \, \text{kg/m}^3
C = V c_p \rho = 4.9726E + 27 J/K
h (today) = 4E-12 J/s-kg
h(0) = 2E-11 J/s-kg (assume factor of 5 higher than today)
H_0 = V h \rho = 1.1E + 14 J/s
k = 2.7 \text{ J/s-m-K}
L = 1E5 \text{ m} (100 \text{ km thick lithosphere})
A = 4 \pi R^2 = 5.14E+14 m^2
K = k A / L = 1.3E+10 J/s-K
a= (K/C)= 2.7E-18 1/s
1/a = 1.2E + 10 yrs
so characteristic cooling time in absence or radioactivity is very long, 10 billion years
b = (H_0/C) = 3.15789E-14 \text{ K/s}
b/a = 1.1667E+04 K
so if radioactivity didn't decay with time, mantle would reach 10,000 K and melt
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Quasi-Kelvin's Age of the earth, time it takes heat flow to decline to given value f J/m2-s in the absence of radioactive heating

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T_0 = (1350 + 273 + 150) = 1773 \text{ K} f= 0.03 \text{ J/m2-s} f= k (T_0 /L) \exp(-at) = f0 \exp(-at) \text{ with } f0 = k (T_0 /L) = 0.045 \text{ J/m2-s} t = -ln(f/f0) /a = 5.6E9 \text{ yrs} So convection alone can give a long age of the earth; radioactivity not needed
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