

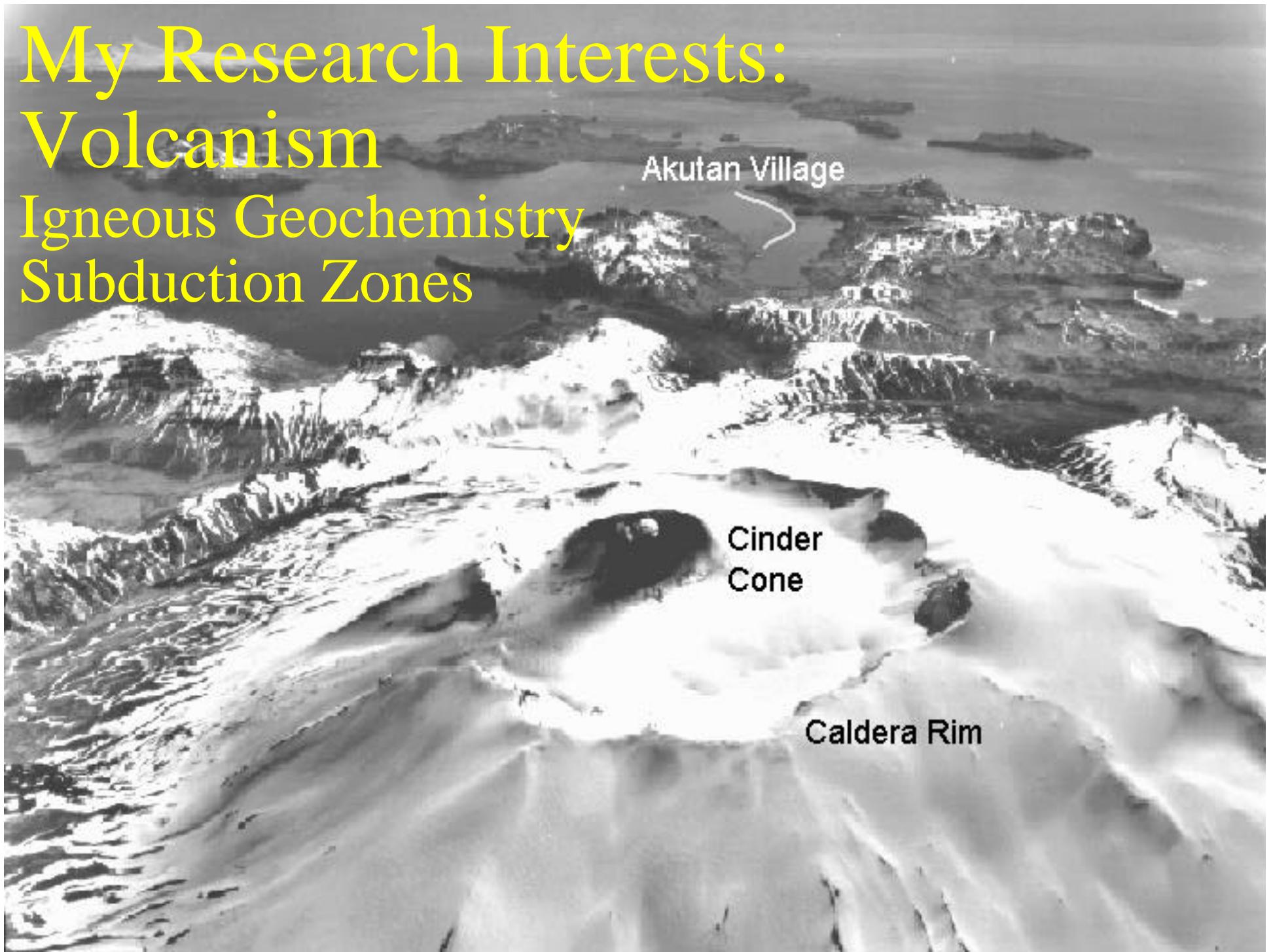
EESC 2200
The Solid Earth System

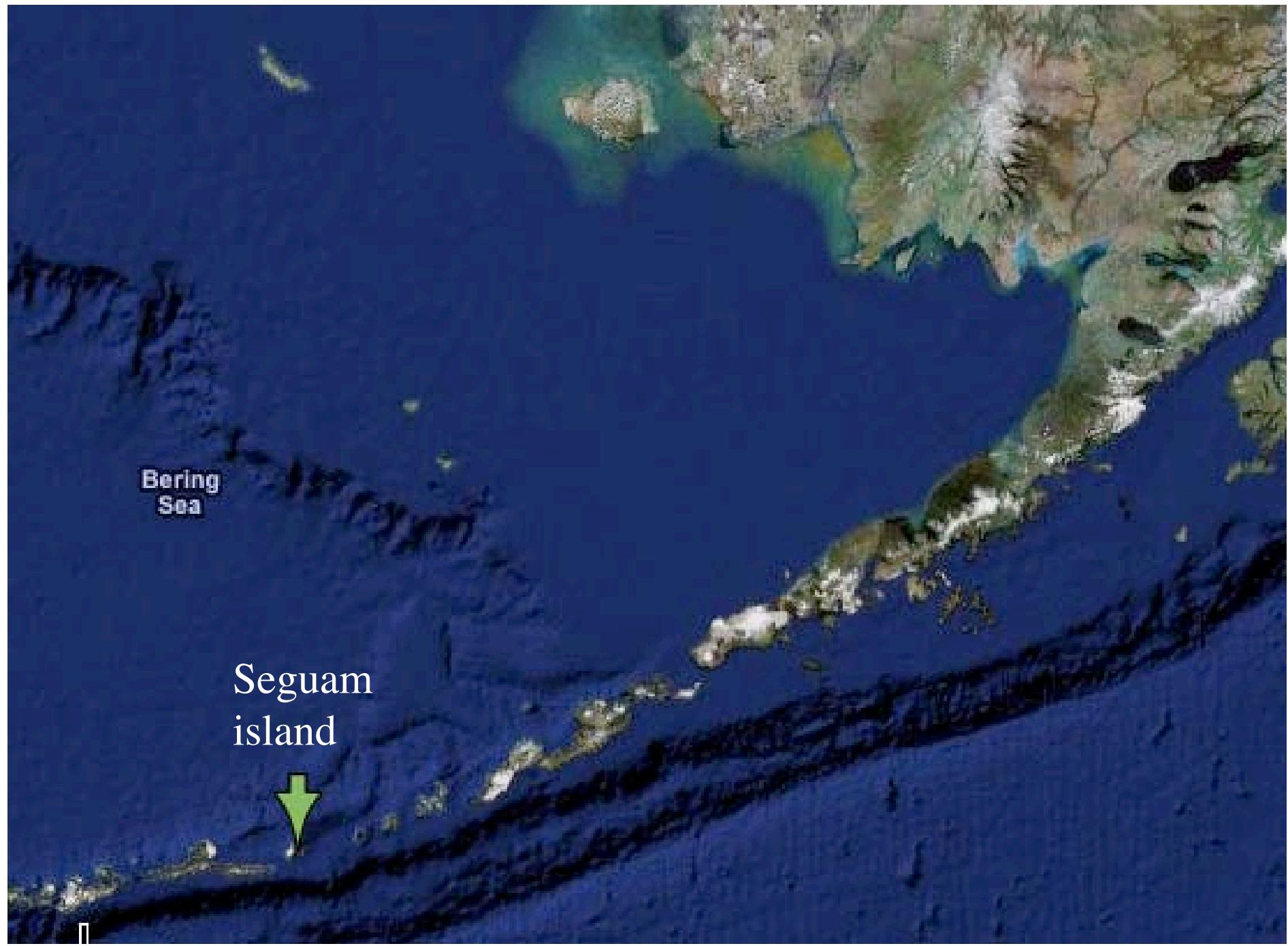
Office Hours: before class, faculty or student lounge

Questionnaire

Volcano field work in Alaska

My Research Interests:
Volcanism
Igneous Geochemistry
Subduction Zones





Bering
Sea

Seguam
island





1977 Eruption on Seguam Island, Alaska



US Coast Guard photo; Pyre Peak; 1054 m

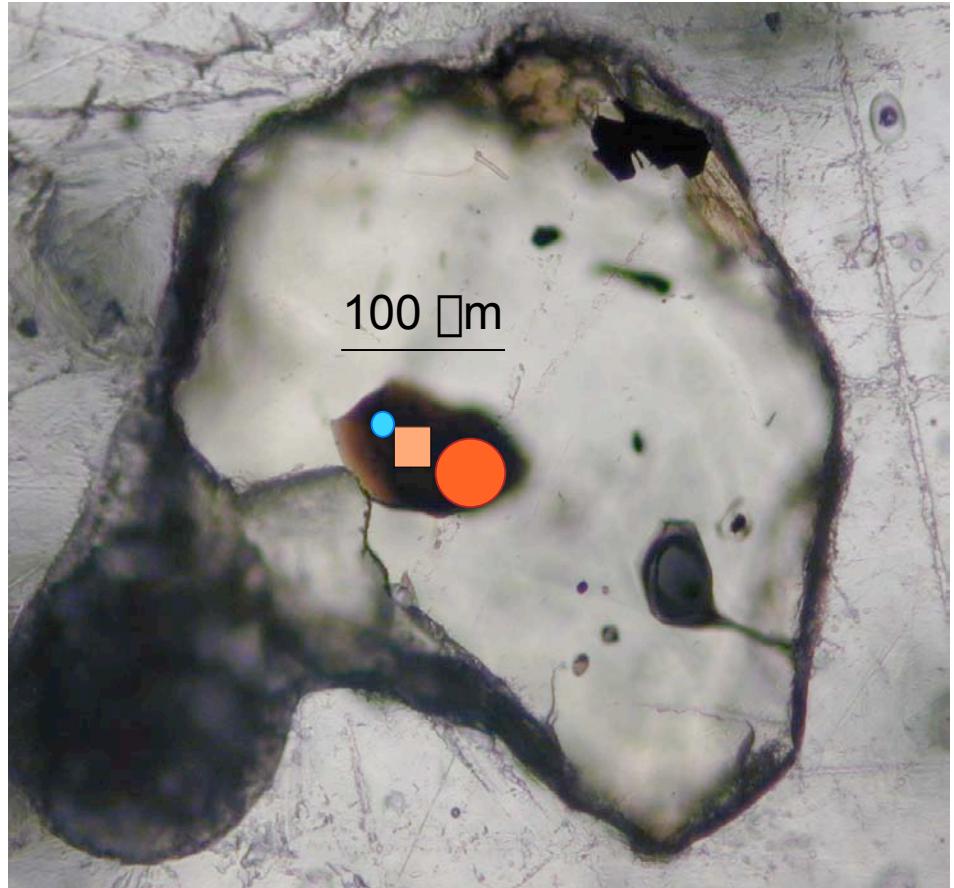






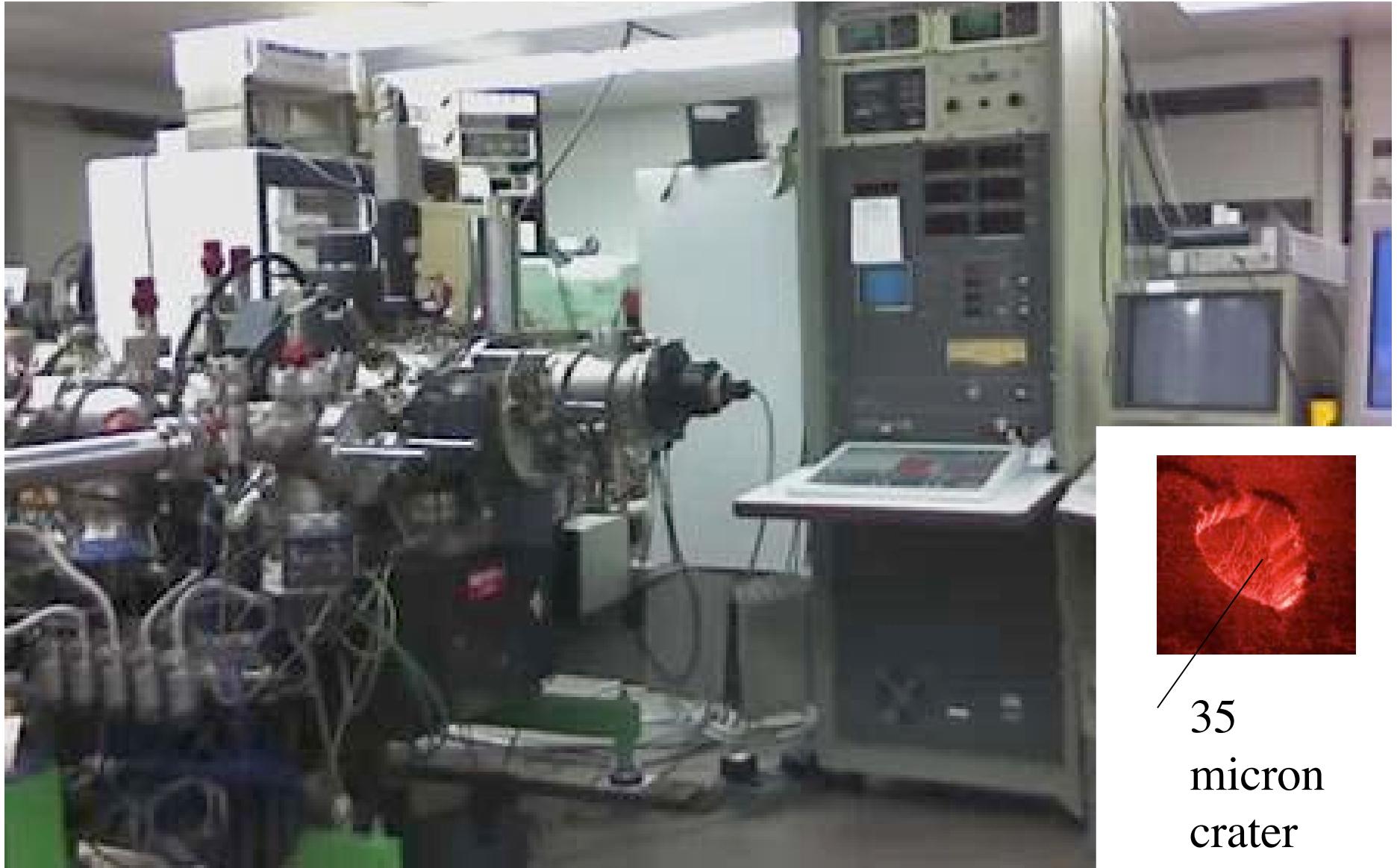


Olivine crystal
With melt inclusion

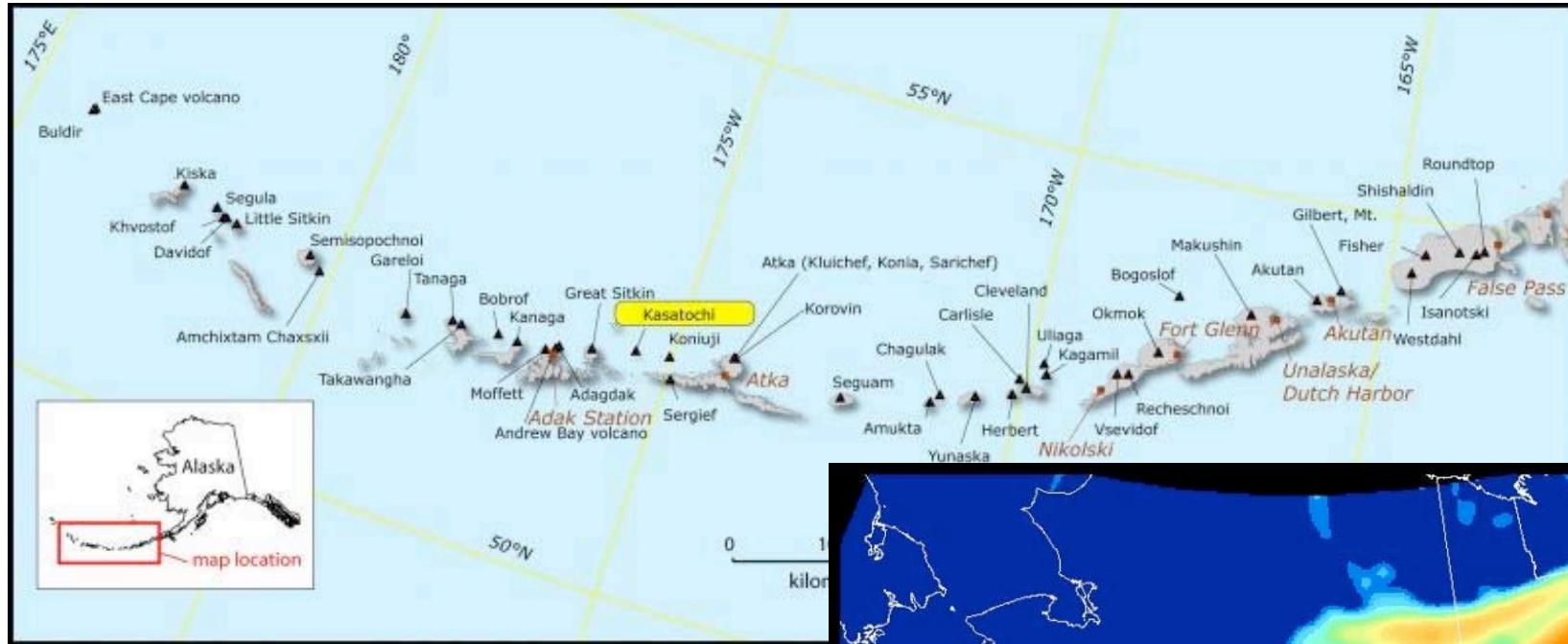


Spots analyzed with
ion microprobe, electron microprobe,
and laser ablation ICPMS

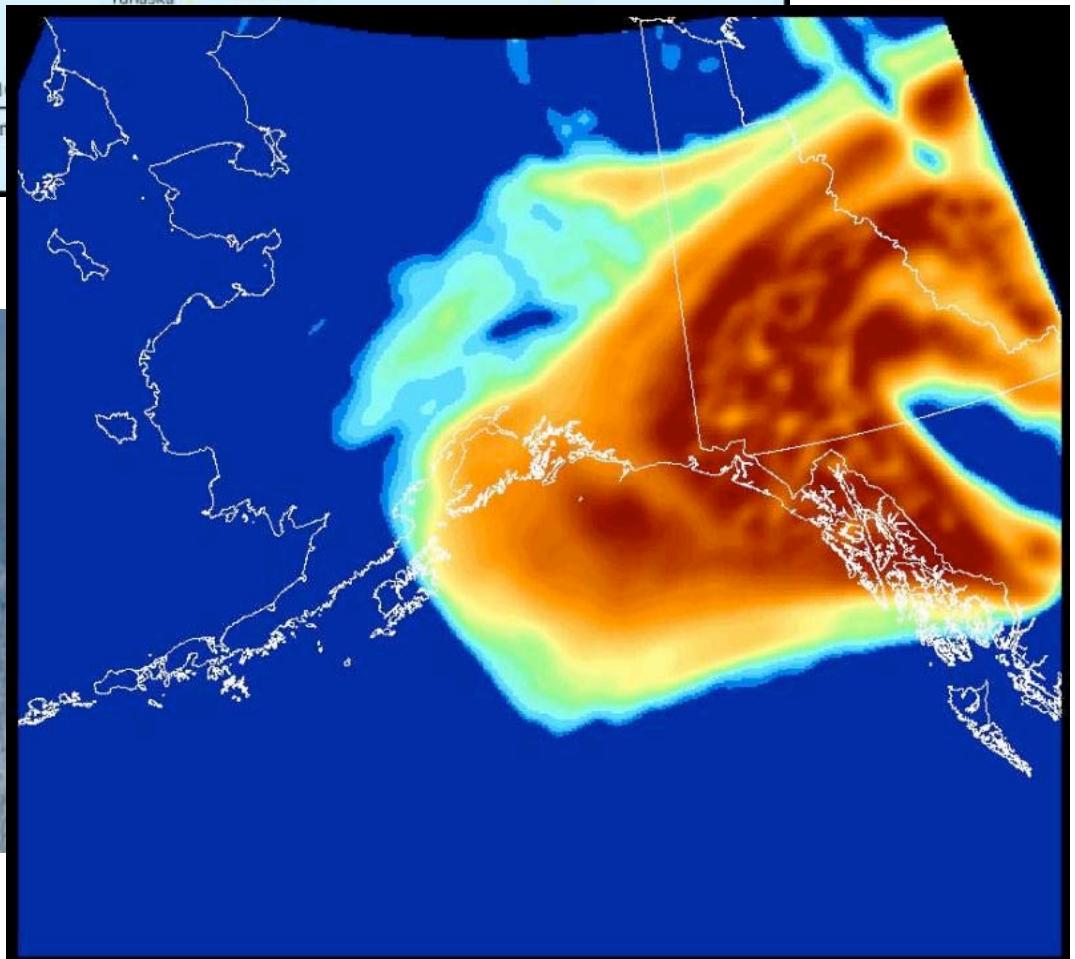
Ion Microprobe, Carnegie Institution of Washington



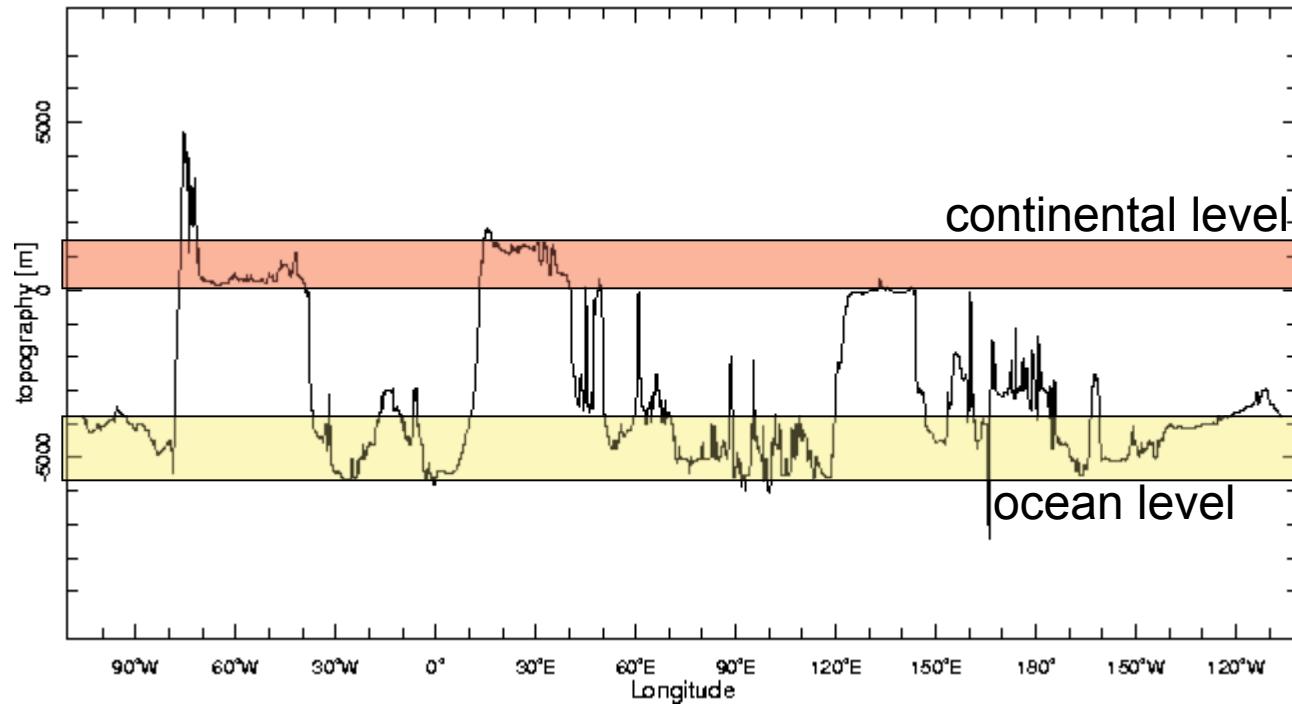
35
micron
crater



August 2008 eruption of Kasatochi



Earth Hypsometry

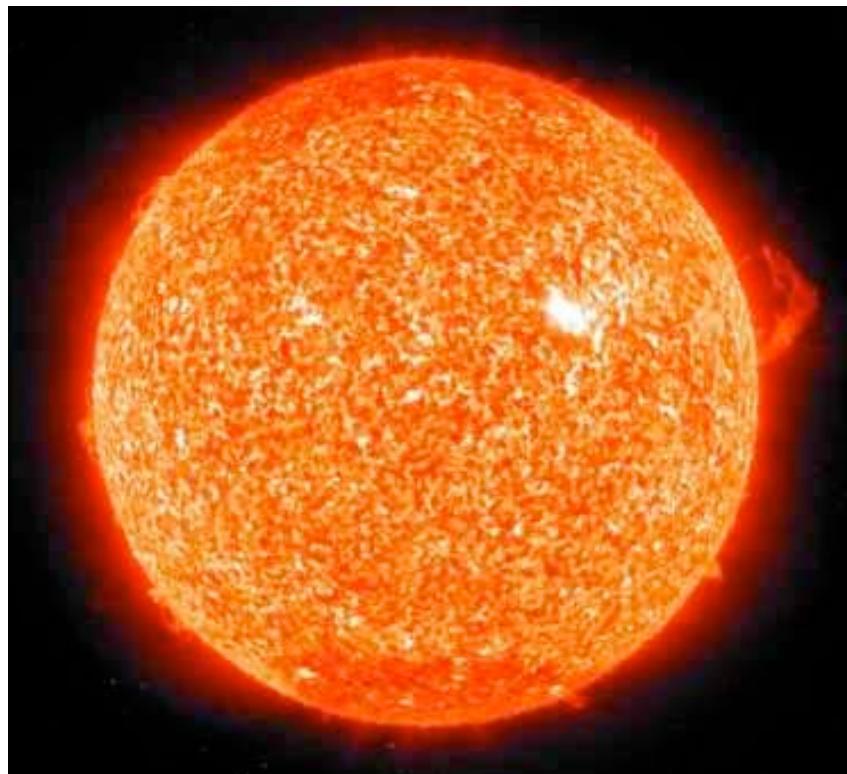


Simple observation.....complex cause



Composition and Structure of the Earth

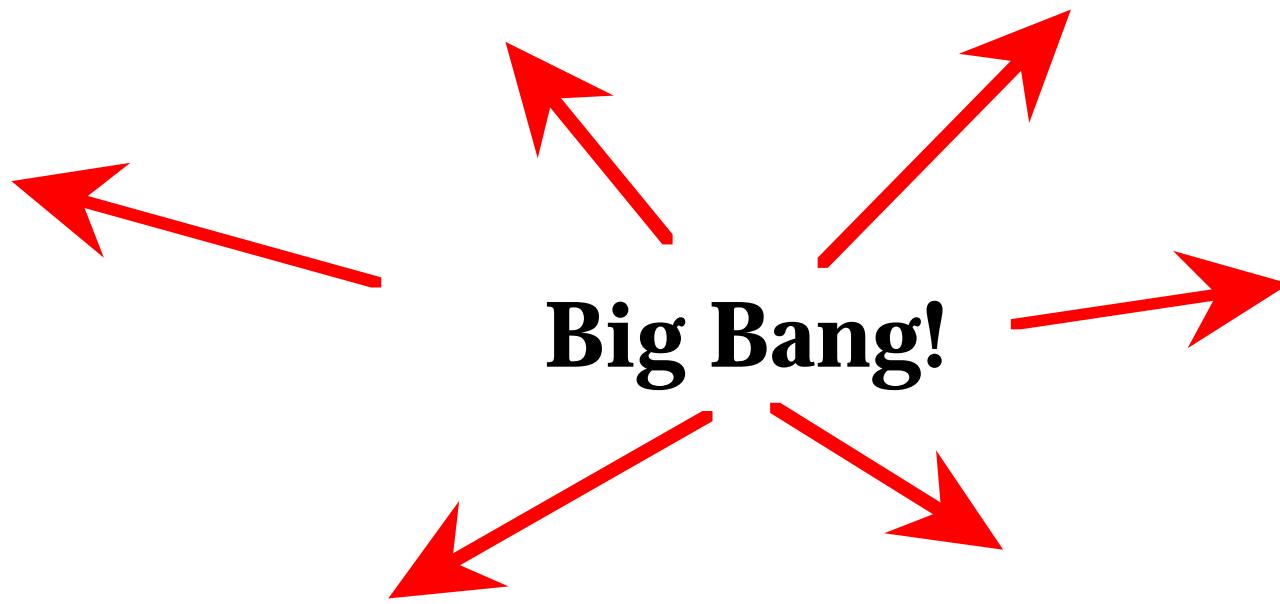
What is the Sun made of?



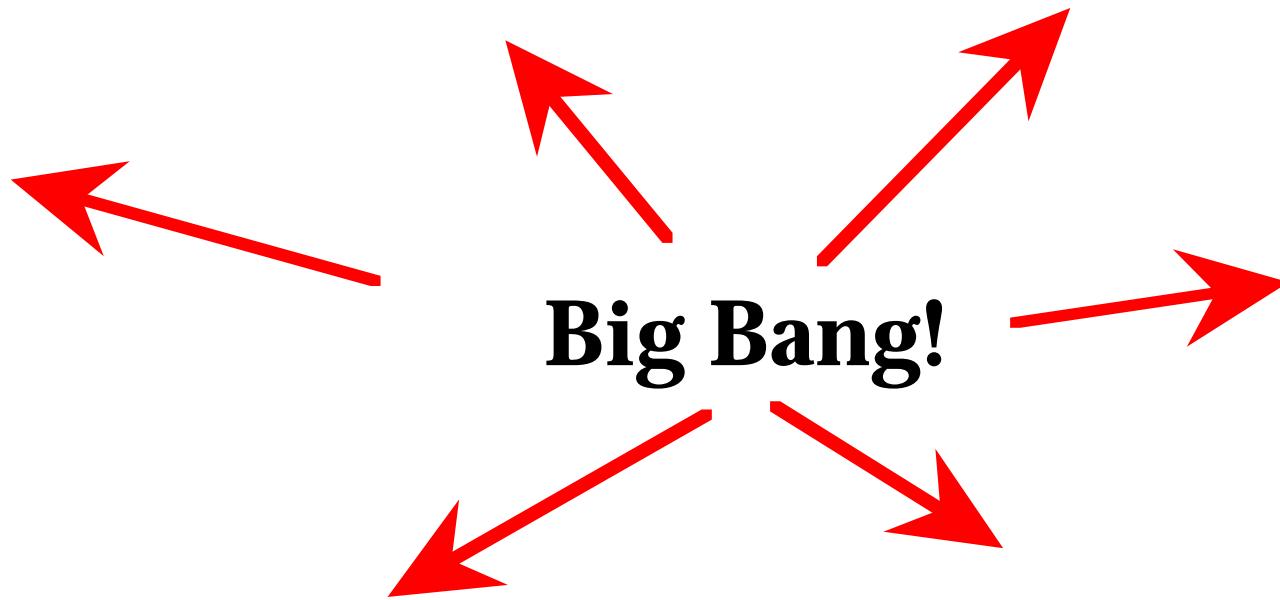
What is the Earth made of?



Top elements are.....



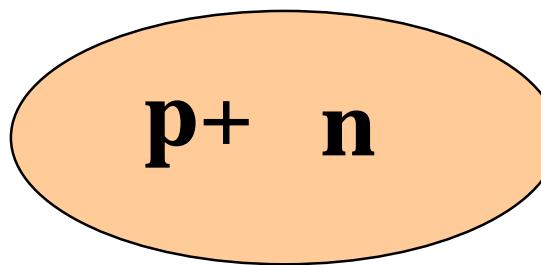
~ 14 Ga
All matter
expanding universe



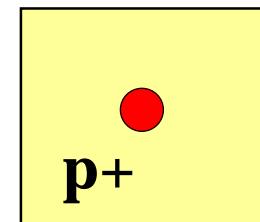
neutrons



half life = 12 min

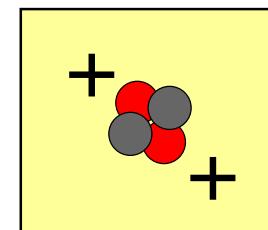


Raw



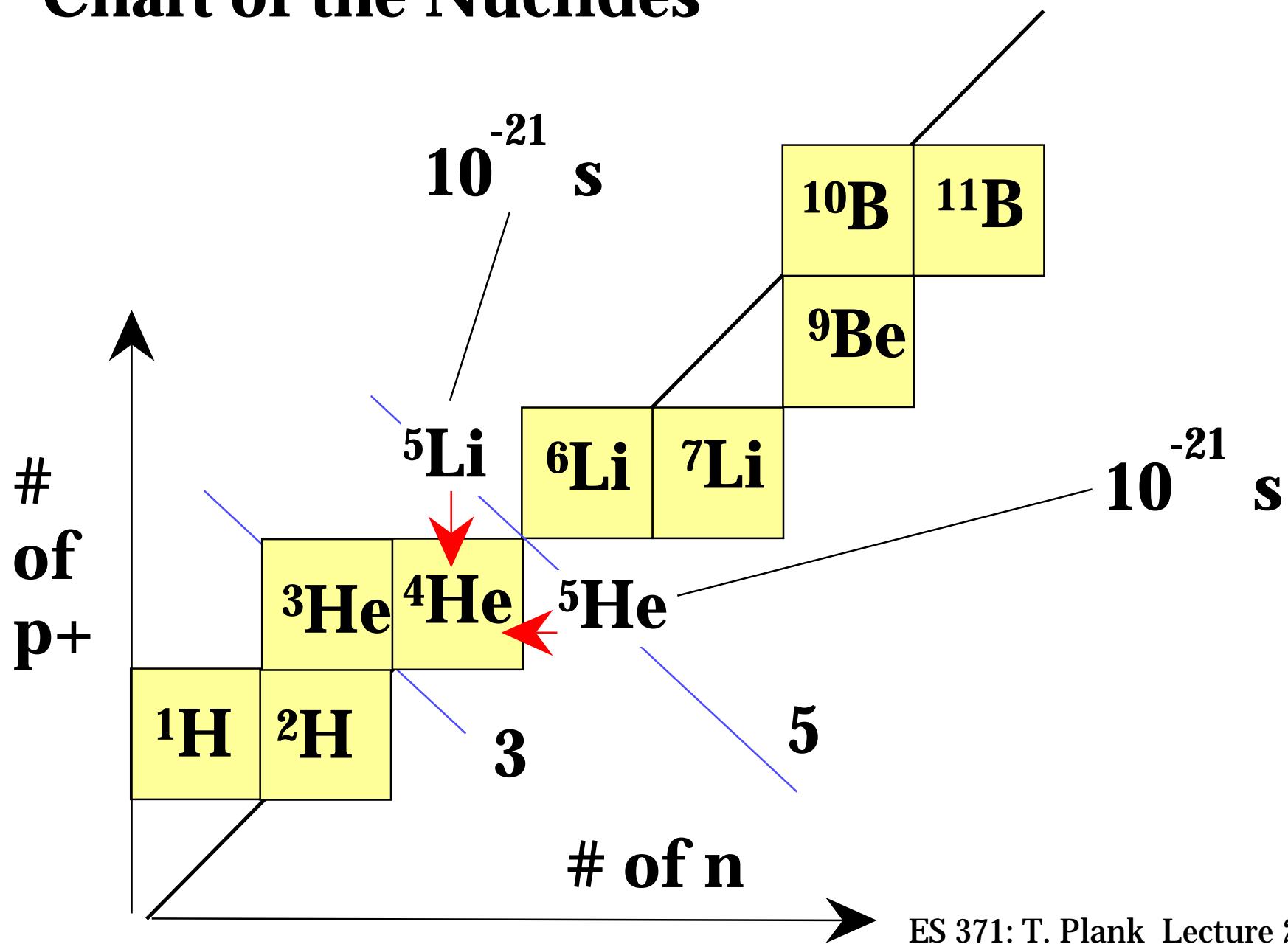
Hydrogen
nucleus

Helium
nucleus



why not keep going?

Chart of the Nuclides



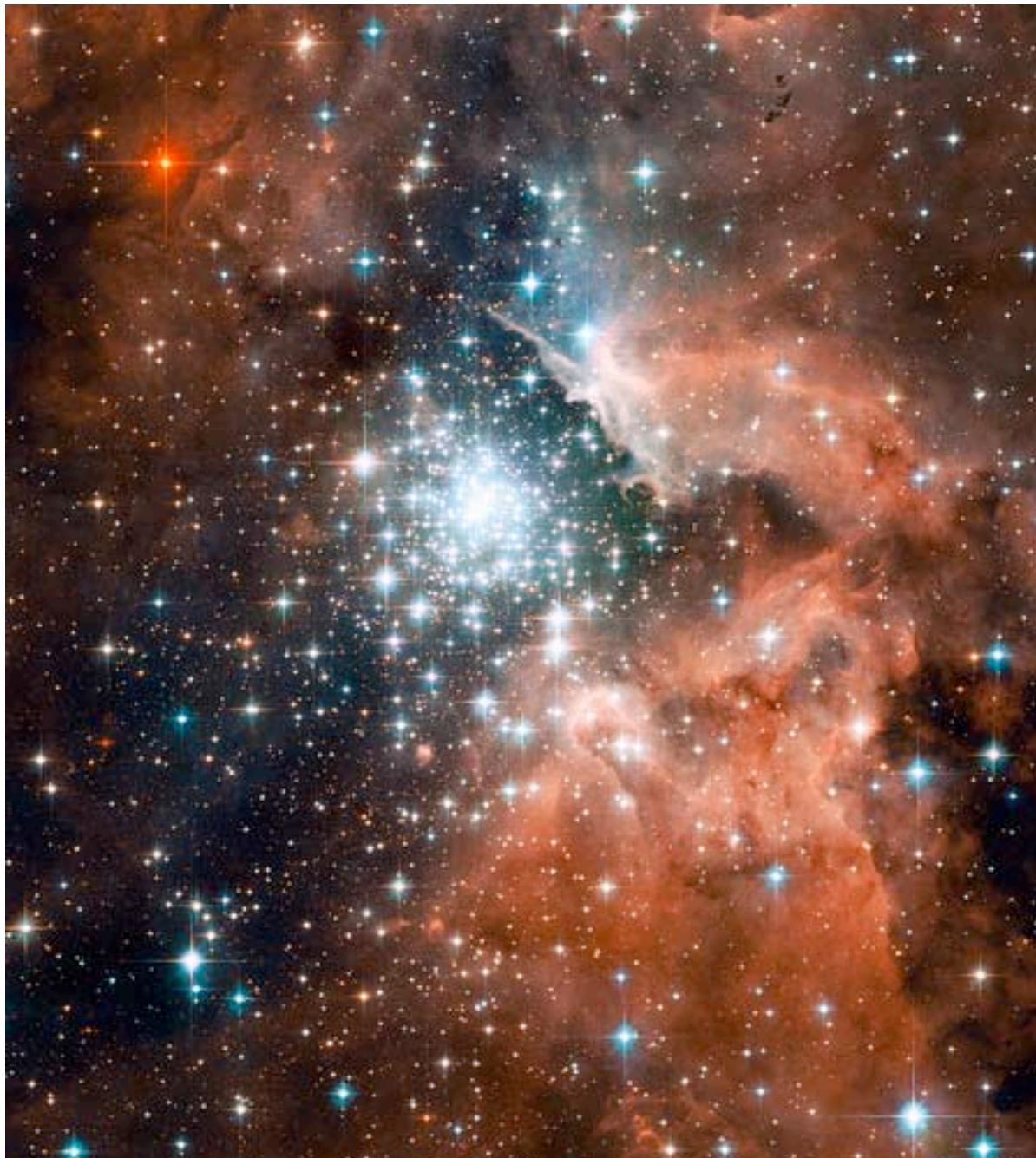
Day One

75% H nucleii

25% He nucleii

Interstellar Gas Clouds



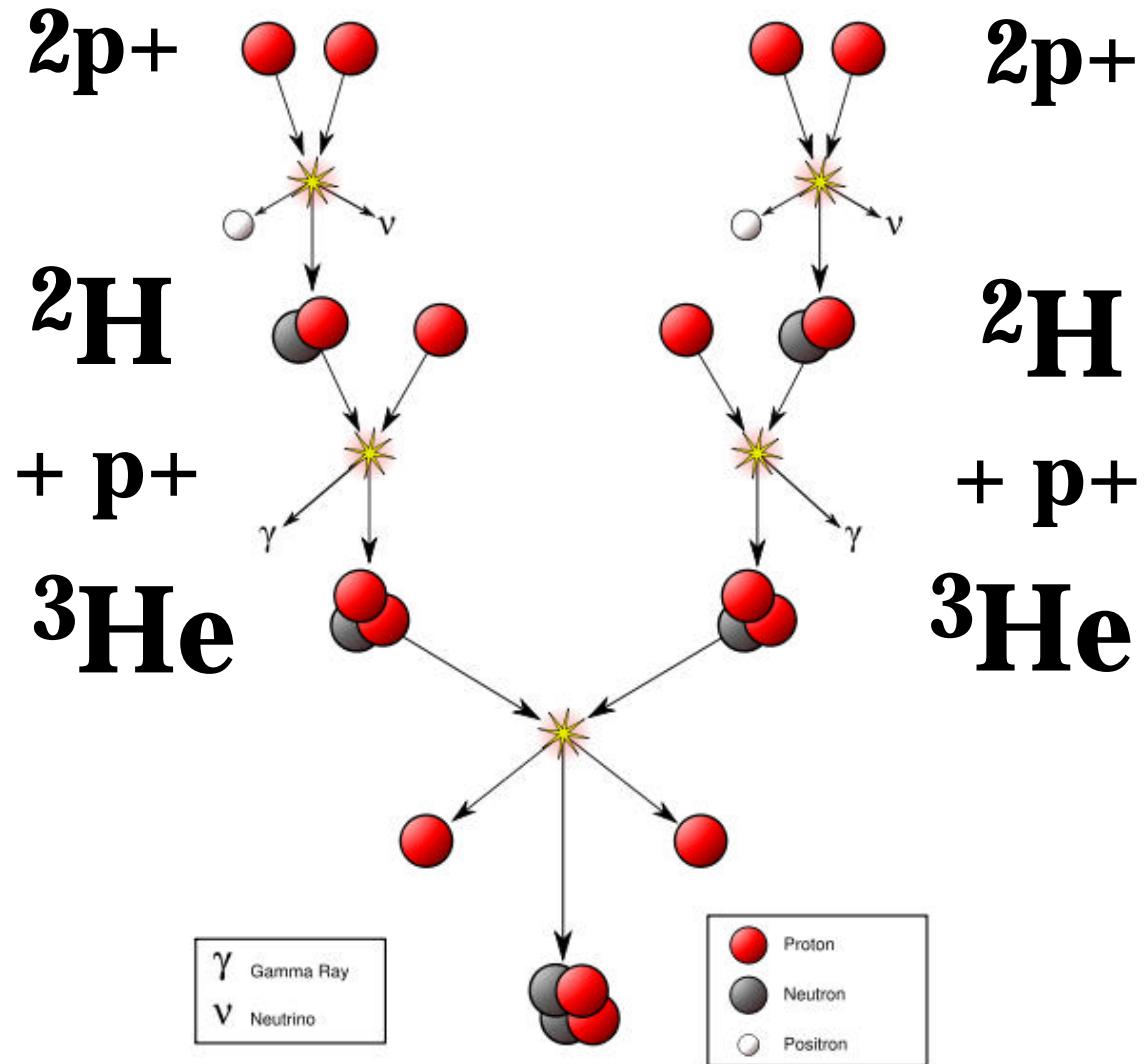


Gas clouds
coalesce
to form
first stars

Density
increases ->
 $T > 60 \text{ M K}$

Fusion!

H burning



net: $4\text{ p}+ \rightarrow ^4\text{He}$

ES 371: T. Plank Lecture 2

H burning

"Nuclear fire"

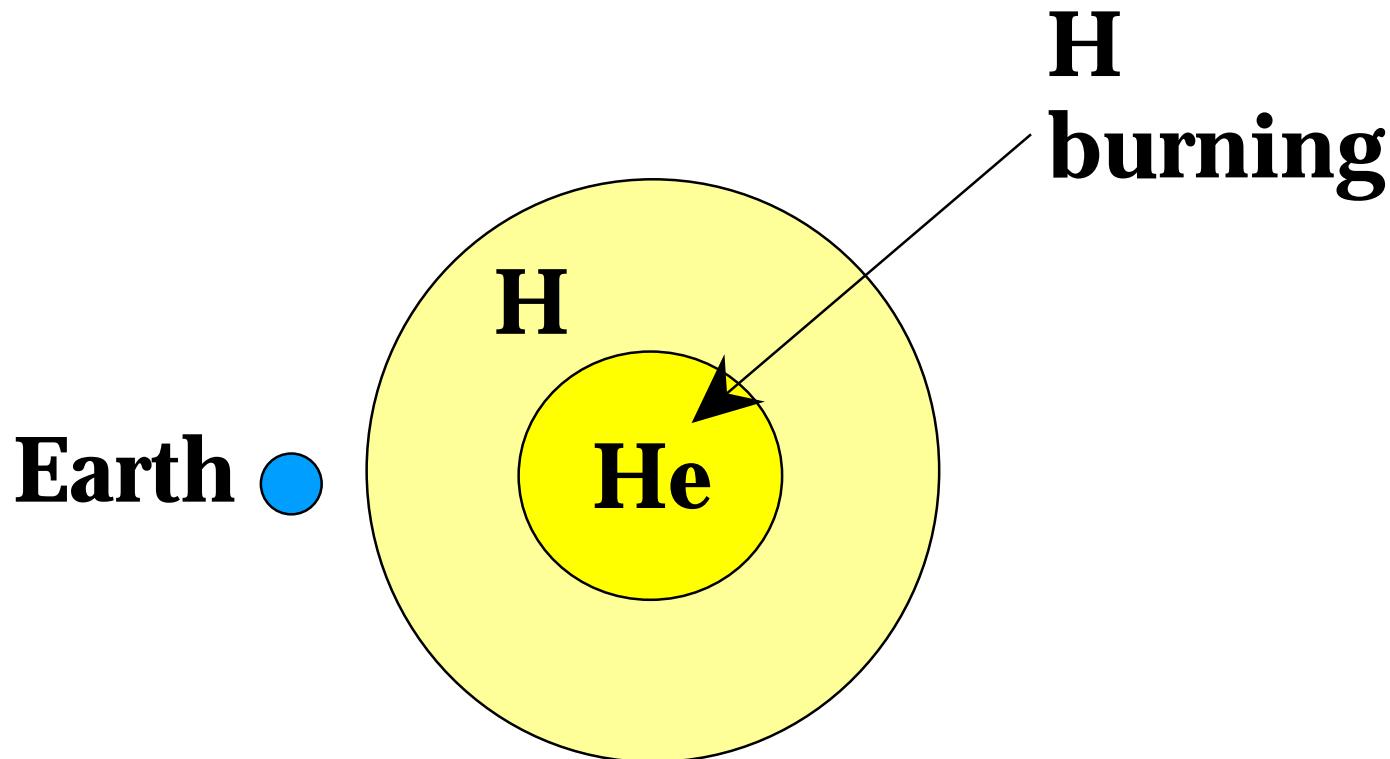
net: $4 \text{ p}+ \rightarrow {}^4\text{He}$

$$4 \times {}^1\text{H} = 6.696 \times 10^{-24} \text{ g}$$

$$1 \times {}^4\text{He} = 6.648 \times 10^{-24} \text{ g}$$

$$\boxed{\mathbf{E = mc^2 !!}} \quad + 26.7 \text{ MeV}$$

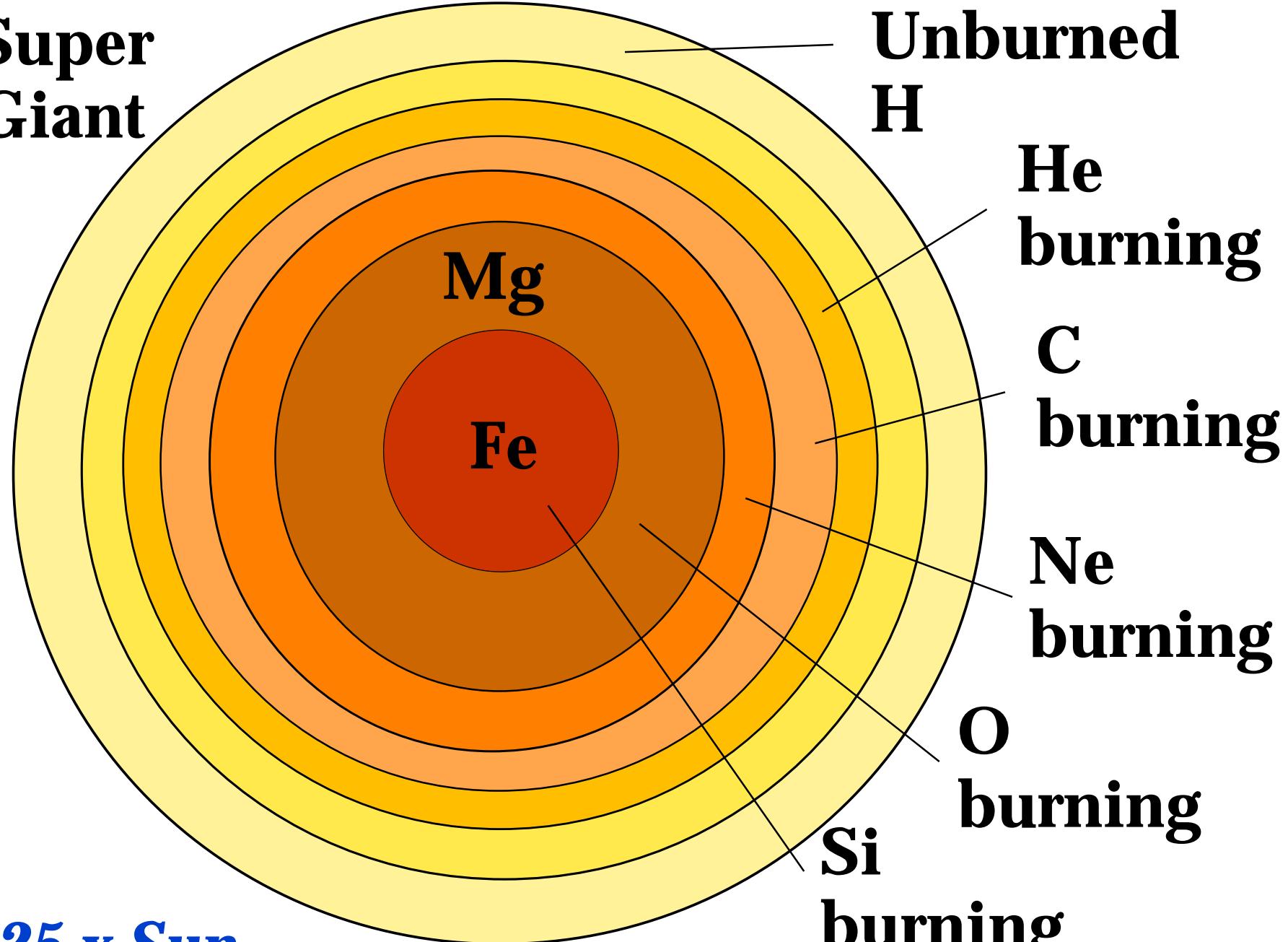
our Sun



**H
burning**

but what happens eventually?

**Super
Giant**

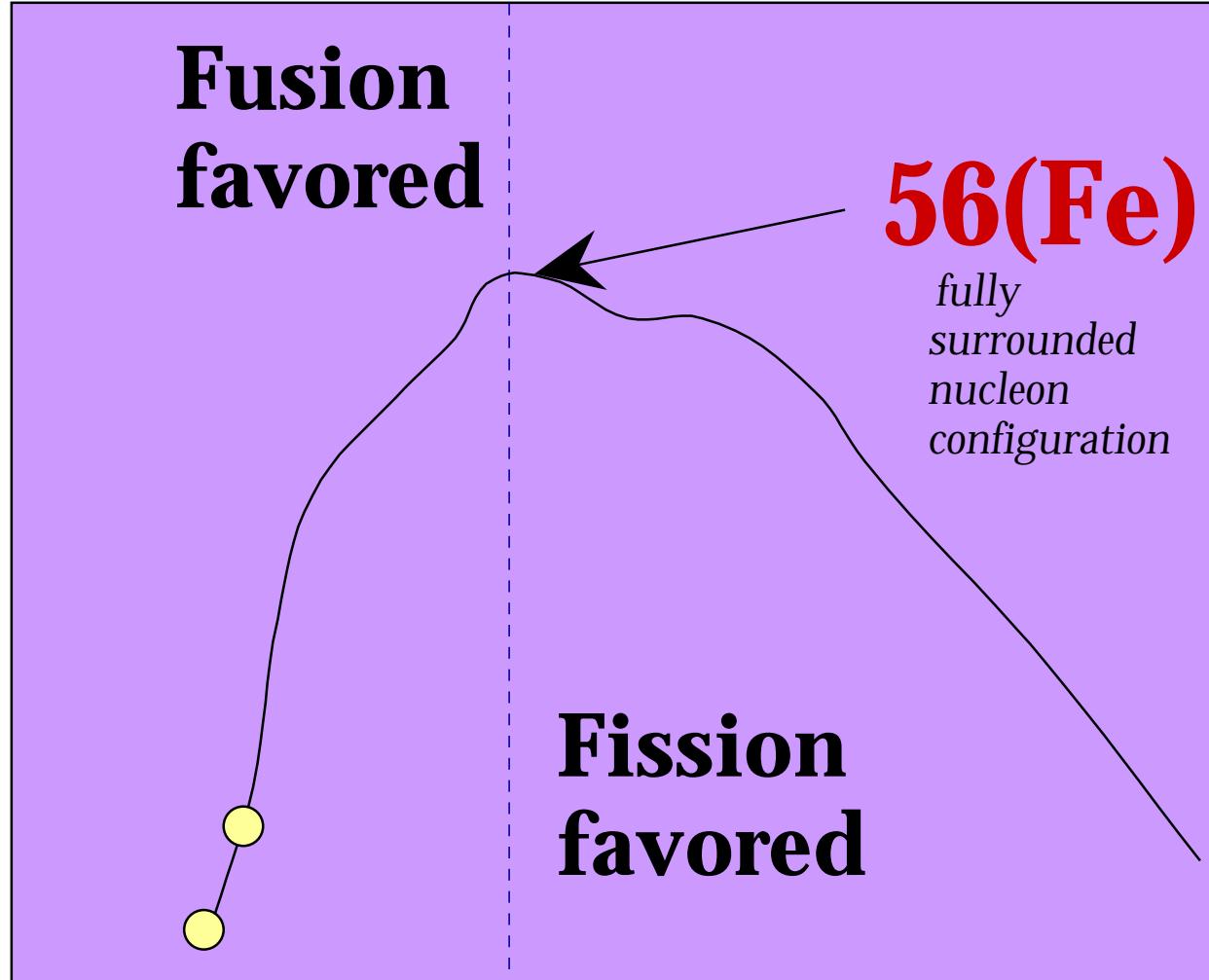


25 x Sun

ES 371: T. Plank Lecture 2

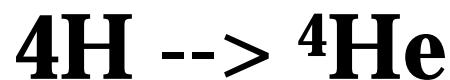
Binding Energy per Nucleon

E needed
to pull
a nucleon
from the
nucleus



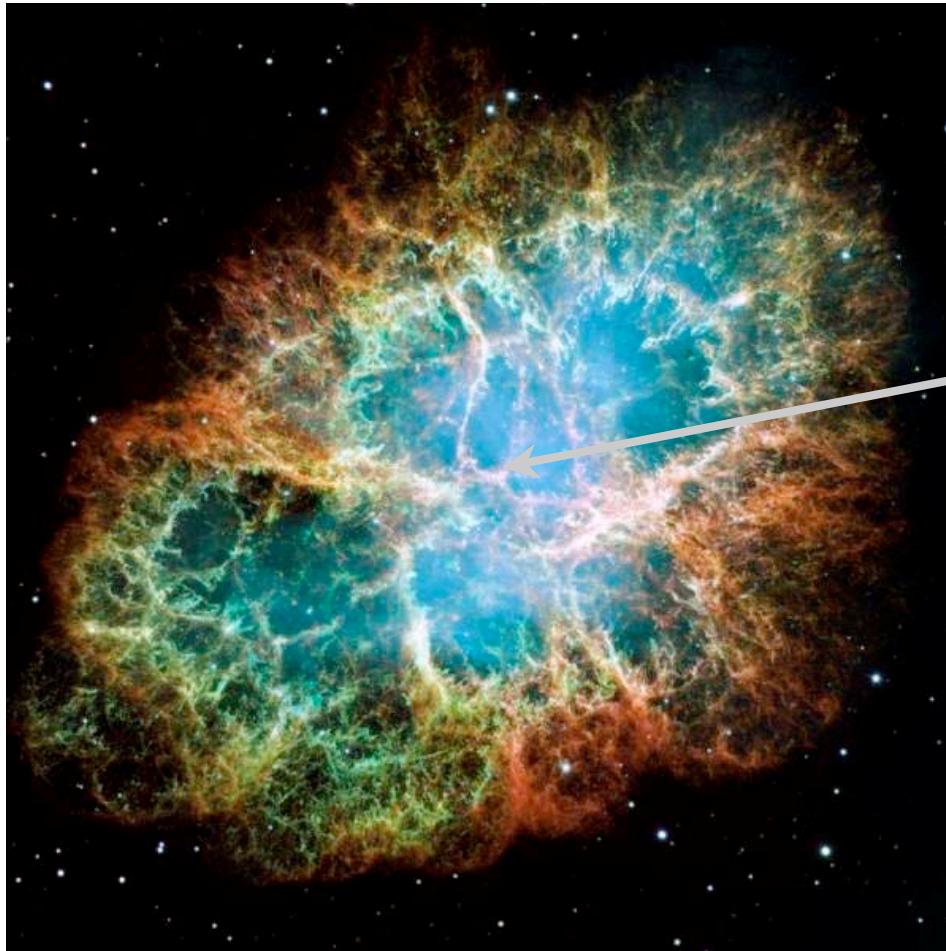
$$A \text{ (p+n)}$$

Strong nuclear force
vs.
electrostatic force



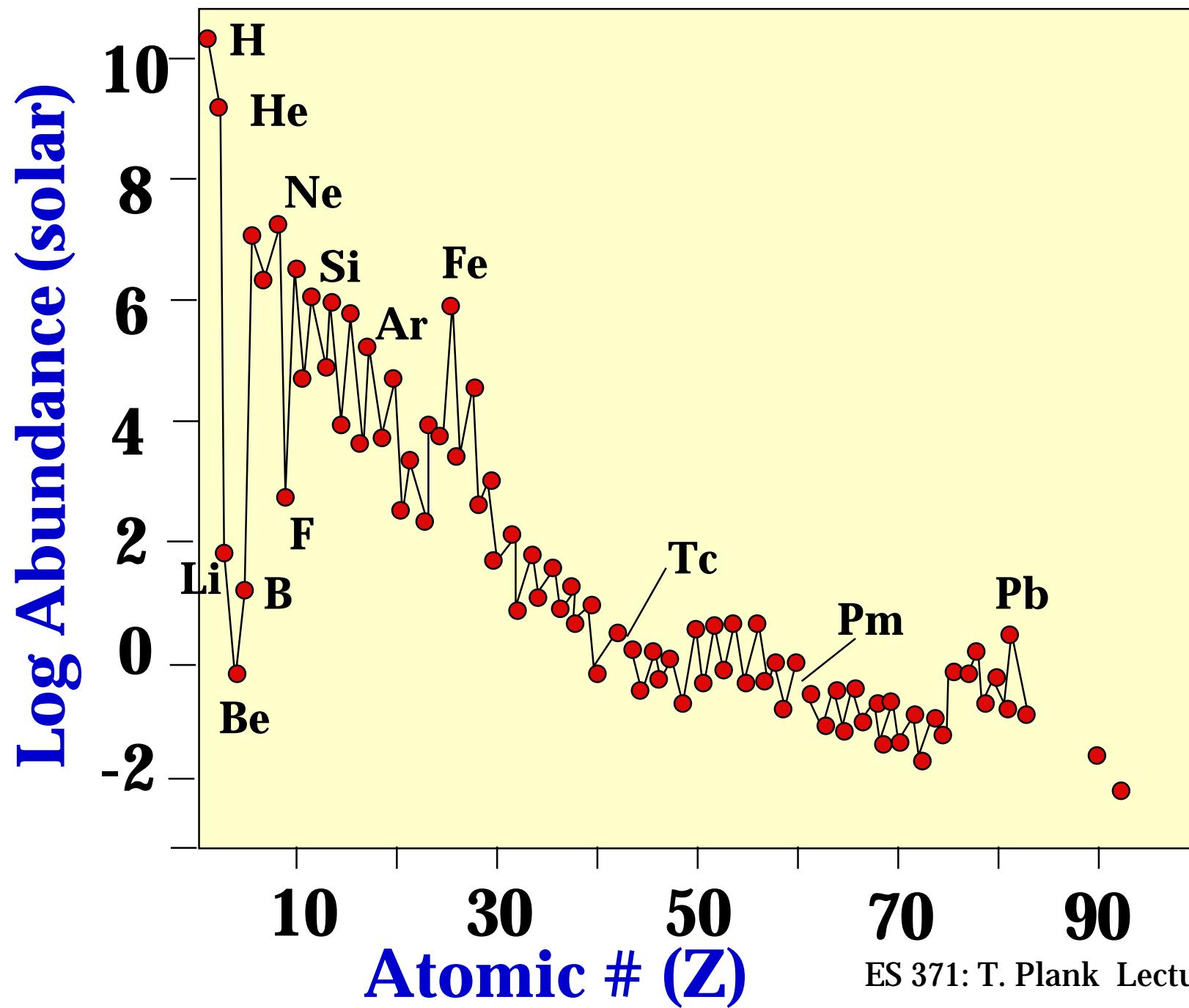
Supernova!

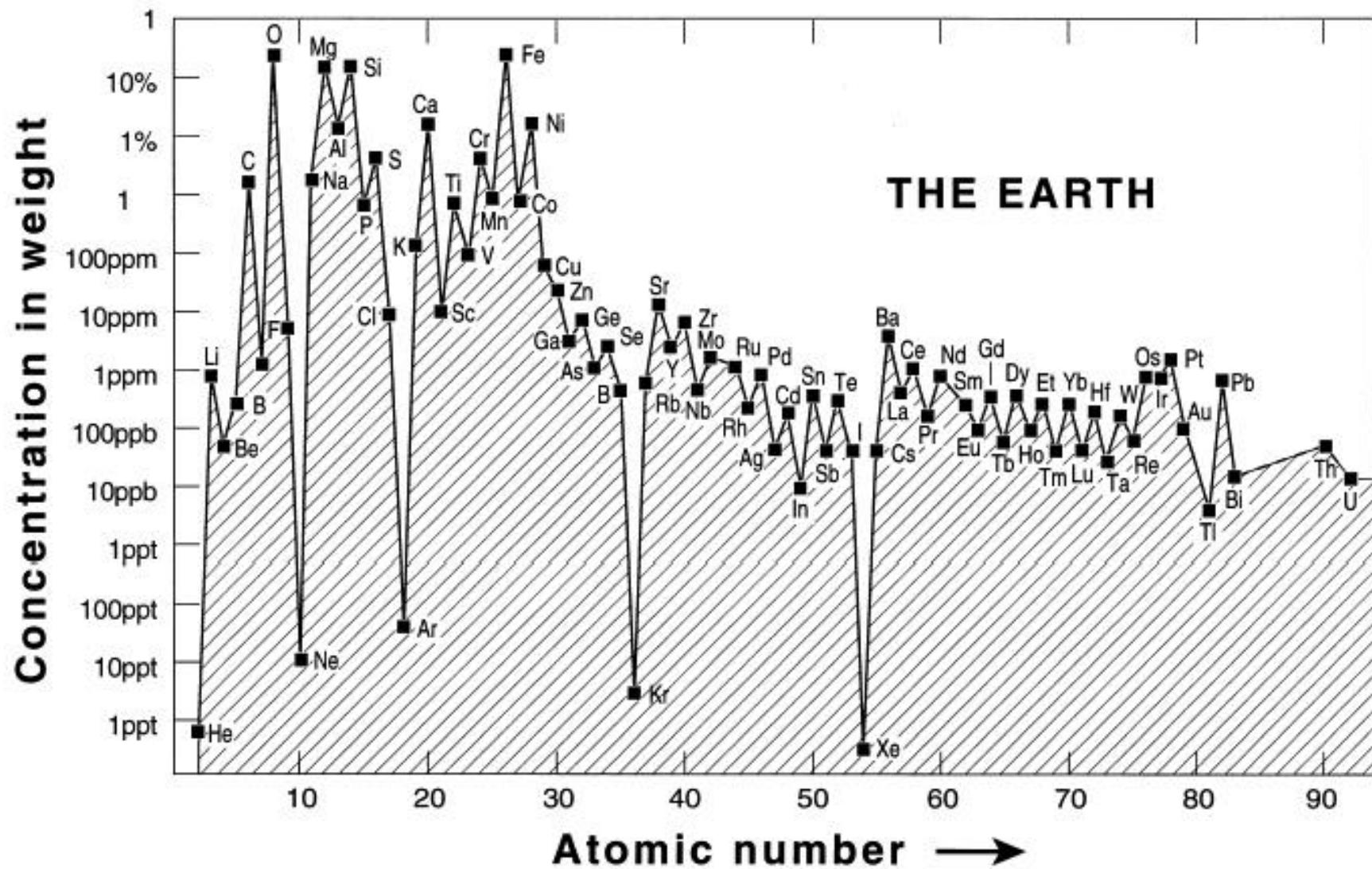
- *Heavy element factory*
- *Element dispersal system*



Original star
is just tiny
dot center

*Crab Nebula: Supernova Remnant 6,500 light years away
seen by Chinese and Japanese astronomers (with naked eye!) in 1054*





Allegre, 2001

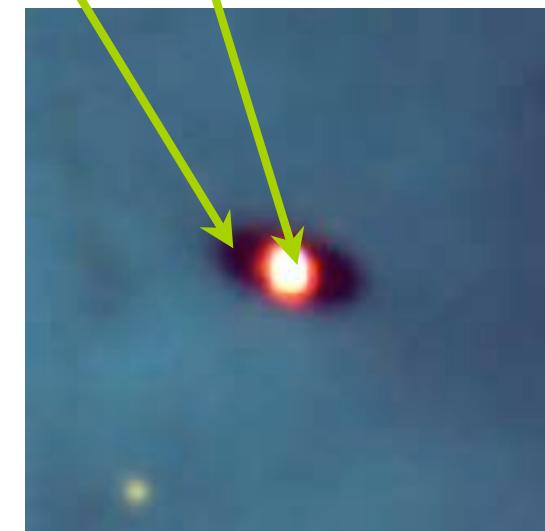
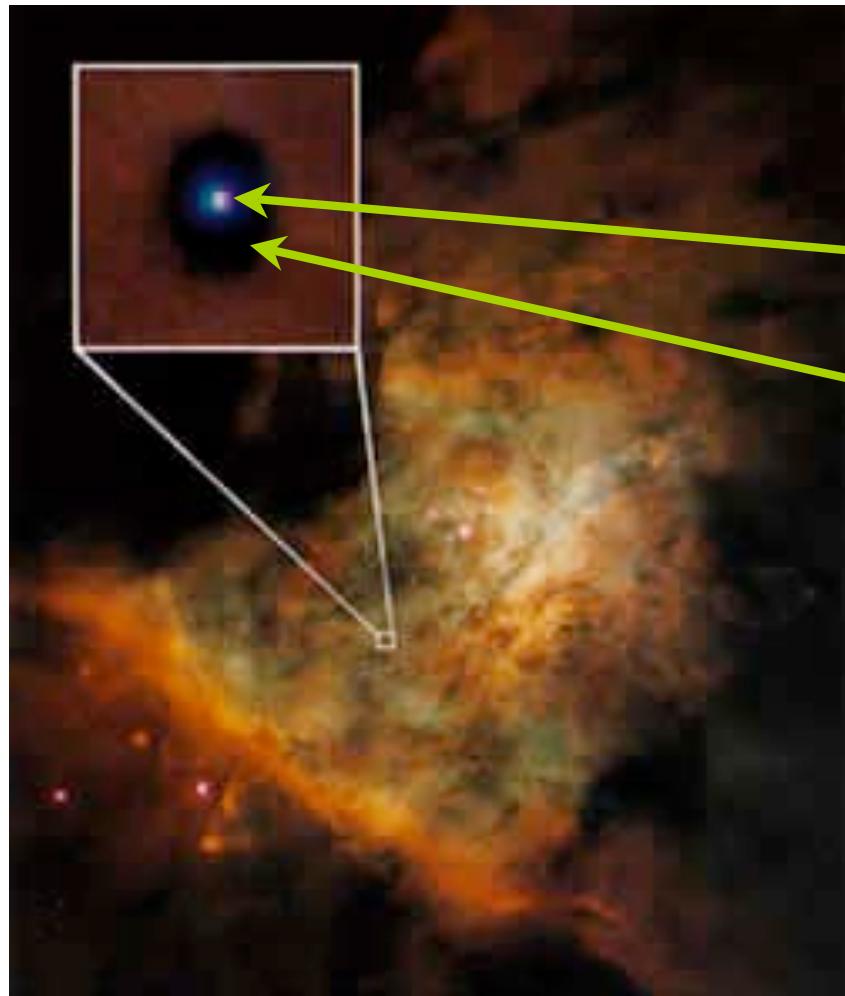
ES 371: T. Plank Lecture 4

Birth of a solar system

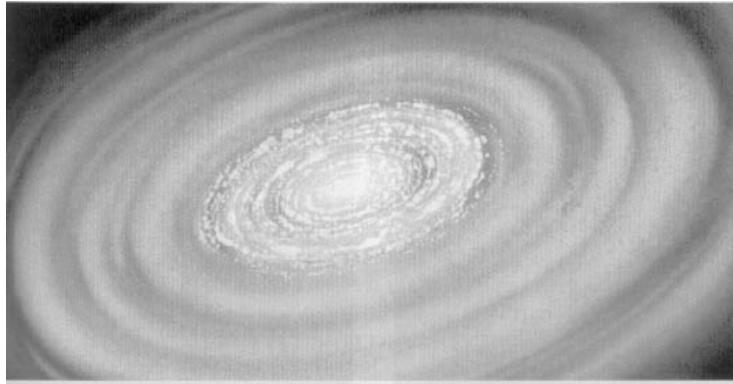


Hubble Space Telescope view of a portion of the Orion Nebula showing five young stars surrounded by gas and dust trapped as the stars formed, that might evolve to planets.

Accretion disk and star

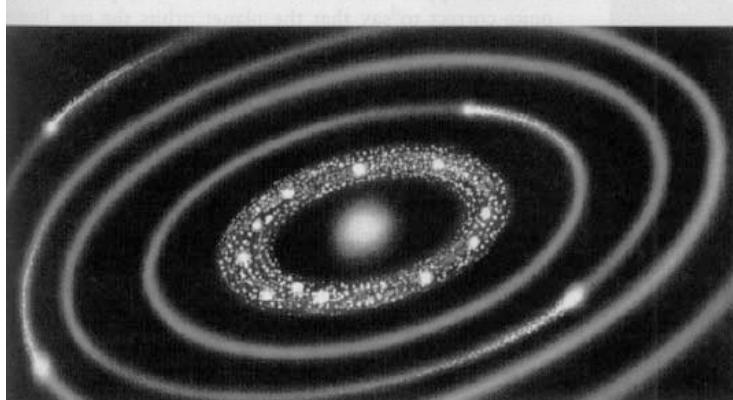


disk
star

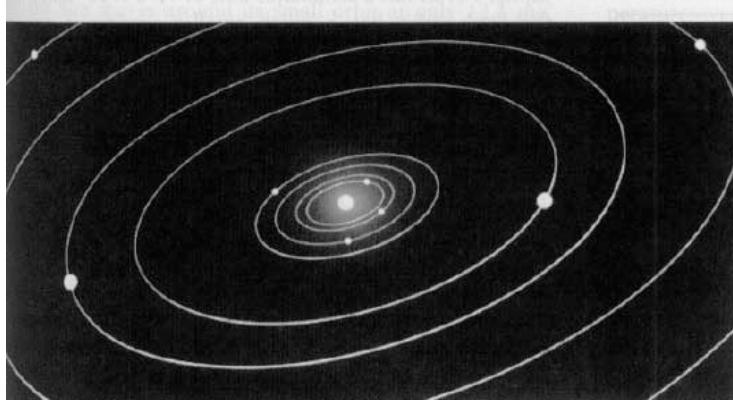


4.56 Ga

Evolution of the accretion disk
into the sun and planets



Explains why all planets orbit
in the same plane and in the
same direction



solar wind blows light
elements away from inner
solar system, explains their
low abundance in inner solar
system

time



Cold Dust

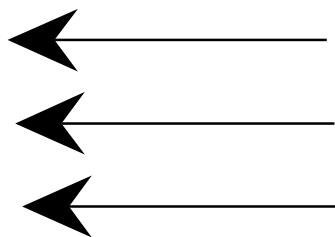
initial

Hot Gas

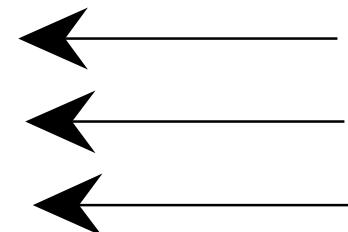
grav.
heating

Partially
Condensed

cooling



solar wind
leaves condensates



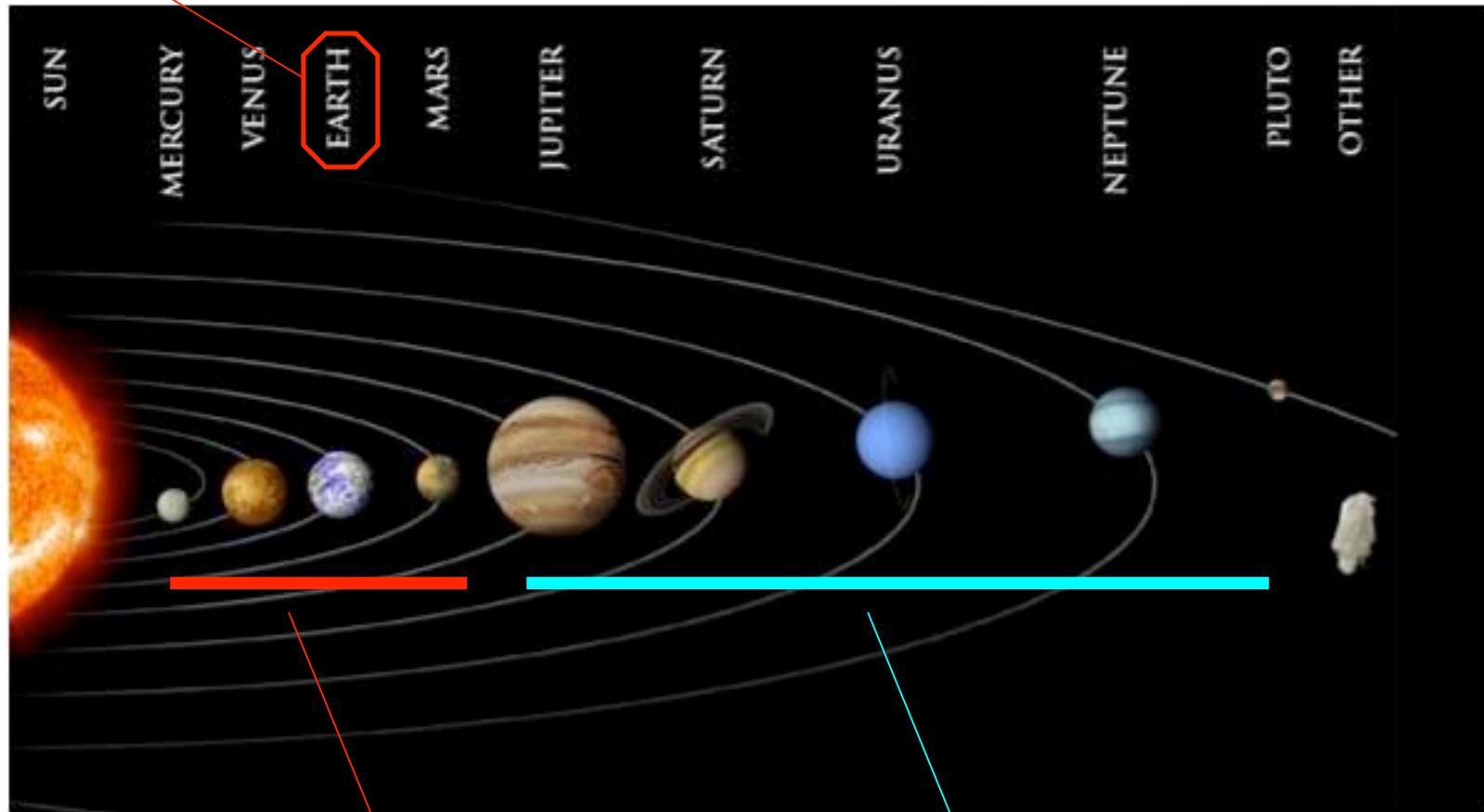
Composition of the Terrestrial Planets

85% = ^{16}O , ^{24}Mg , ^{28}Si , ^{56}Fe

- Primary solar abundances
- Solid vs. Gas

home sweet home

our solar system



Terrestrial planets

Mostly heavy nuclei

Outer solar system

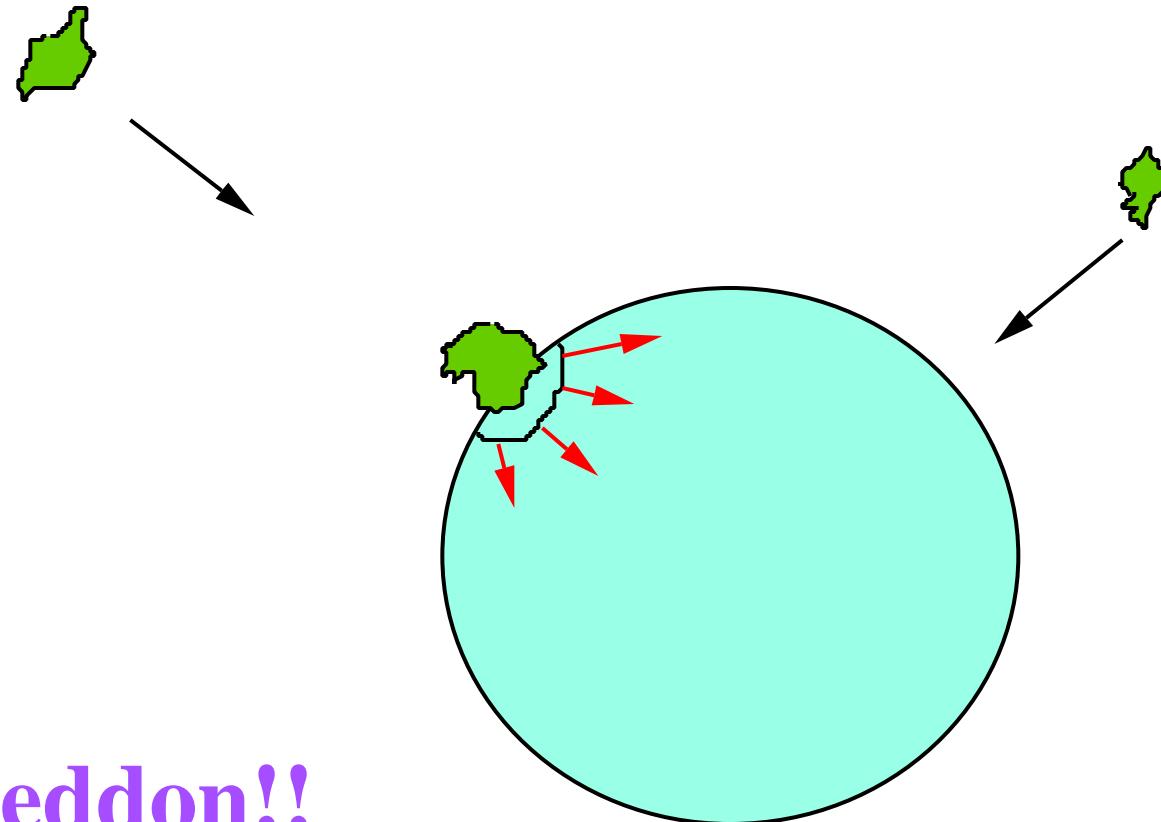
Lots of H and He

Eris (Xena): New Dwarf Planet '06



Accretion = sticking together

1. *impact heating*



Armageddon!!

ES 101 - Lect 1

One Giant Impact Formed the Moon!

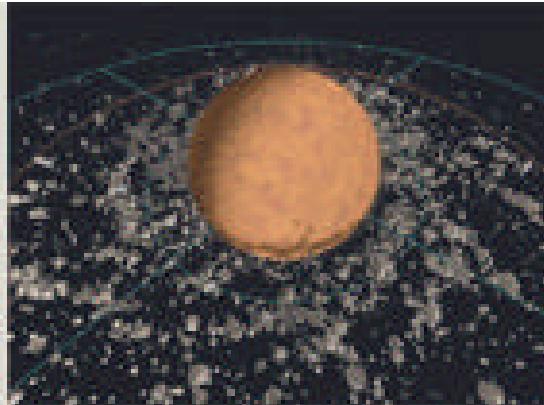


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Giant Impact Theory



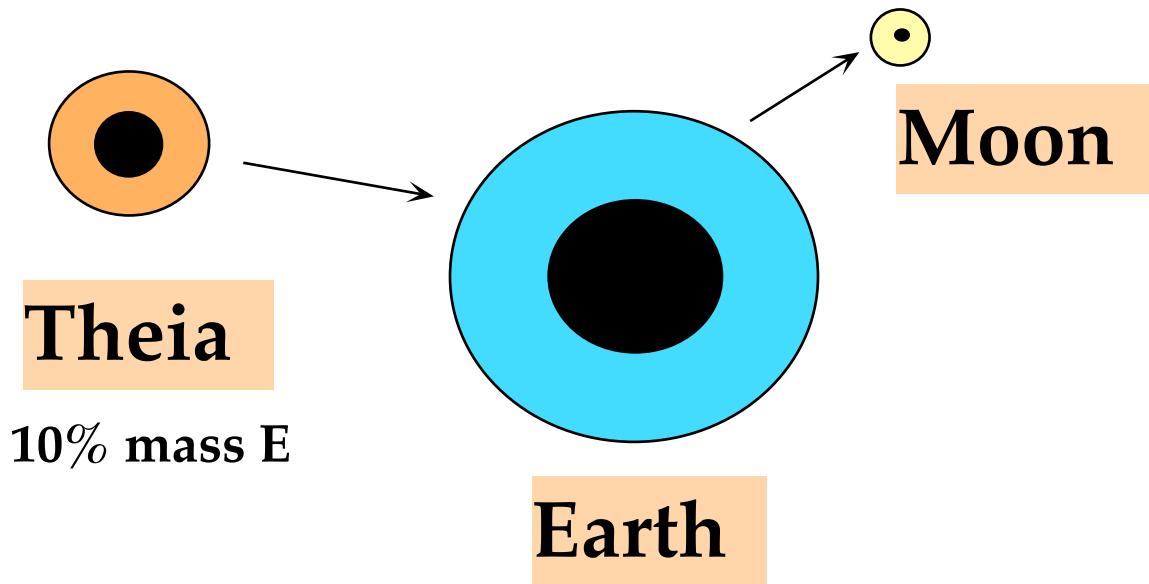
debris disk
from
collision
of
Theia



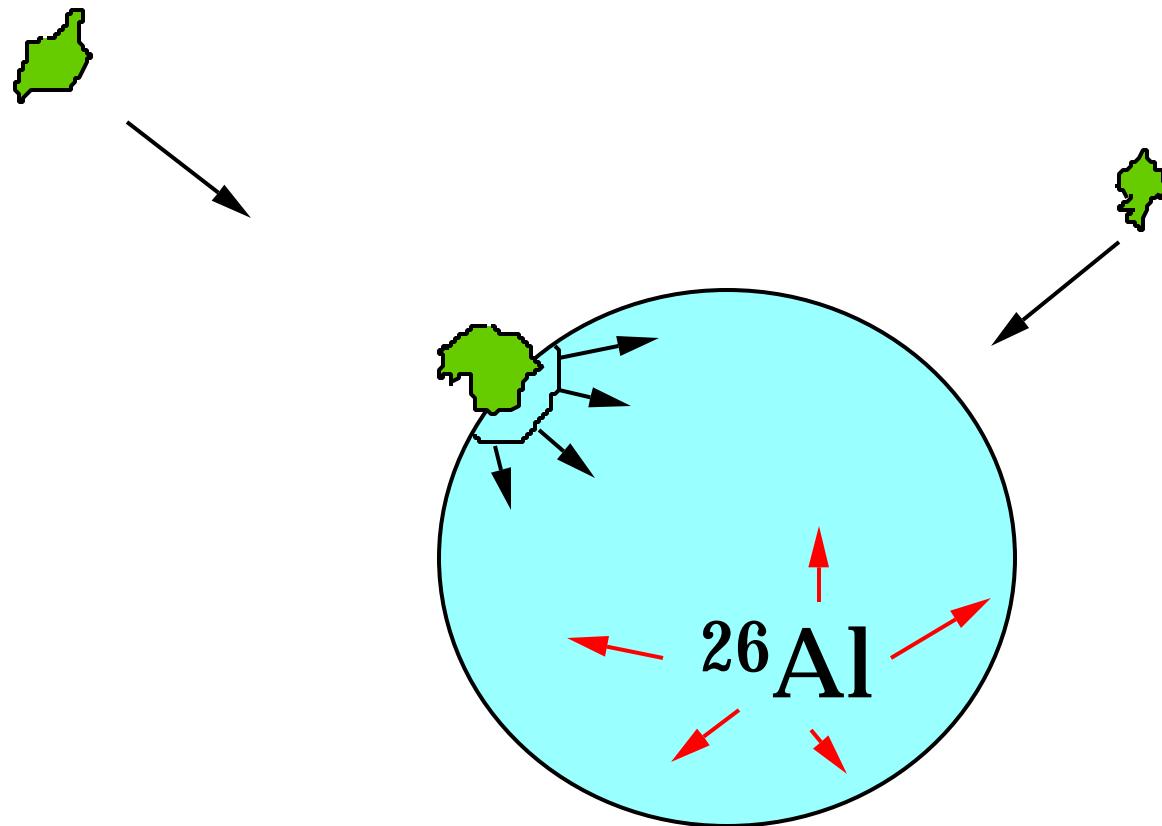
20 hrs later -
spiral arm
structure



200 hrs later -
a single large
moon



1. impact heating
2. Radioactive heating

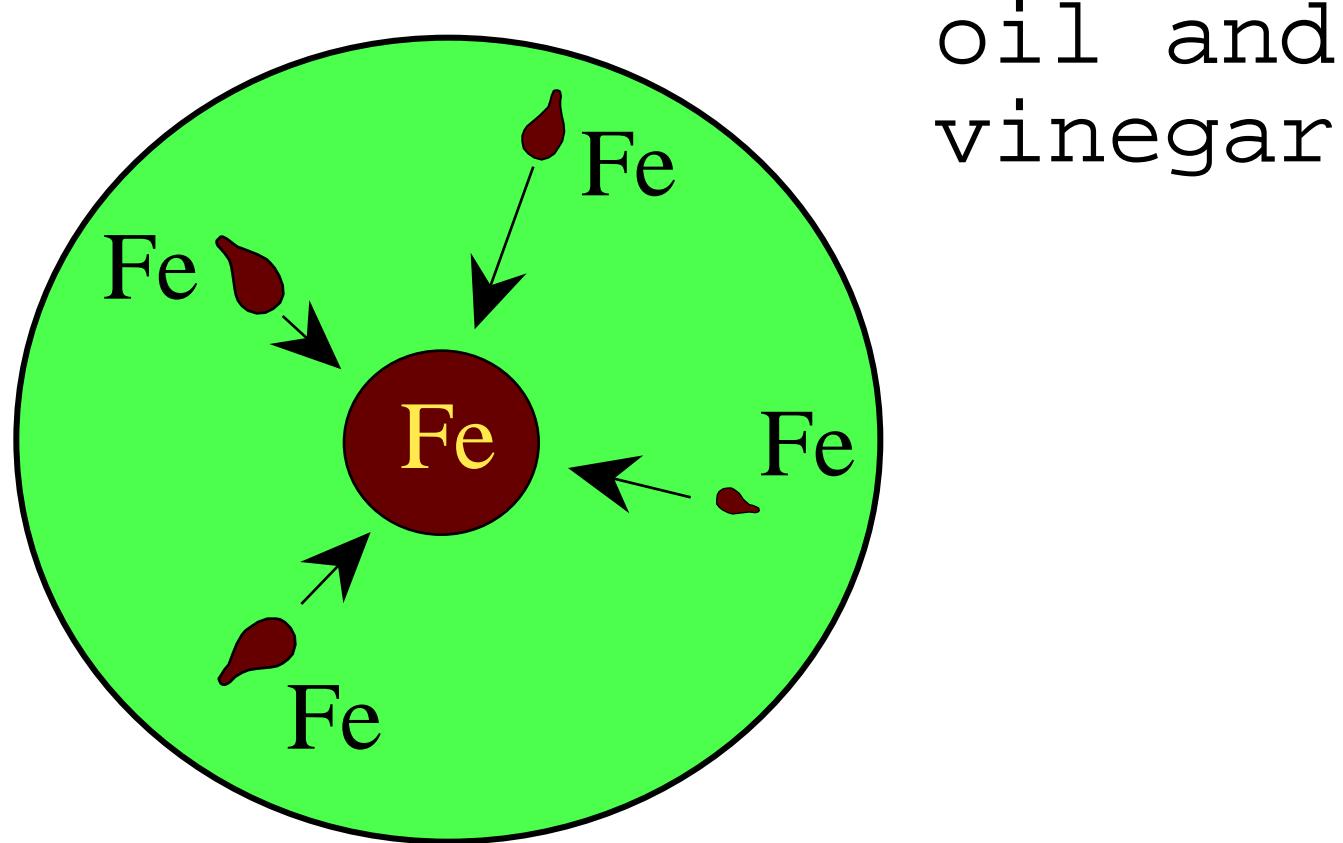


earth started with magma ocean



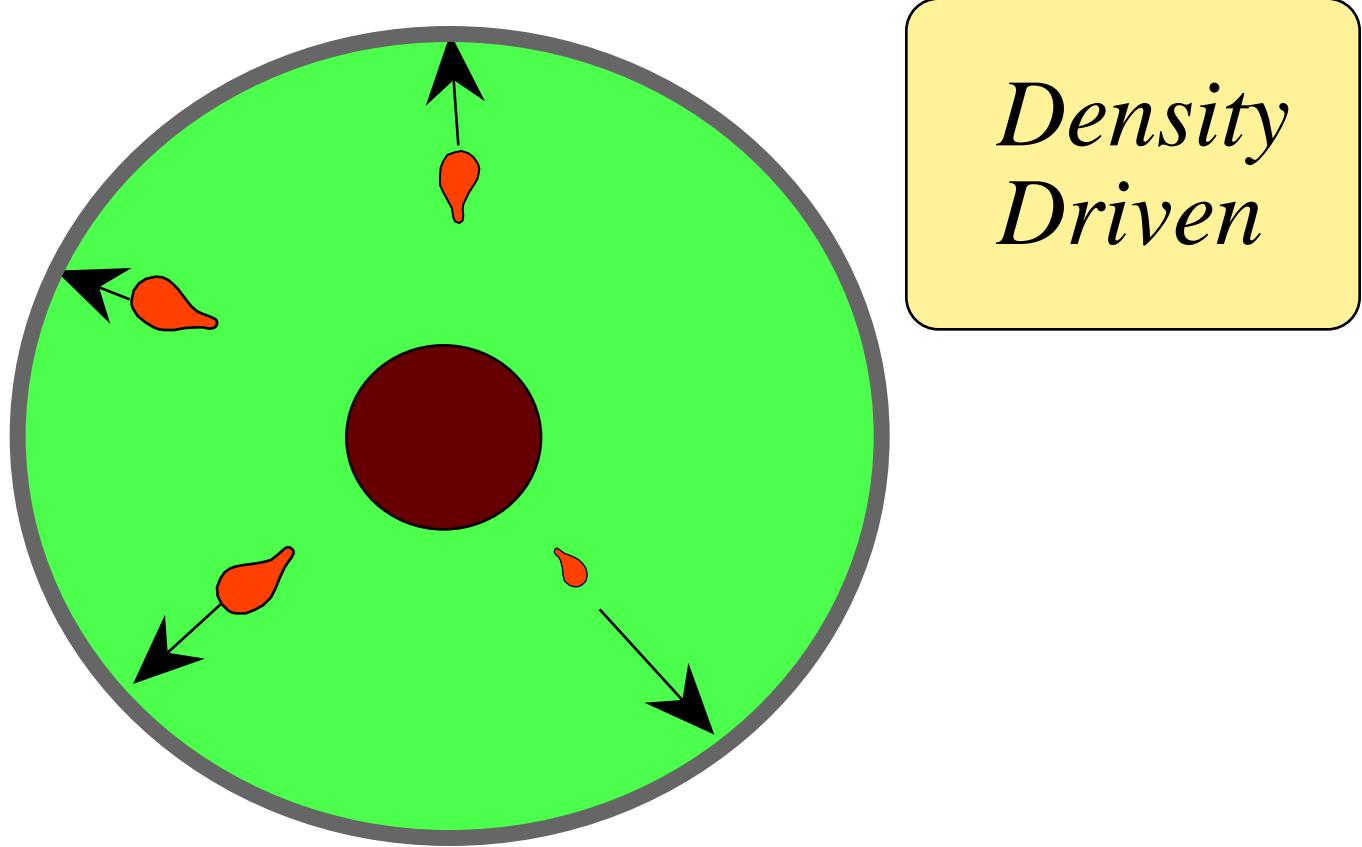
(artist's rendition)

Earth Heats Up and Differentiates



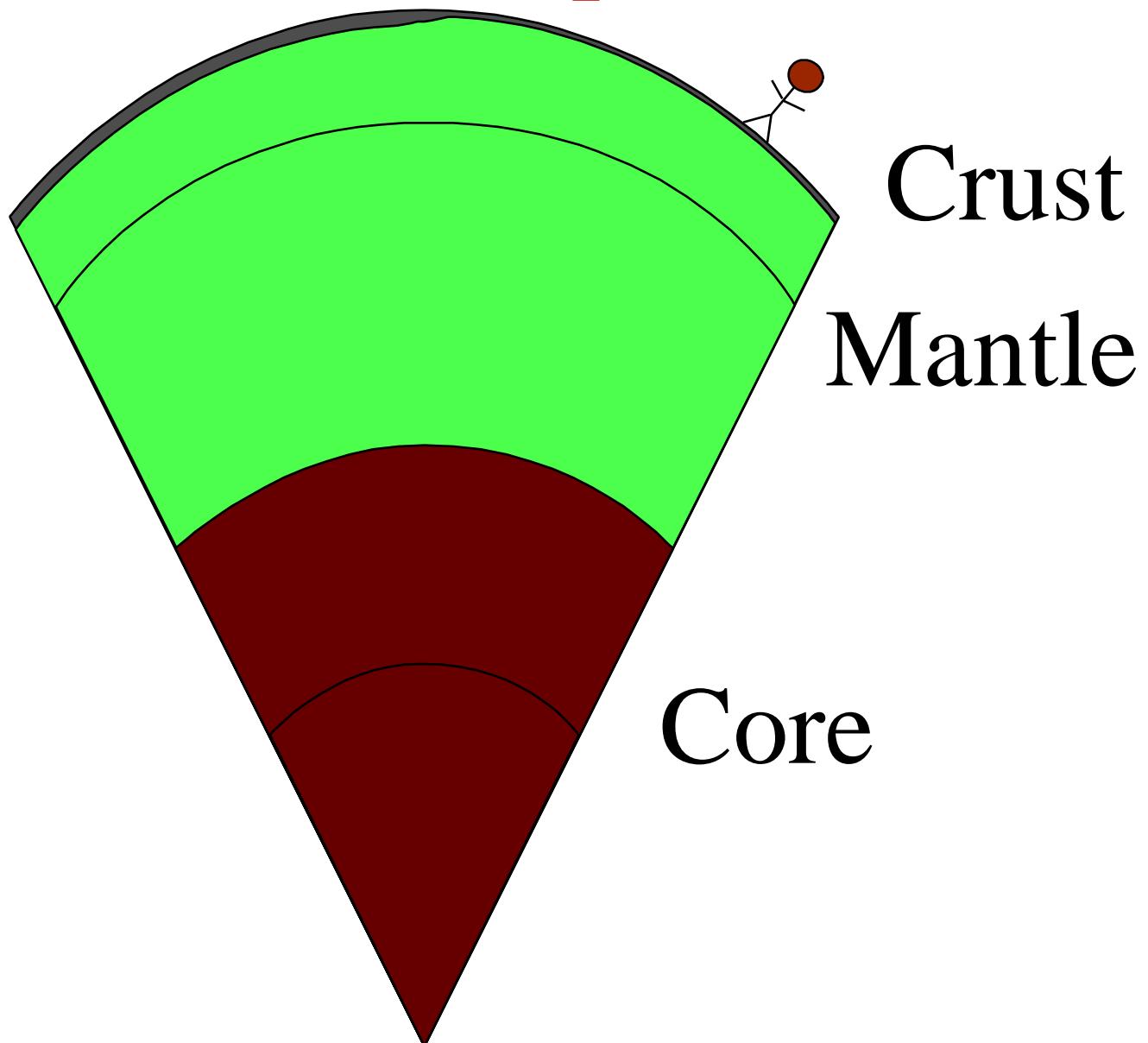
Iron liquid forms and sinks

Earth Heats Up and Differentiates

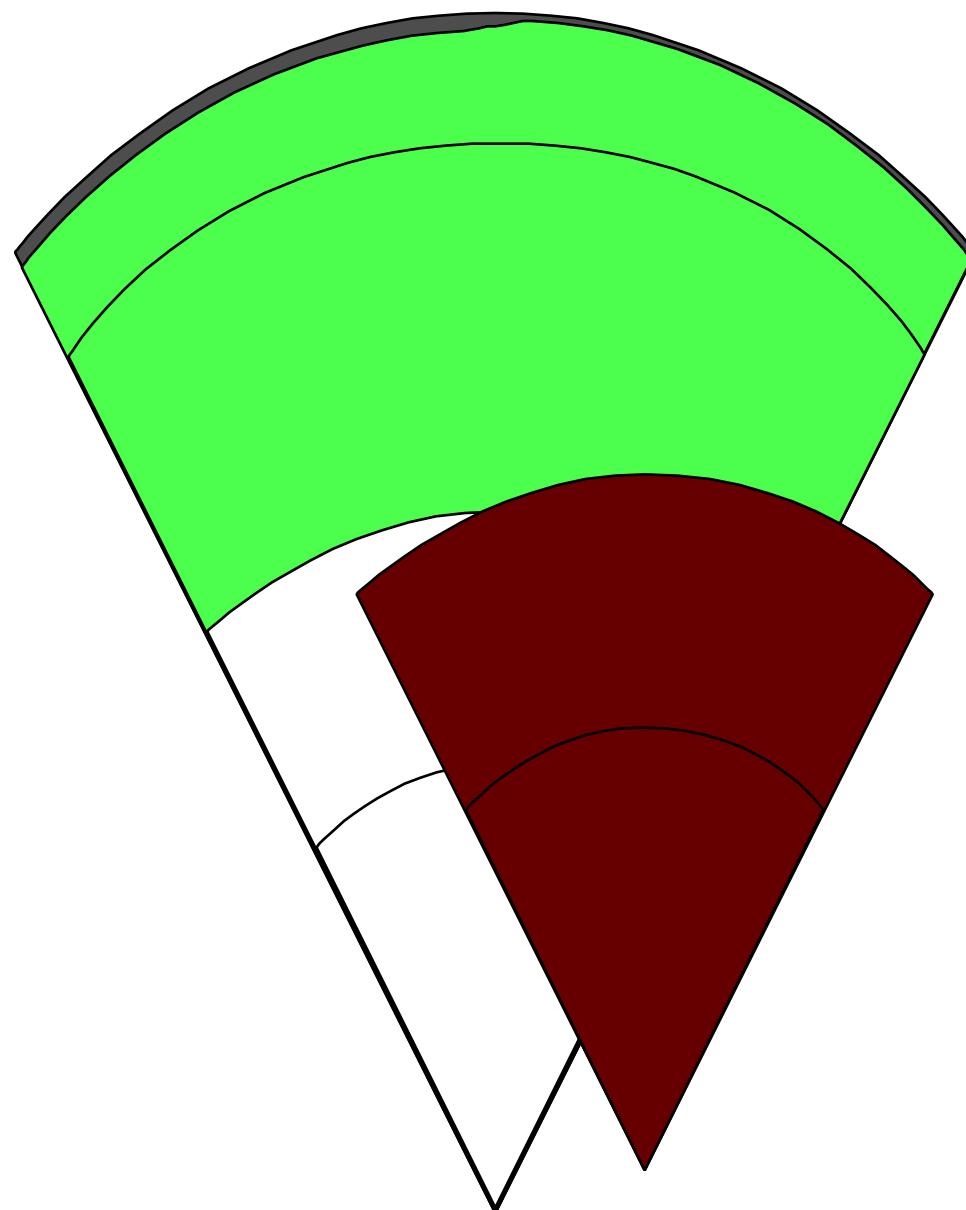


Silica-rich melt rises and forms crust

Differentiation-> Compositional Layering



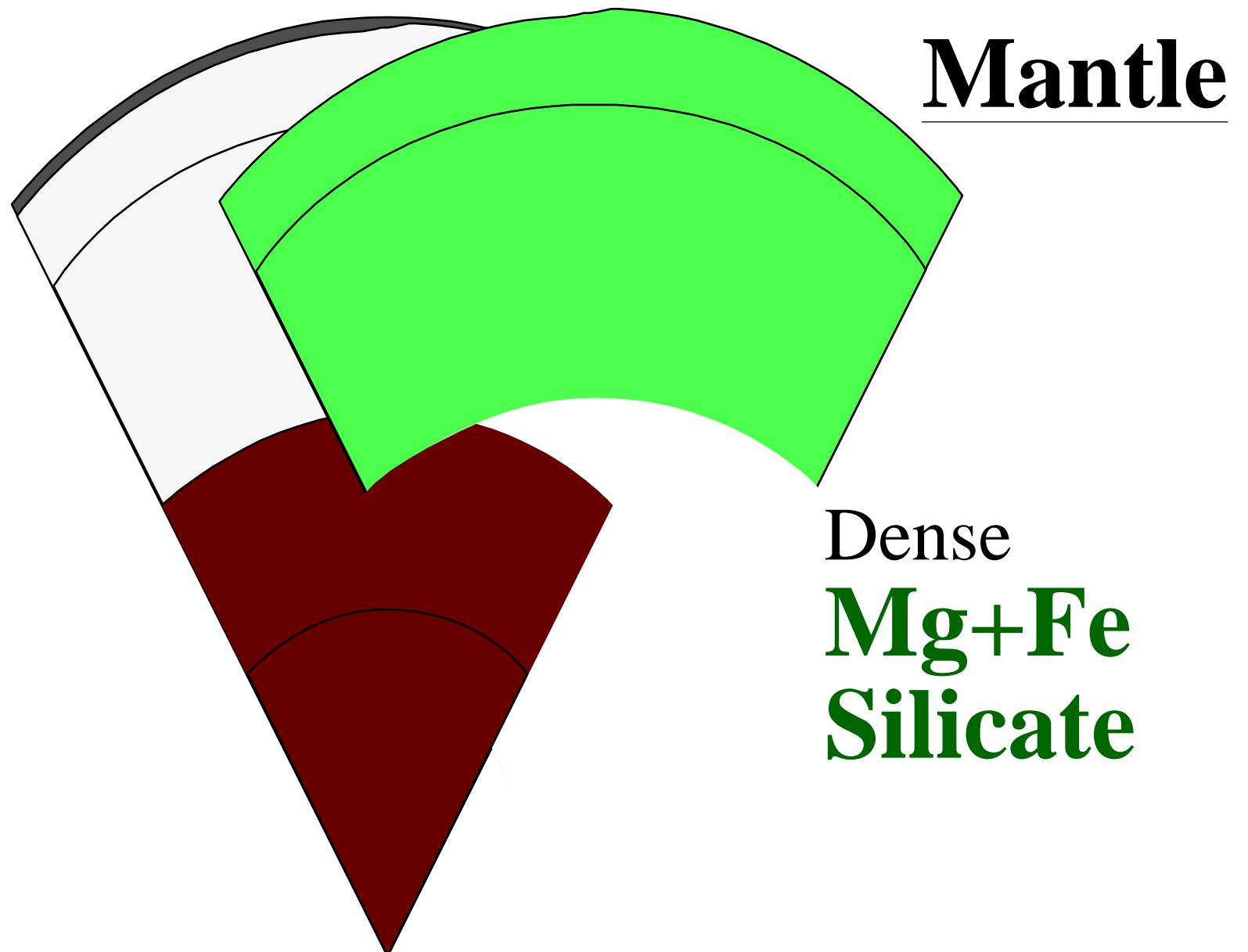
Compositional Layering



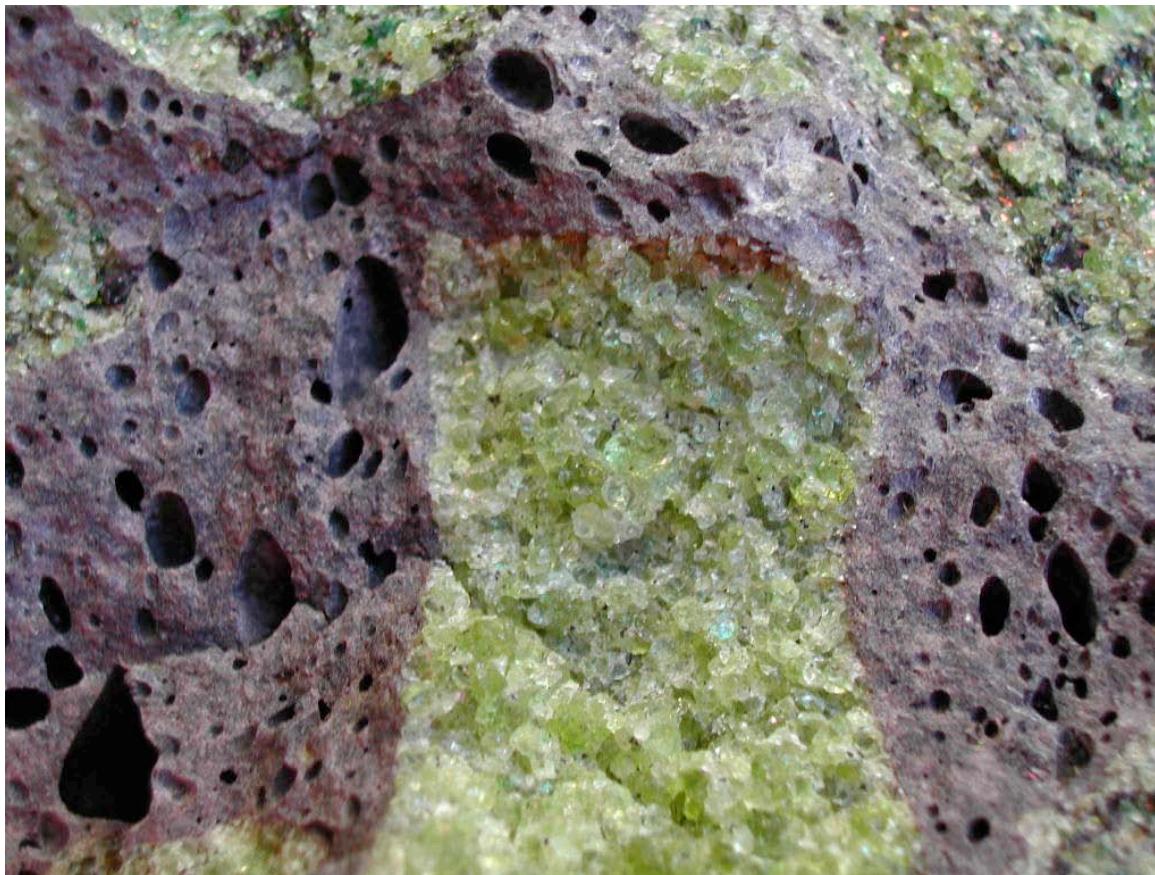
Core

Mostly **Iron**
Some Nickel

Compositional Layering



Pieces of Earth's mantle (green) caught in basaltic lava

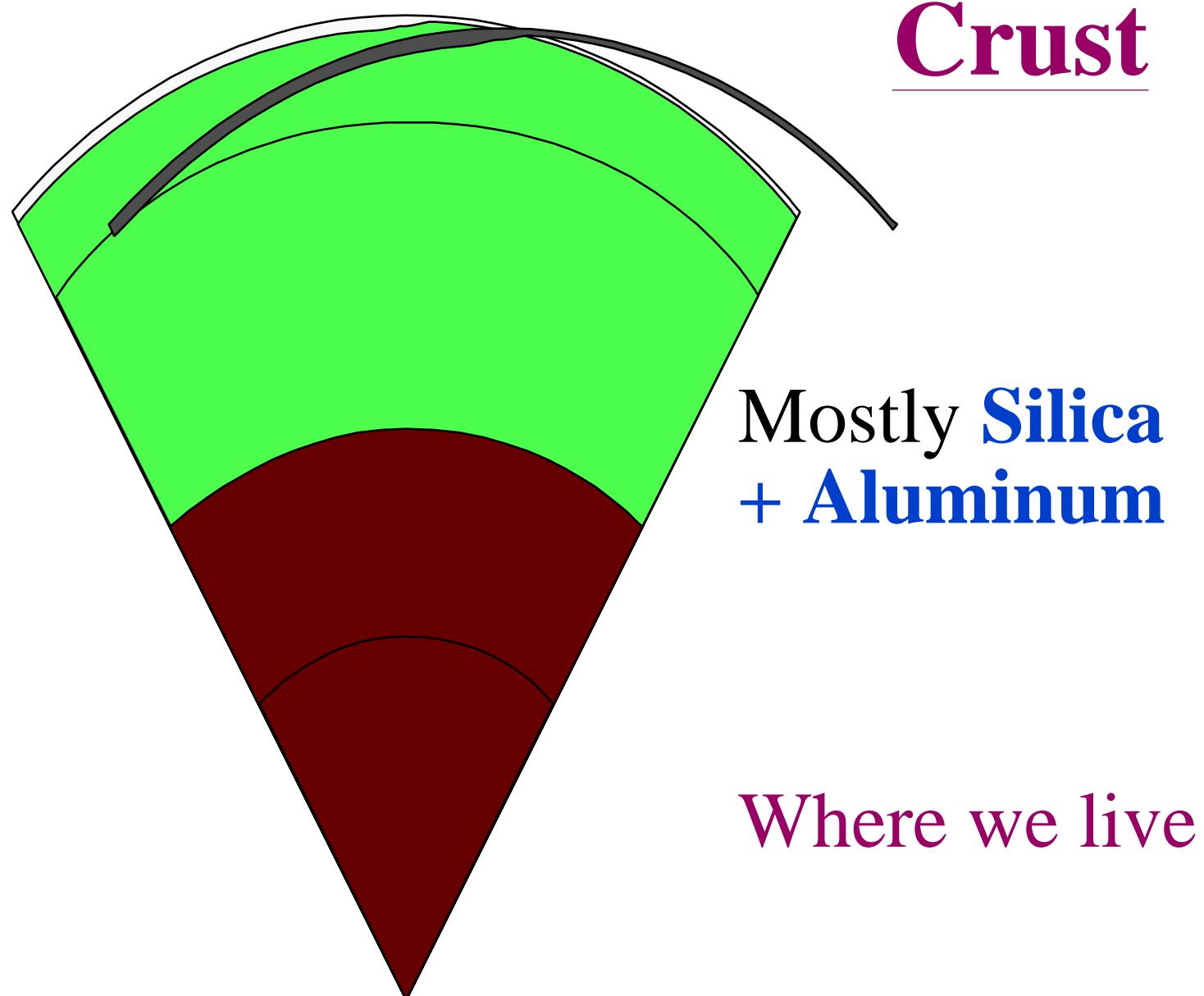


Core-mantle boundary of Asteroid

Mantle minerals in
Fe matrix



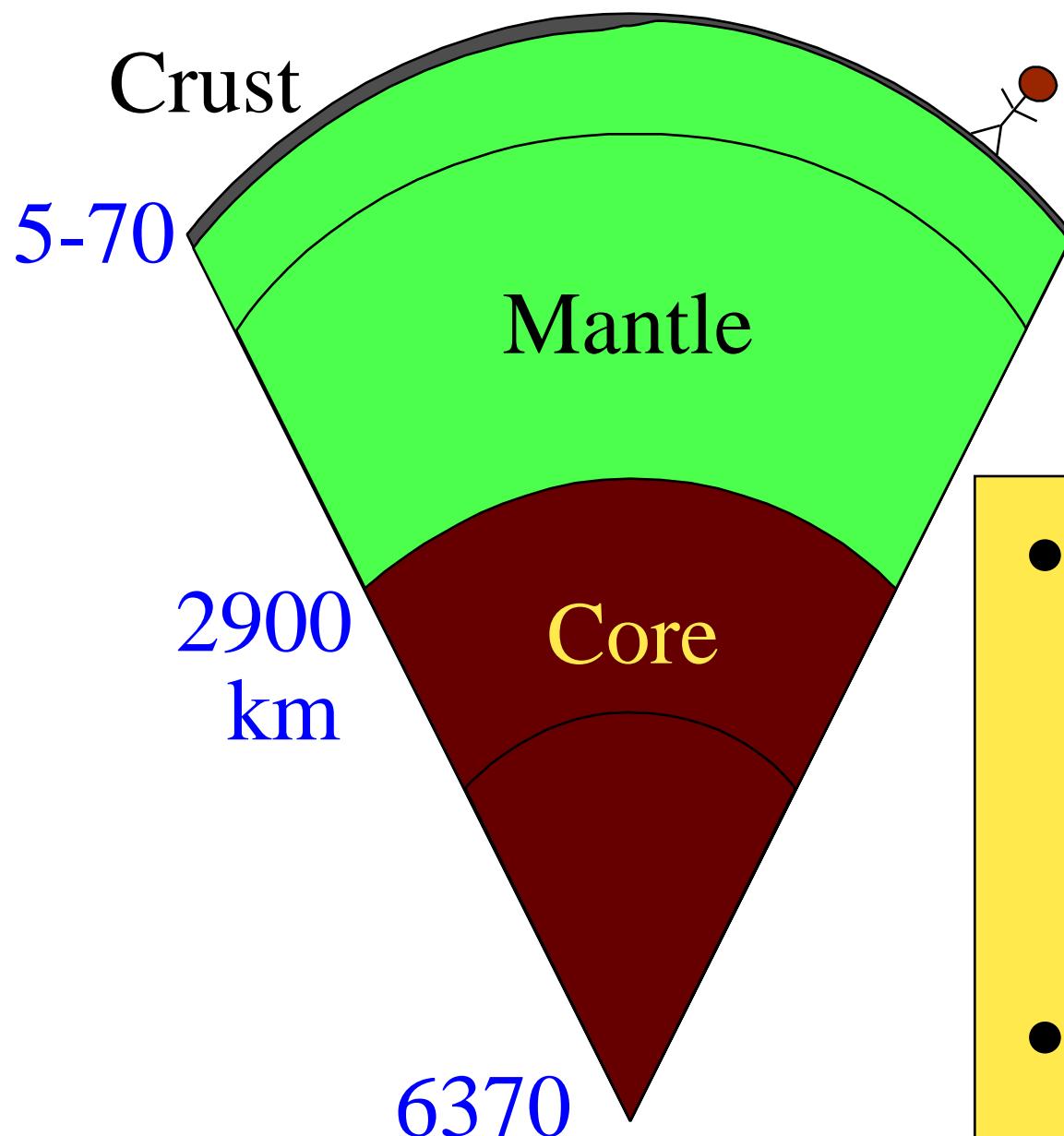
Compositional Layering



Where we live

How do we know?

How do we Know this?



- seismic wave velocities
- meteorites

Compositional Layering

vs.

Mechanical Layering

How Strong or Weak?