

Tropical Convective Forcing of the Southern Hemisphere Circulation: Modulation of Polar Regions via Planetary Wave Trains in the UTL

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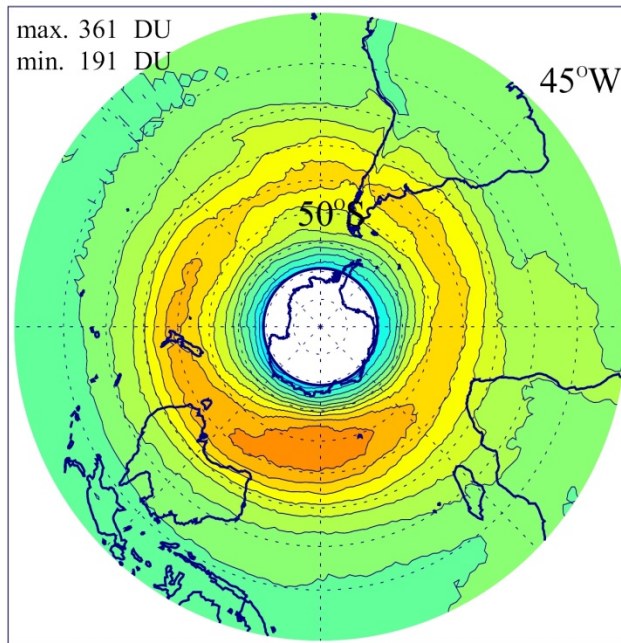
1. Planetary wave radiation from Indonesian convection
 - * Transition from winter to spring (Hitchman and Rogal, 2010a)
 - * August vs. October height anomalies

2. Influence of ENSO on SH circulation
(Rogal and Hitchman 2010b; Rogal et al., 2010)
 - * 3d wave activity flux diagnostics (Kinoshita and Sato 2013a,b)
 - La Nina: column ozone maximum shifted westward $\sim 50^\circ$
 - stronger poleward wave activity flux from Indonesia
 - stronger polar vortex
 - stronger planetary wave one
 - enhanced Ross Sea low

3. Cold East Antarctica during La Nina (Welhouse, 2011)
 - UWNMS simulations Dec. 1-9, 1988

1. Planetary wave radiation from Indonesian convection
 - * Transition from winter to spring (Hitchman and Rogal, 2010a)
 - * August vs. October height anomalies

August 2000

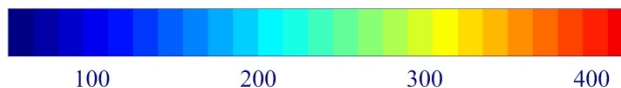
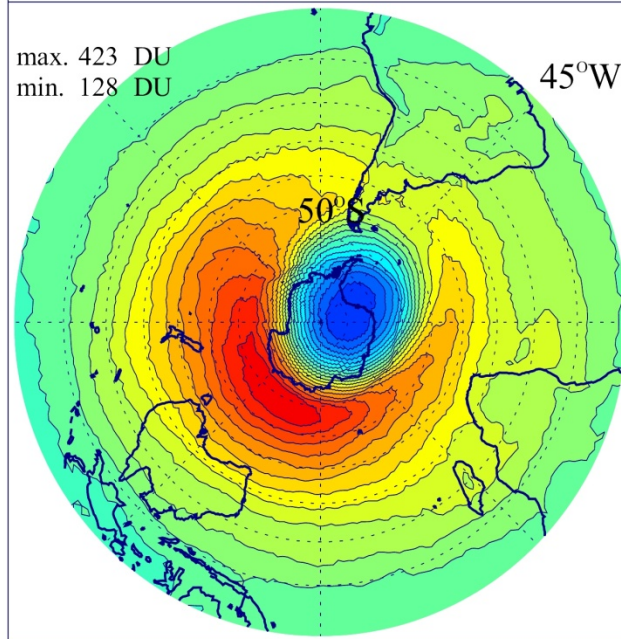


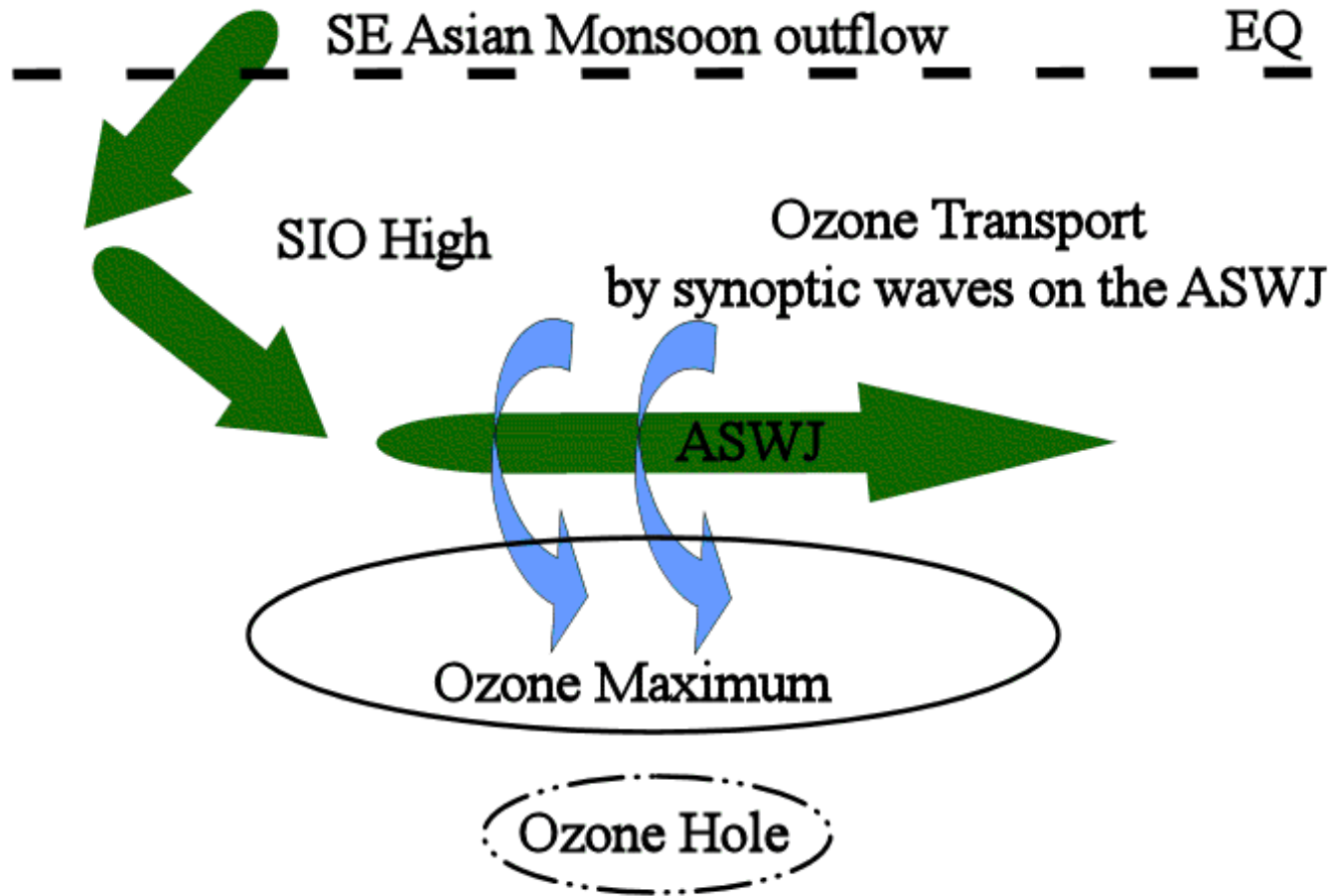
What causes the zonal asymmetry in monthly mean column ozone?

Why is the maximum south of Australia?

What causes the seasonal sharpening of the ozone maximum and its eastward shift during the winter to spring transition?

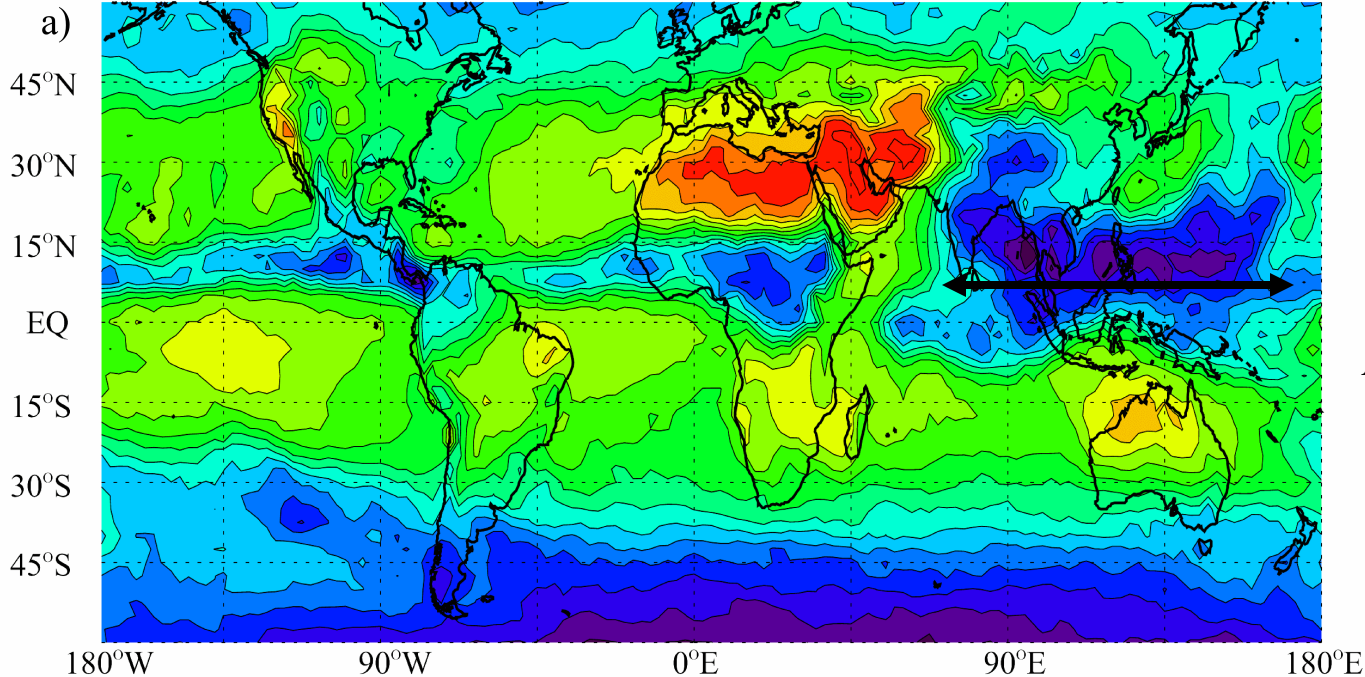
October 2000





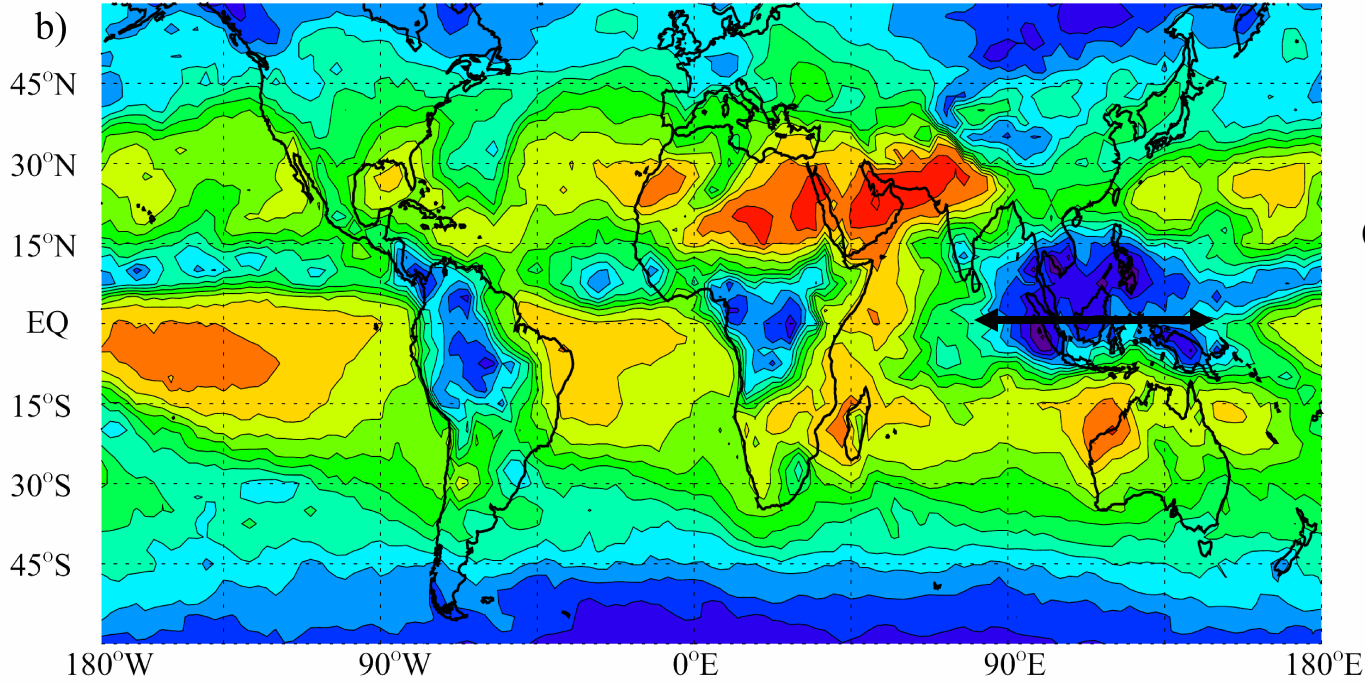
Column ozone maximizes poleward of the Australian Subtropical Westerly Jet in the subsidence warm anomaly (cf. idealized linear model by Wirth, 1993).

The *geographical location* of the ozone croissant and associated wave one structure is partly controlled by convection in Southeast Asia.



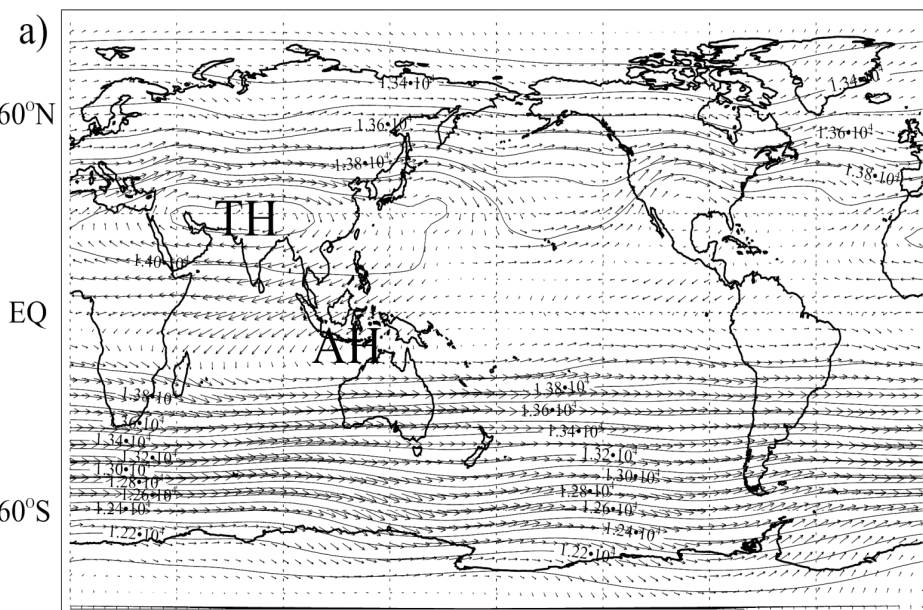
Outgoing Longwave
Radiation

August 2000

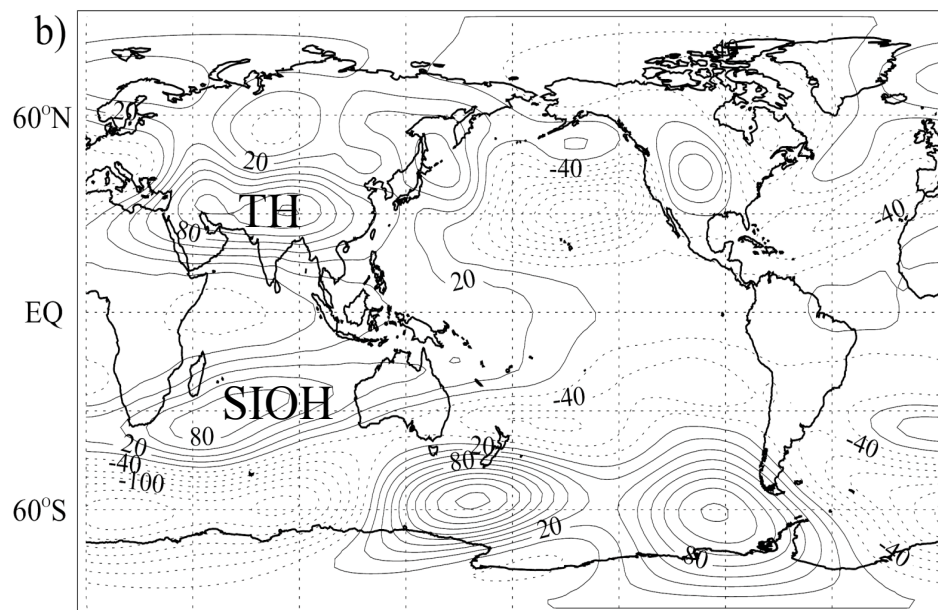


October 2000

ASO 2000 ECMWF



Z 150 hPa heights and winds
interval 100 m

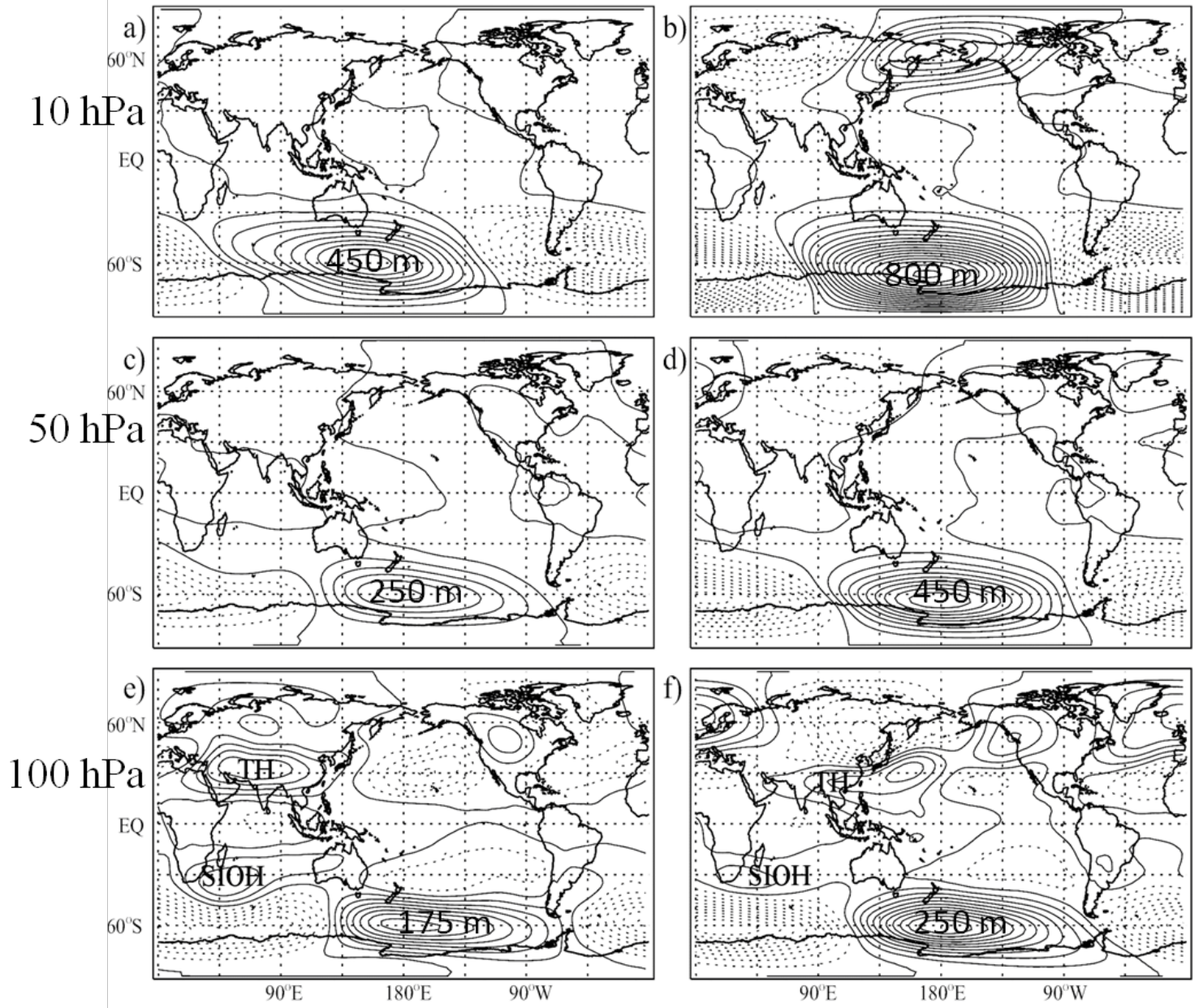


Z' 150 hPa eddy geopotential height
interval 20 m

A planetary wave train emanates from Indonesia into the westerlies of the SH lower stratosphere.

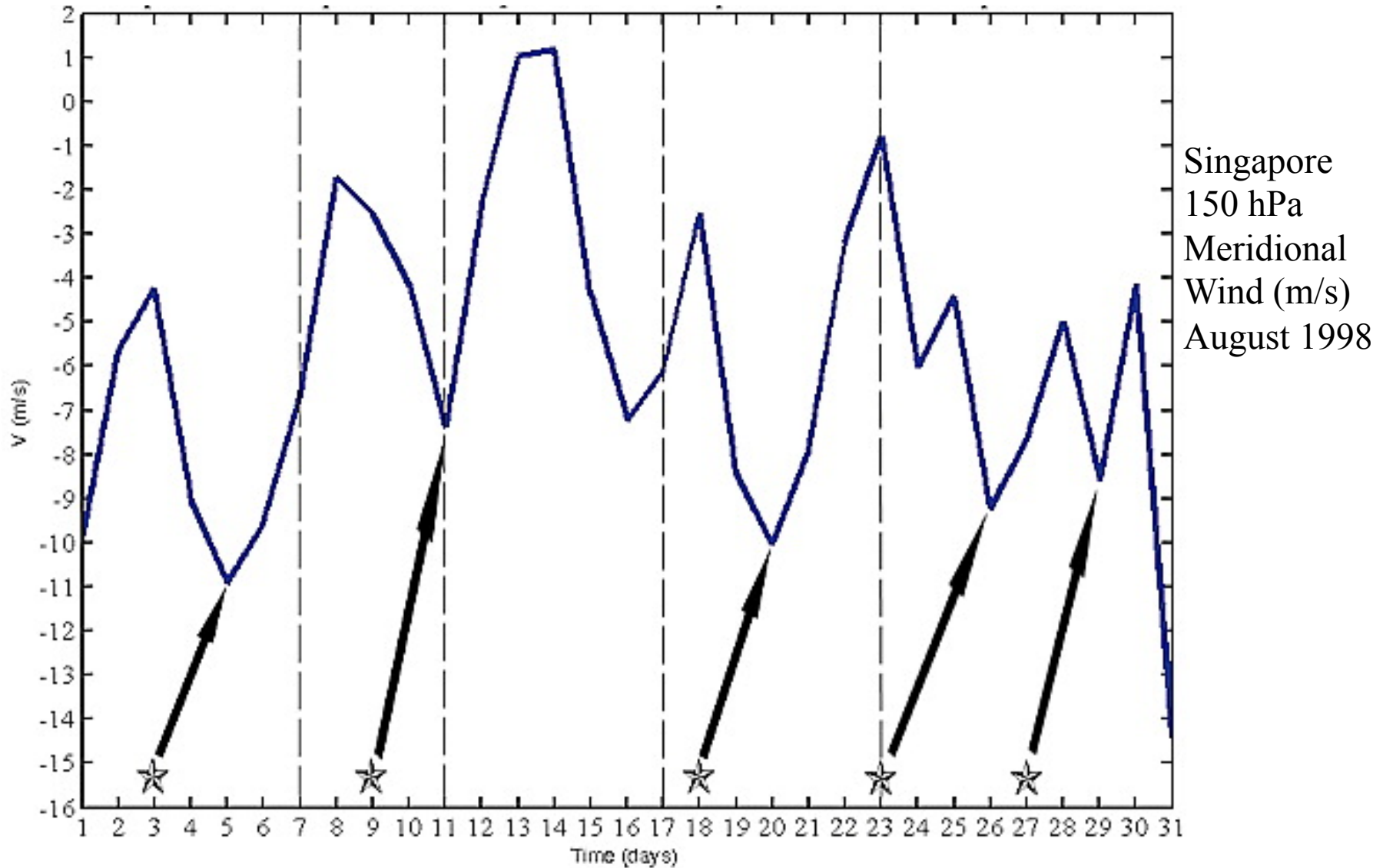
August

October



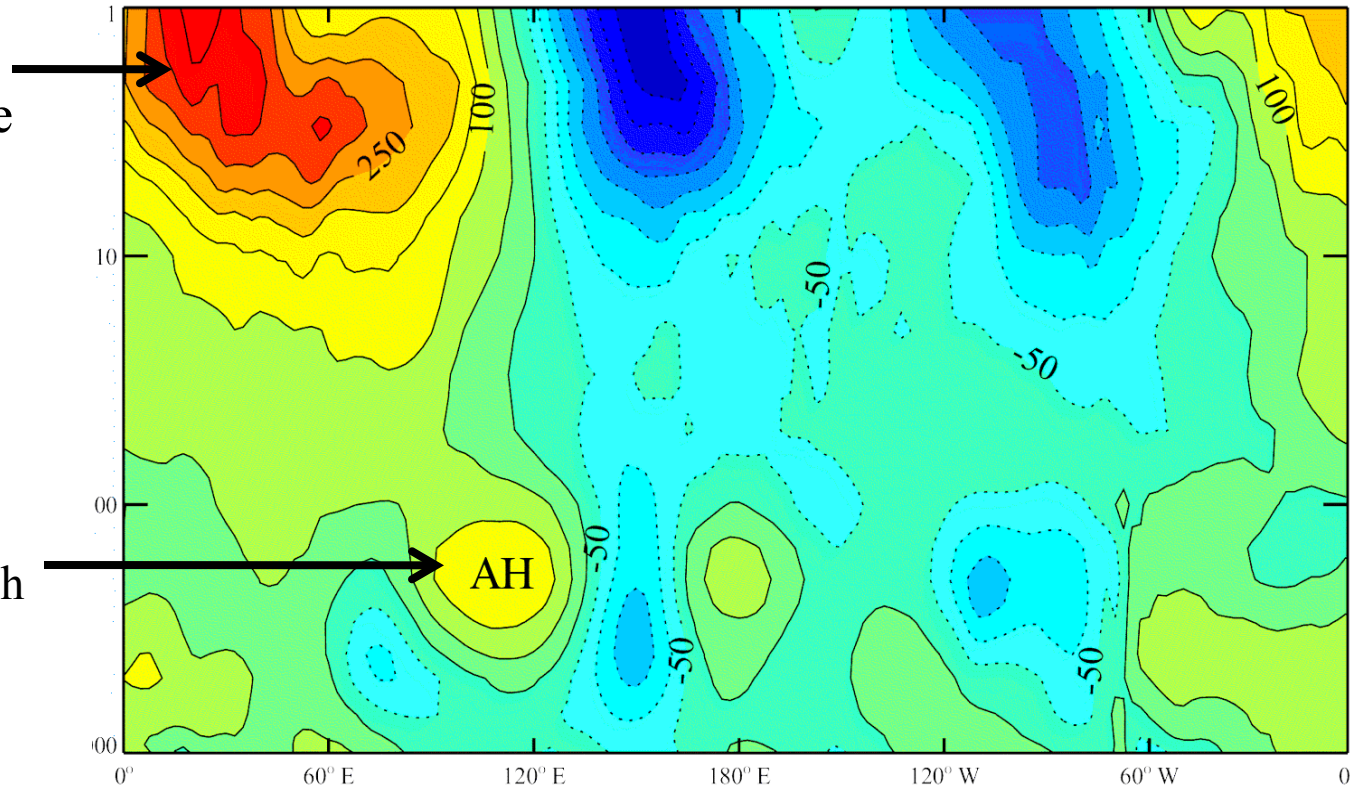
ECMWF monthly mean height anomalies during 1995-2004.

Convective maxima lead to outflow surges into the Southern Indian Ocean UTLS



Asterisk = Indonesia OLR minima

Amplified
Planetary Wave
Ridge

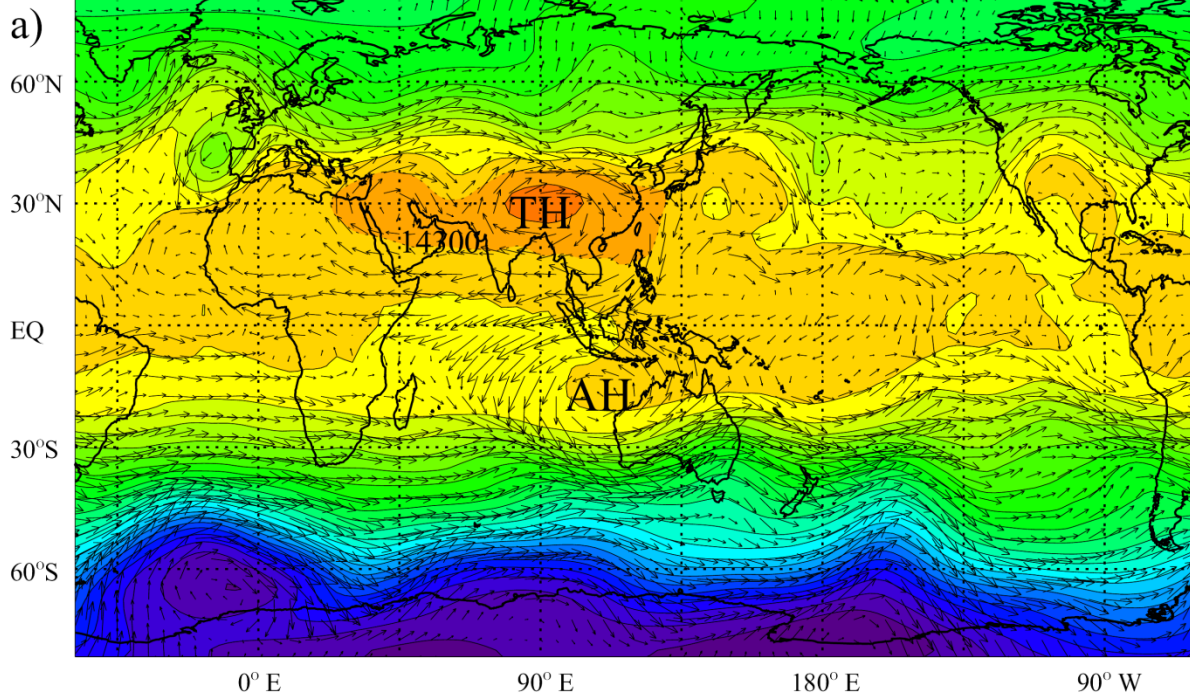


22°S

Amplified
Subtropical High

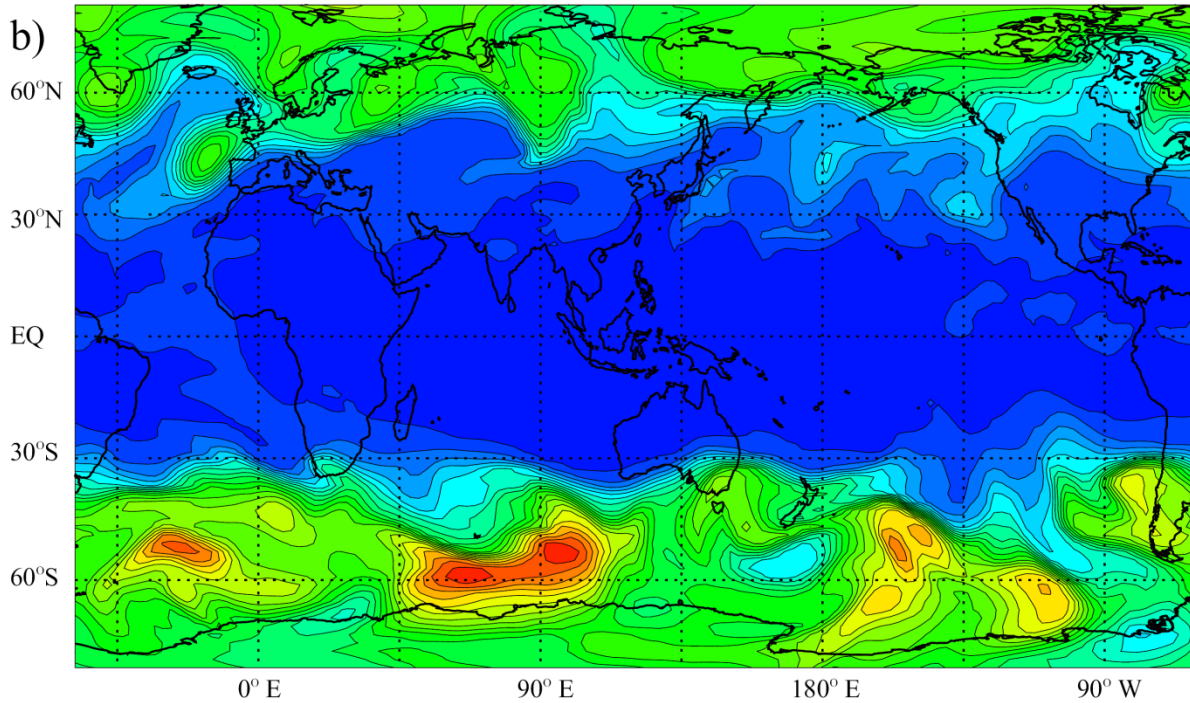
Longitude-altitude section from the surface to the stratopause of eddy geopotential height at 22°S (contour interval 50 m) for 24 August 2000.

Outflow surges into the UTLS over the Indian Ocean amplify the Australian High, which merges with a stalling planetary wave ridge in the winter stratosphere.

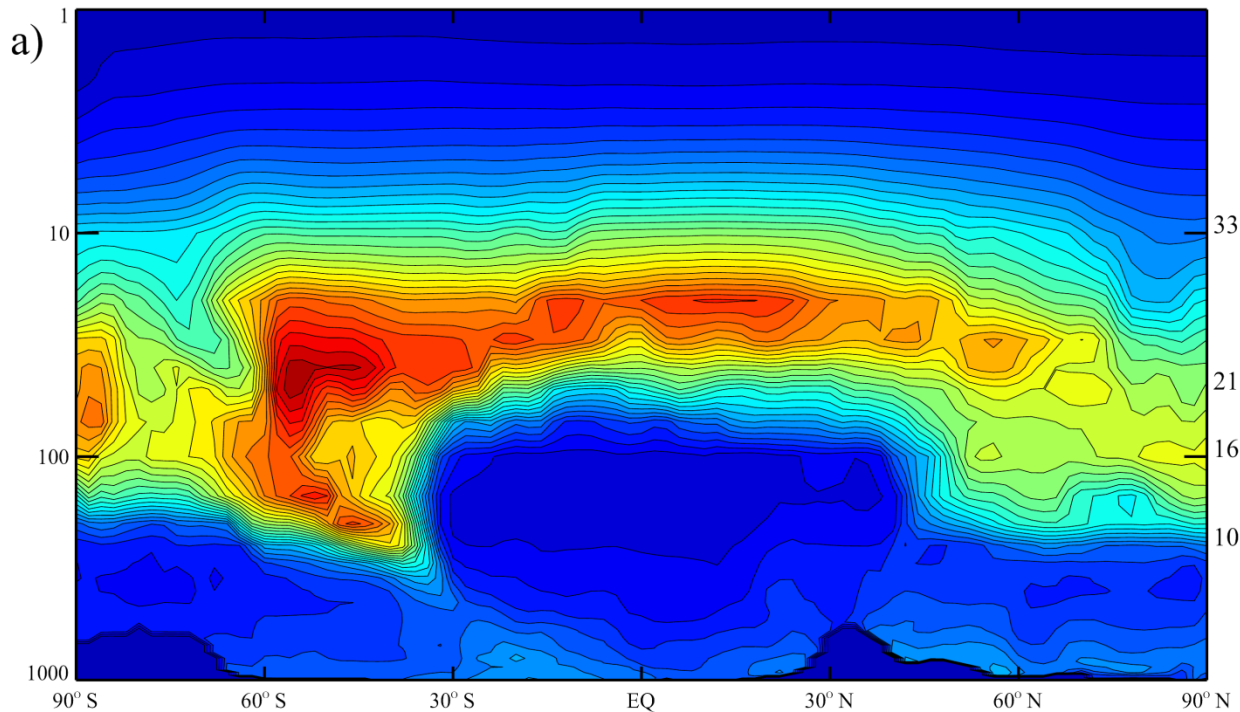


August 24, 2000

150 hPa Z, winds

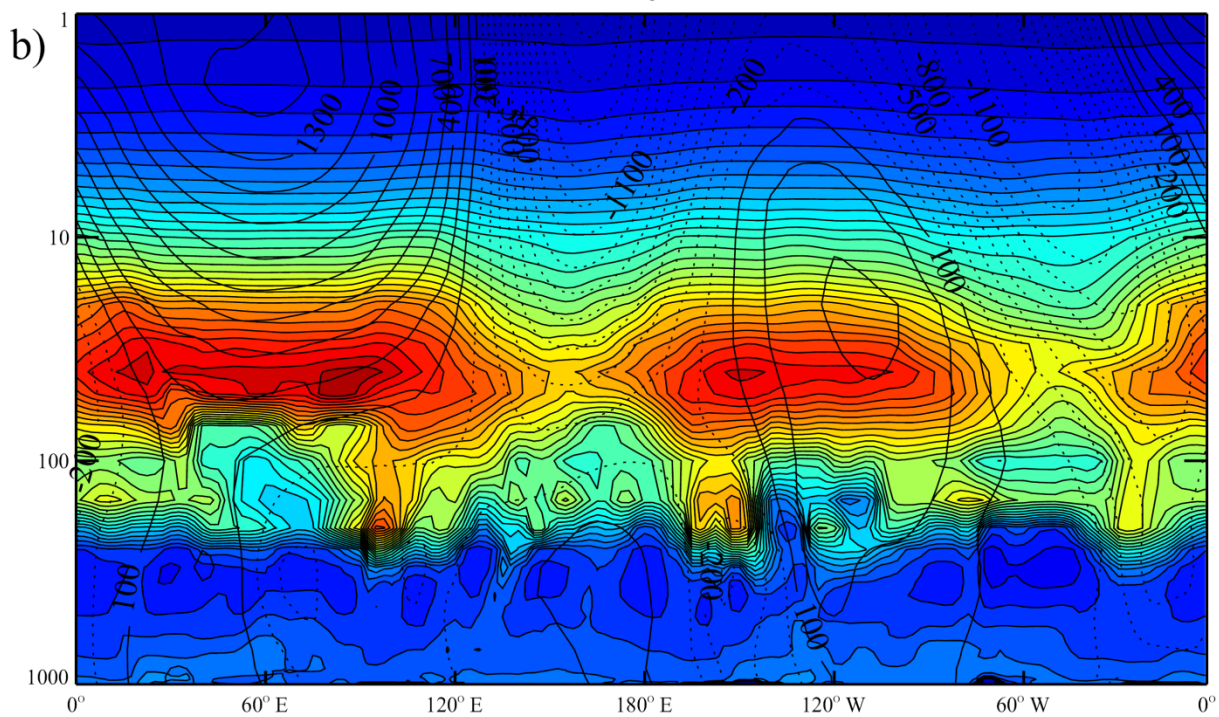


150 hPa GMAO
ozone mixing ratio.



August 24, 2000

GMAO ozone concentration
at 95°E

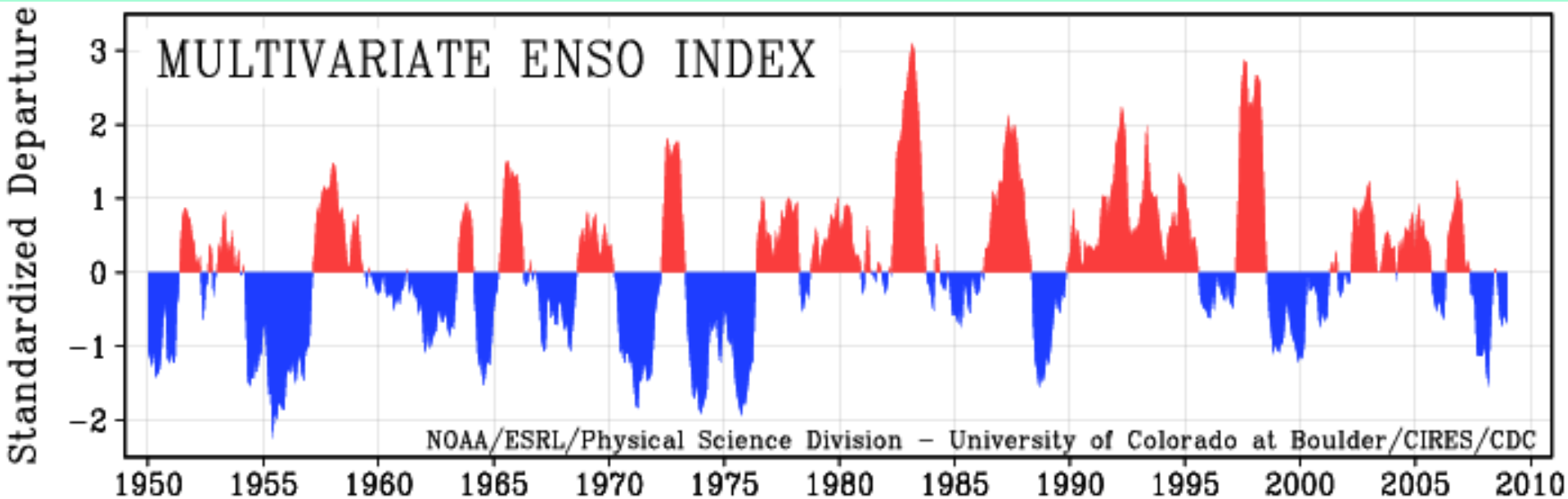


Eddy geopotential height and
Ozone (color) at 47°S

2. Influence of ENSO on SH circulation during SON

- * effects on planetary waves and polar vortex
- * ozone maximum shifted westward during La Nina

ENSO Influences On SH Planetary Waves and Ozone

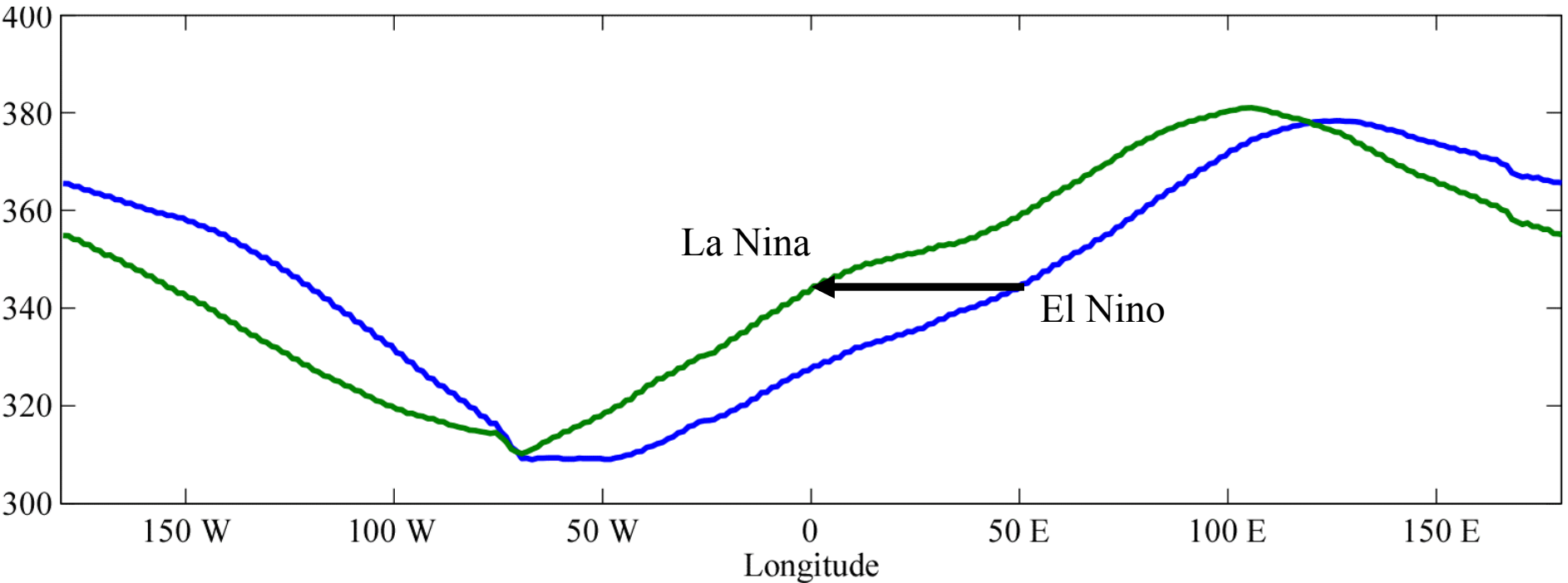


El Nino 1982, 1986, 1987, 1991, 1997, 2002, 2004

La Nina 1983, 1984, 1988, 1998, 1999, 2000

TOMS column ozone 1982-2004

August-September-October 45-60S



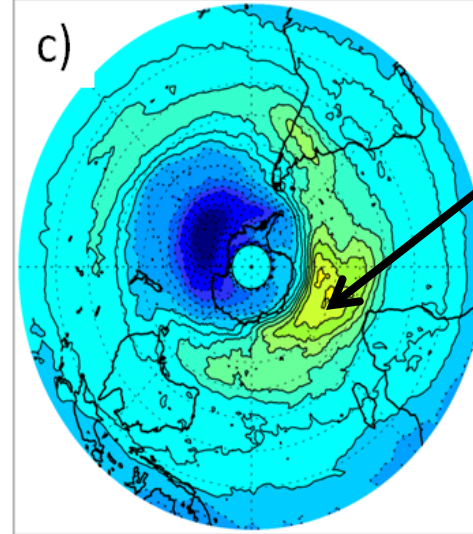
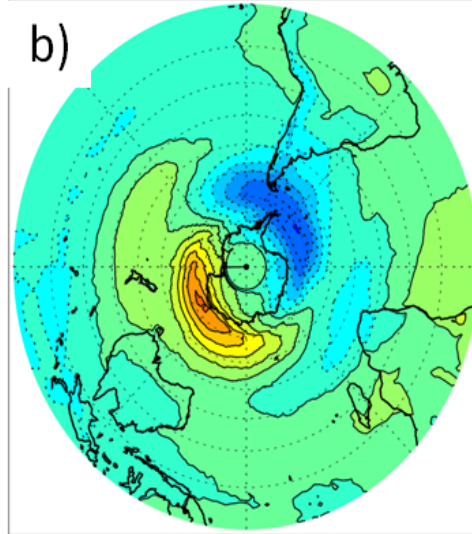
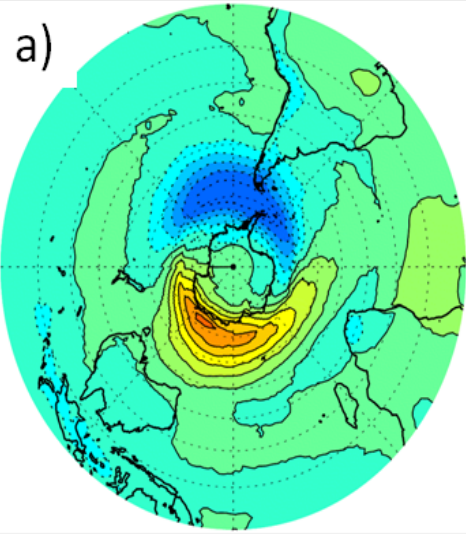
SH column ozone shifts westward with convection during La Nina.

LN

EN

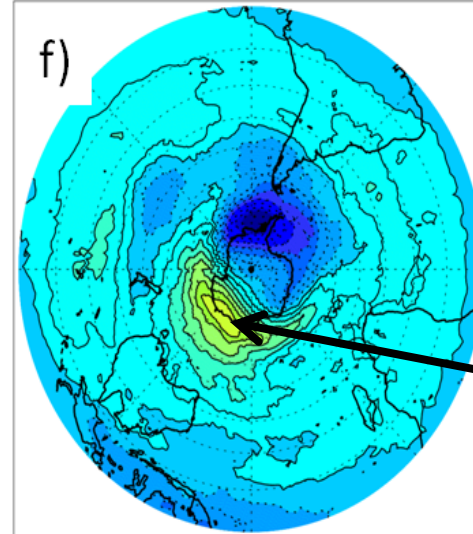
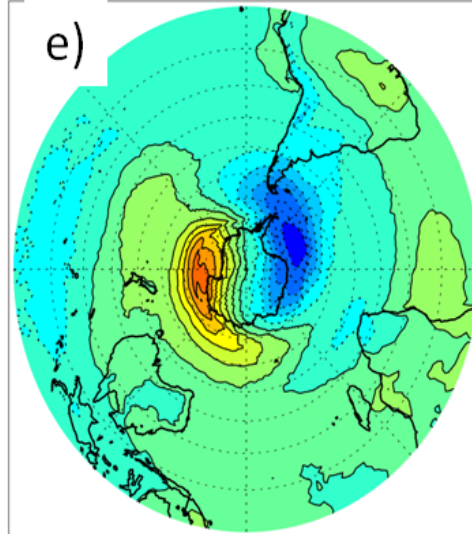
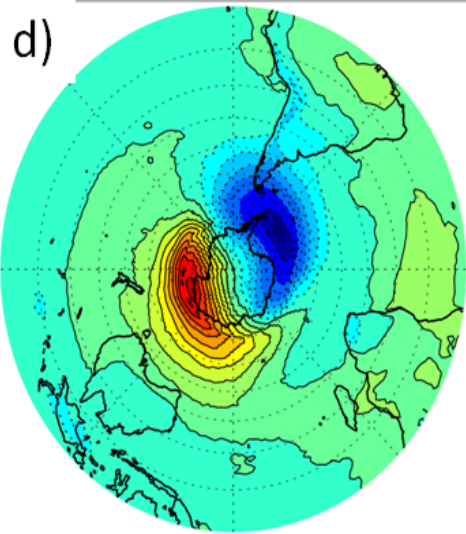
LN - EN

SEPT



Less during
El Niño

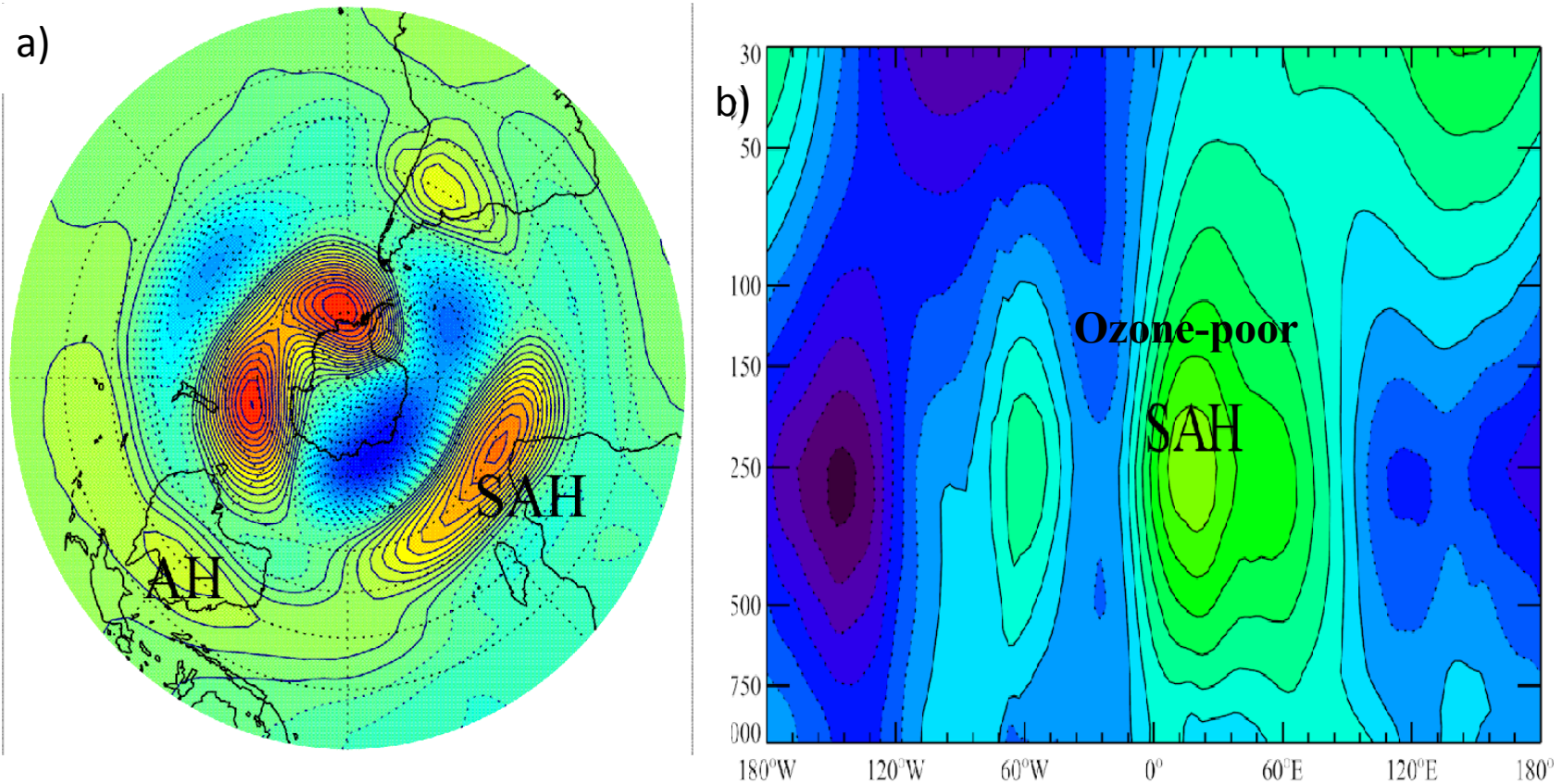
OCT



More during
La Niña

 O_3'

The reason for the westward shift during LN depends on month.



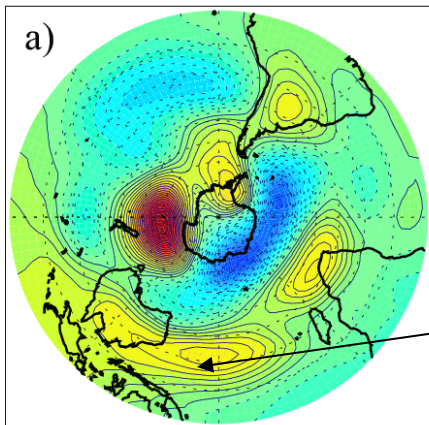
September Z' for El Niño at a) 150 hPa and b) 35°S.

EN favors a stronger South African High (SAH) in the UTLs, which displaces the tropopause upward, causing a cold anomaly in the lower stratosphere and reduced column ozone.

150 hPa Z'

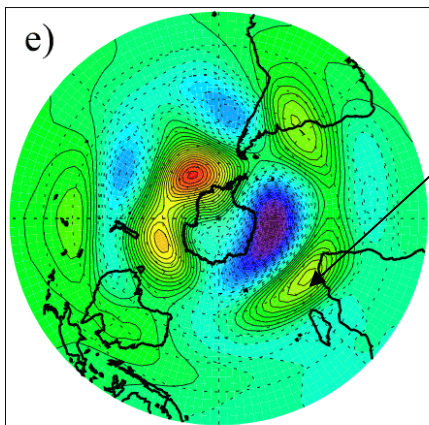
October

La Nina



Strong
Australian
High

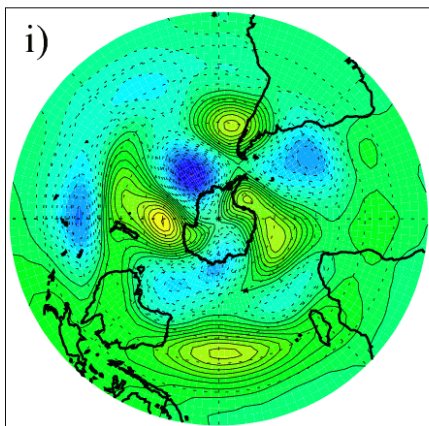
El Nino



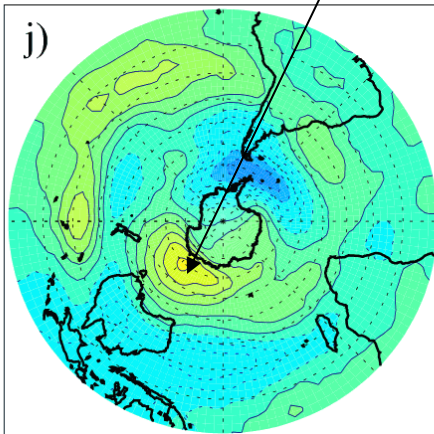
South
African
High

La Nina:
deep barotropic warm feature
high column ozone
South of Australia

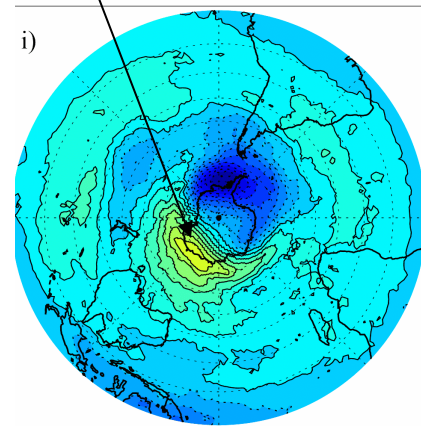
La - El
Nino



100 hPa T'



TOMS ozone

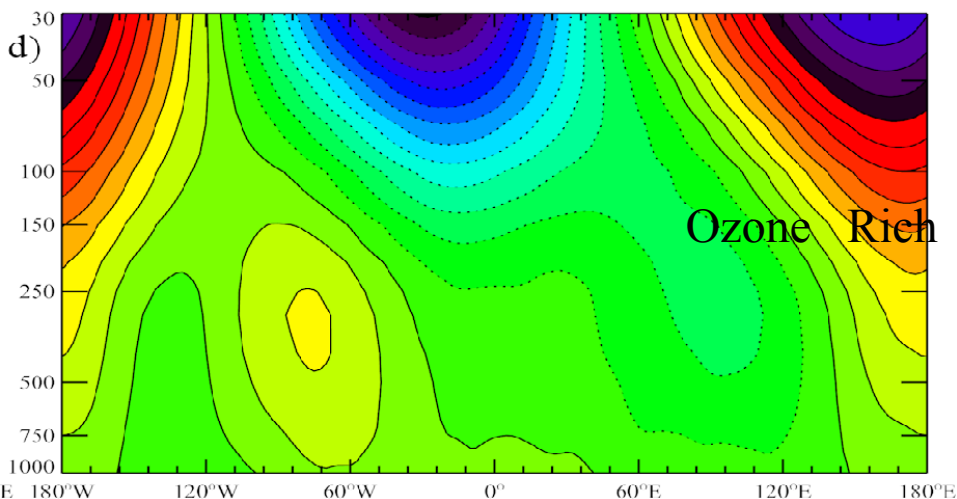
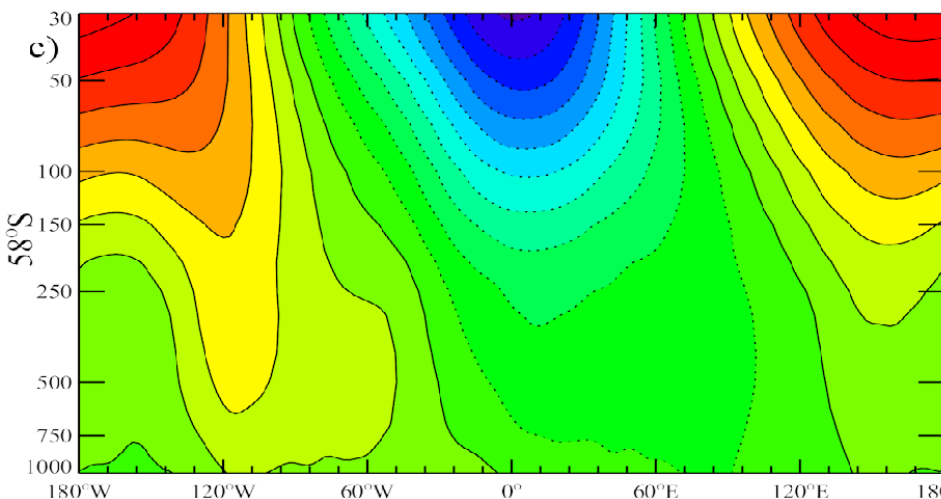
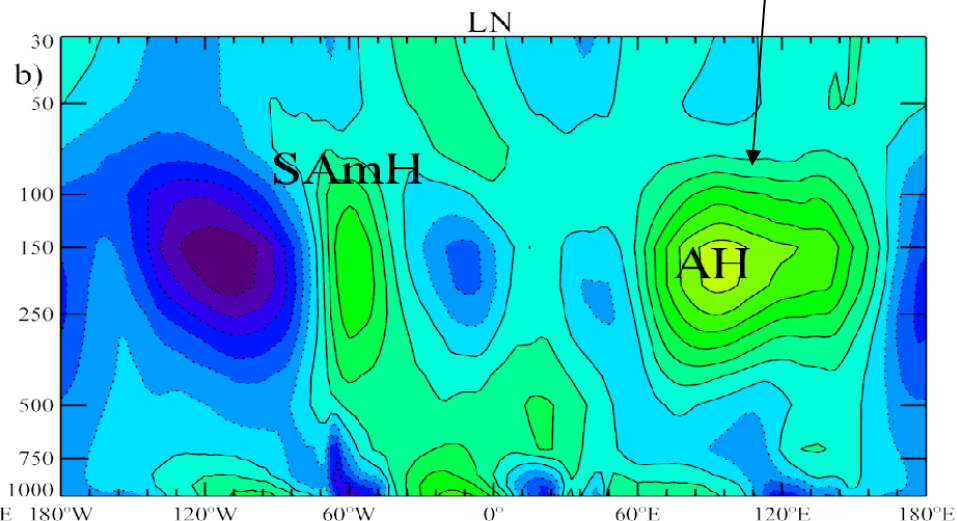
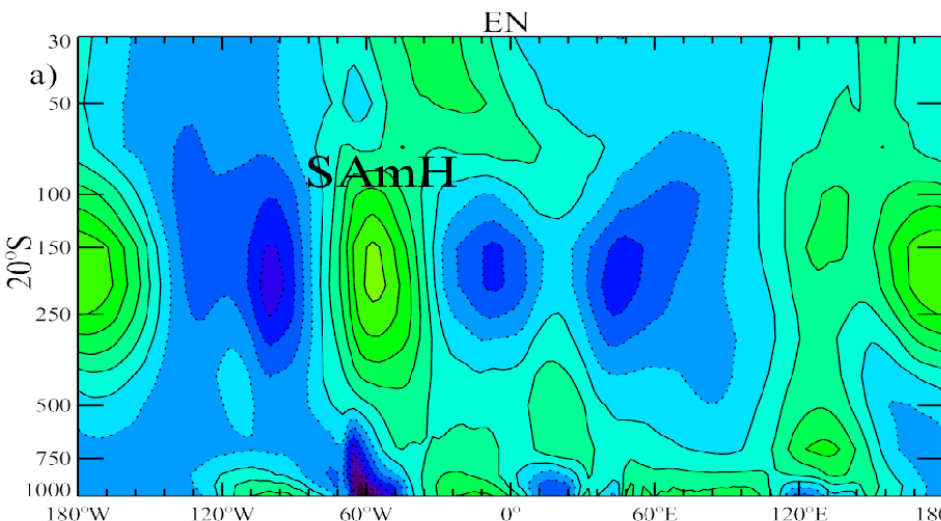


El Nino

October Z'

La Nina

Australian High amplified



Stronger More Asymmetric
Vortex during La Nina

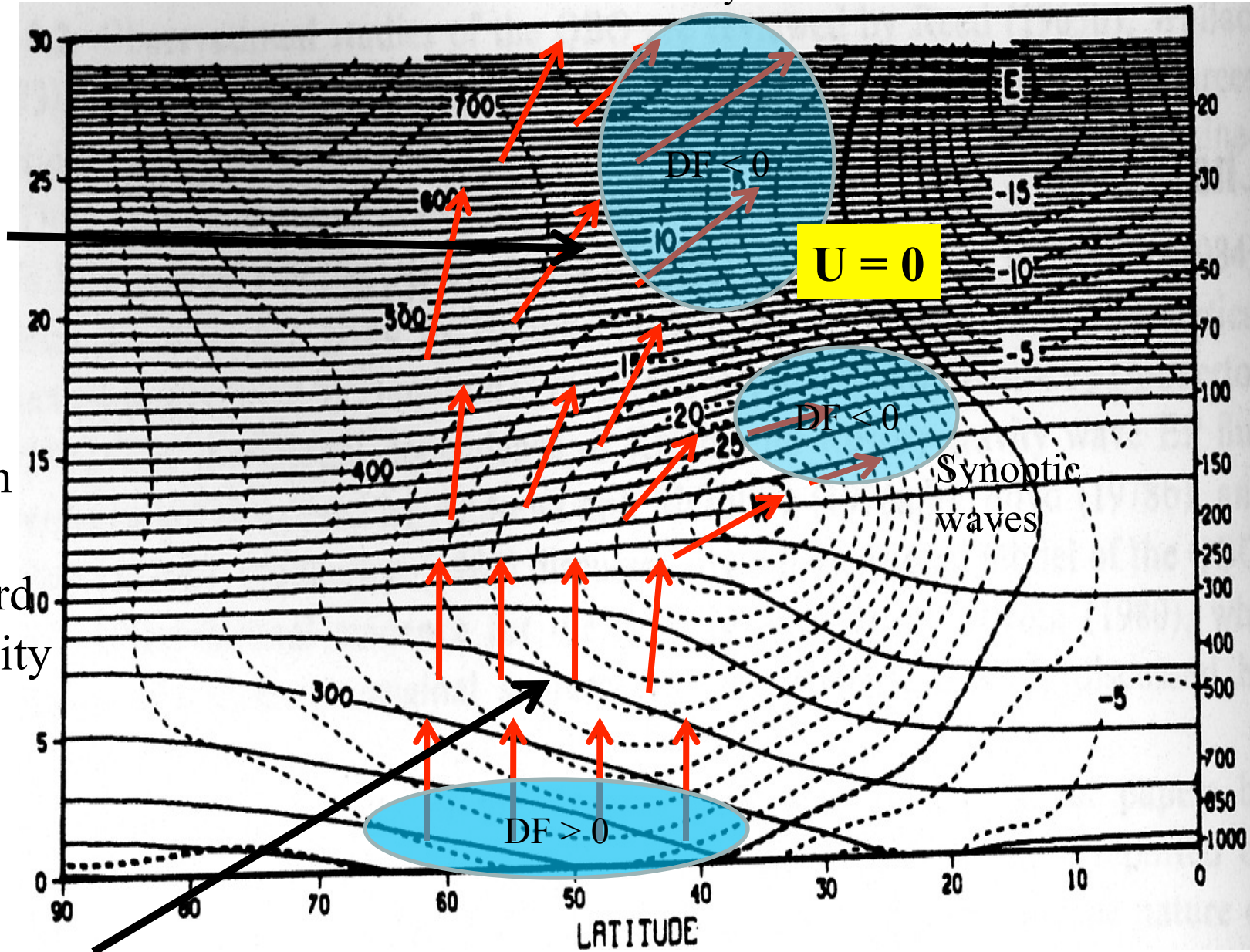
2. Influence of ENSO on SH circulation during SON

* 3d wave activity flux diagnostics (Kinoshita and Sato 2013a,b)

Planetary waves

$$F_y \sim -u'v' \\ \sim G_z A$$

poleward
momentum
flux =
equatorward
wave activity
flux



$$F_z \sim f v'\theta' \sim G_z A \quad \text{poleward heat flux} = \text{upward wave activity flux}$$

3D wave activity flux expressing wave-mean interaction (left) and wave propagations (right)

$$F_{1M(x)} \equiv \rho_0 \left(\overline{u'^2} + \frac{\bar{u}_y - f}{f} \overline{S_{(p)}} - \bar{u}_z \frac{\overline{u' \Phi'_z}}{N^2} \right),$$

$$F_{1M(y)} \equiv \rho_0 \left(\overline{u'v'} - \frac{\bar{u}_x}{f} \overline{S_{(p)}} - \bar{u}_z \frac{\overline{v' \Phi'_z}}{N^2} \right),$$

$$F_{1M(z)} \equiv \rho_0 \left(\overline{u'w'} + \bar{u}_x \frac{\overline{u' \Phi'_z}}{N^2} + (\bar{u}_y - f) \frac{\overline{v' \Phi'_z}}{N^2} \right)$$

$$F_{1W(x)} \equiv \rho_0 (\overline{u'^2} - \bar{S}),$$

$$F_{1W(y)} \equiv \rho_0 (\overline{u'v'}),$$

$$F_{1W(z)} \equiv \rho_0 \left(\overline{u'w'} - f \frac{\overline{v' \Phi'_z}}{N^2} \right),$$

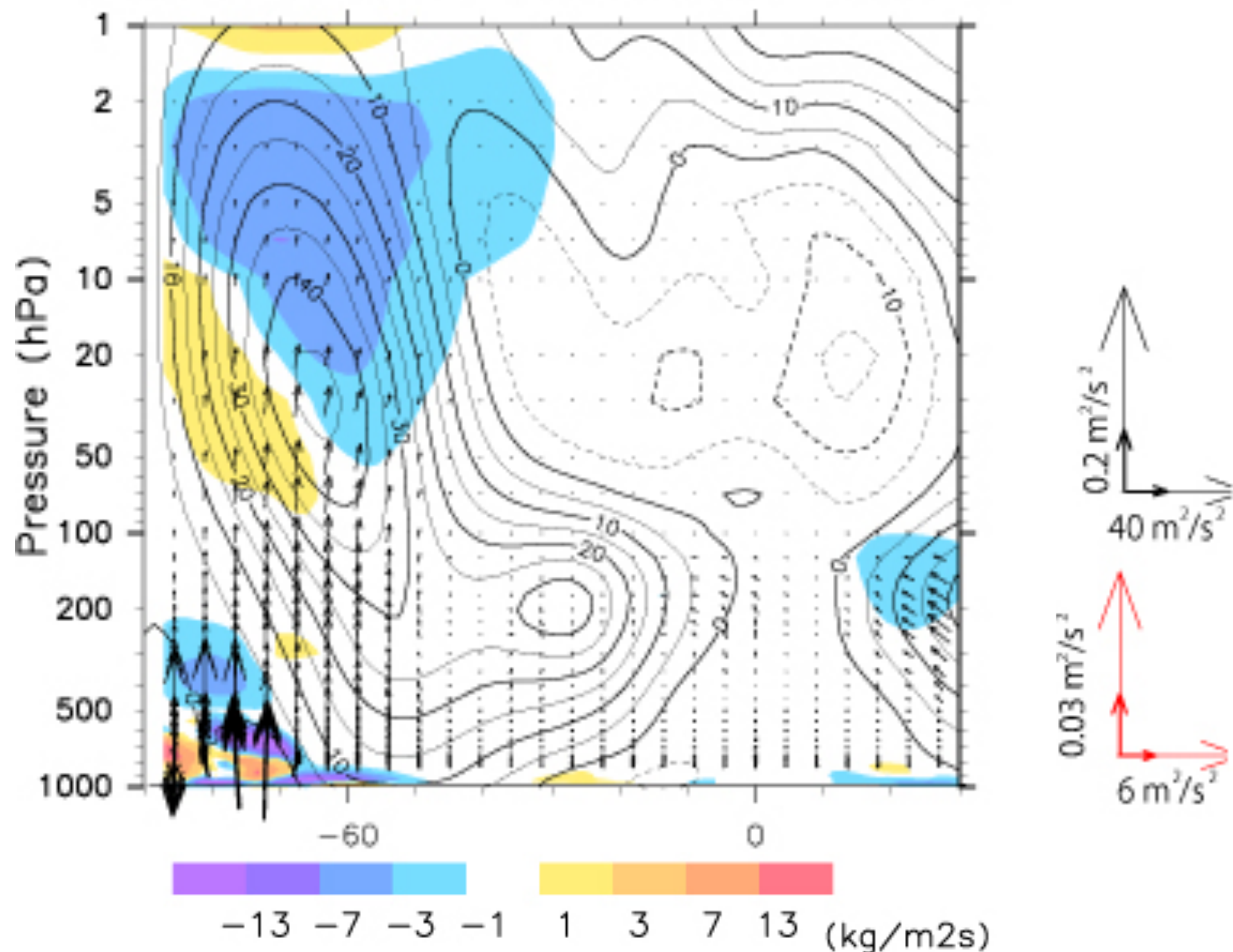
$$\mathbf{F}_{1W} = \hat{\mathbf{C}}_g \left(\frac{E'}{\hat{\mathbf{C}}_{(x)}} \right)$$

To examine quasi-stationary waves, we use the extend Hilbert transform instead of time mean (Sato et al. 2013)

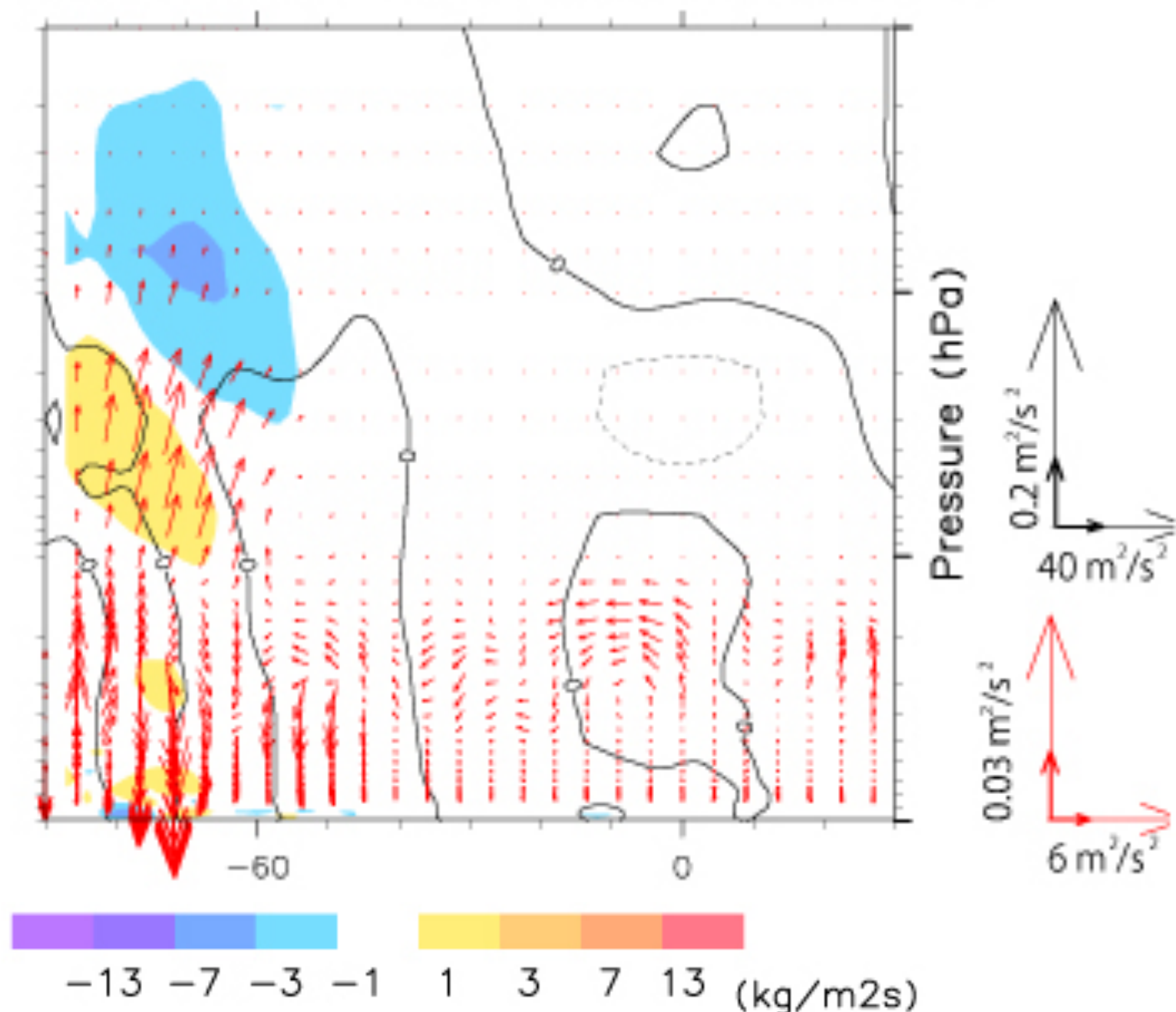
- Data : ERA-Interim reanalysis data
- Focus on : August ~ October
 - ◆ El Nino (EN) : 1991, 1997, 2002, 2004, 2006, 2009
 - ◆ La Nina (LN) : 1998, 1999, 2000, 2007

We use Ocean Nino Index (ONI) in NOAA as an index of ENSO.

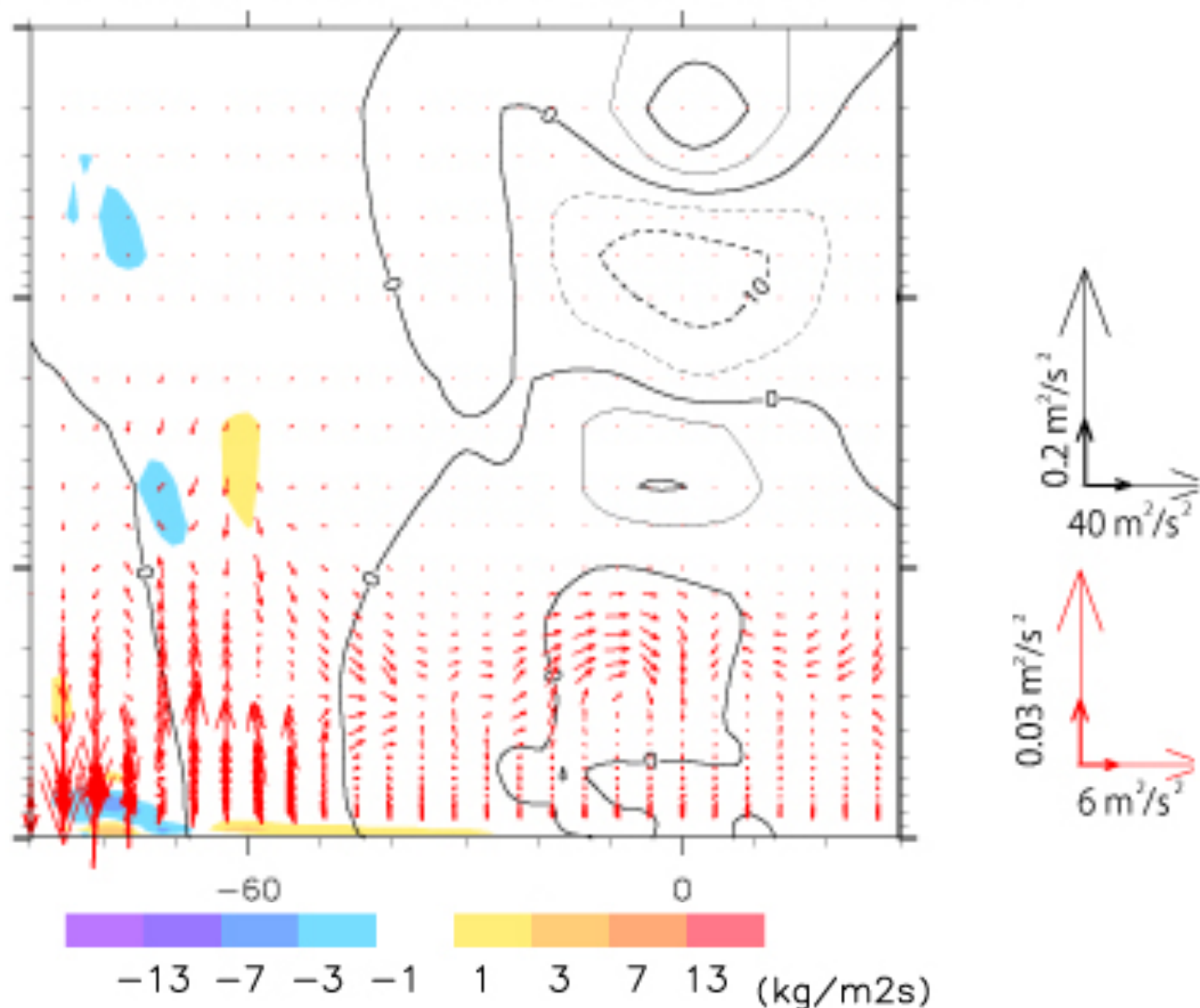
Oct-CLIM Zonal Mean Wave Activ. Flux (s1-3)

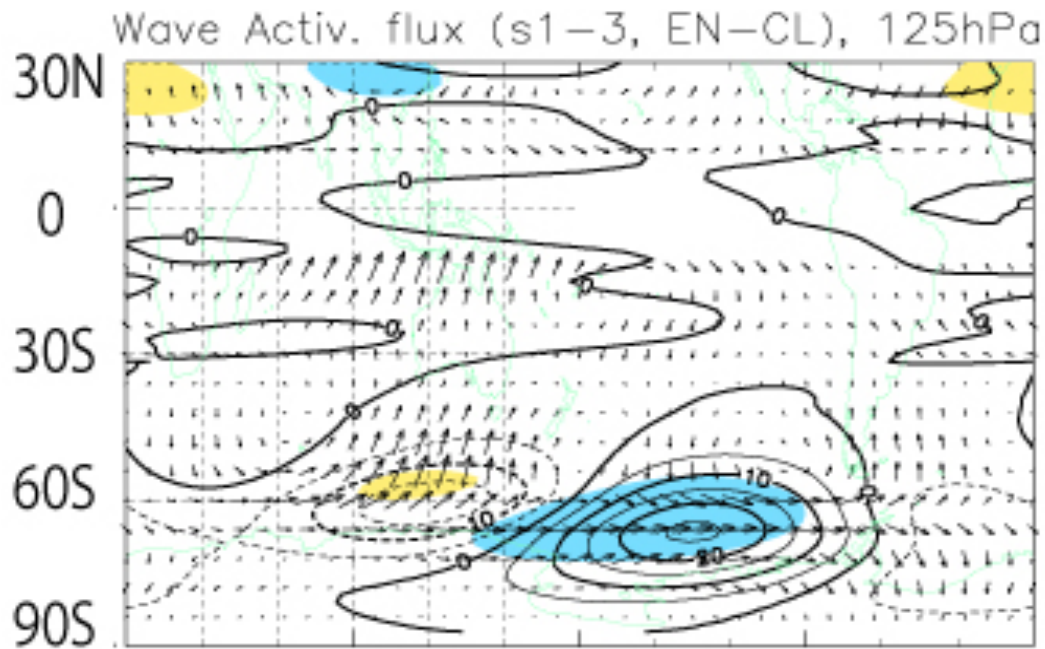


Oct-LN-CLIM Zonal Mean Wave Activ. Flux (s1-3)



Oct-EN-CLIM Zonal Mean Wave Activ. Flux (s1-3)

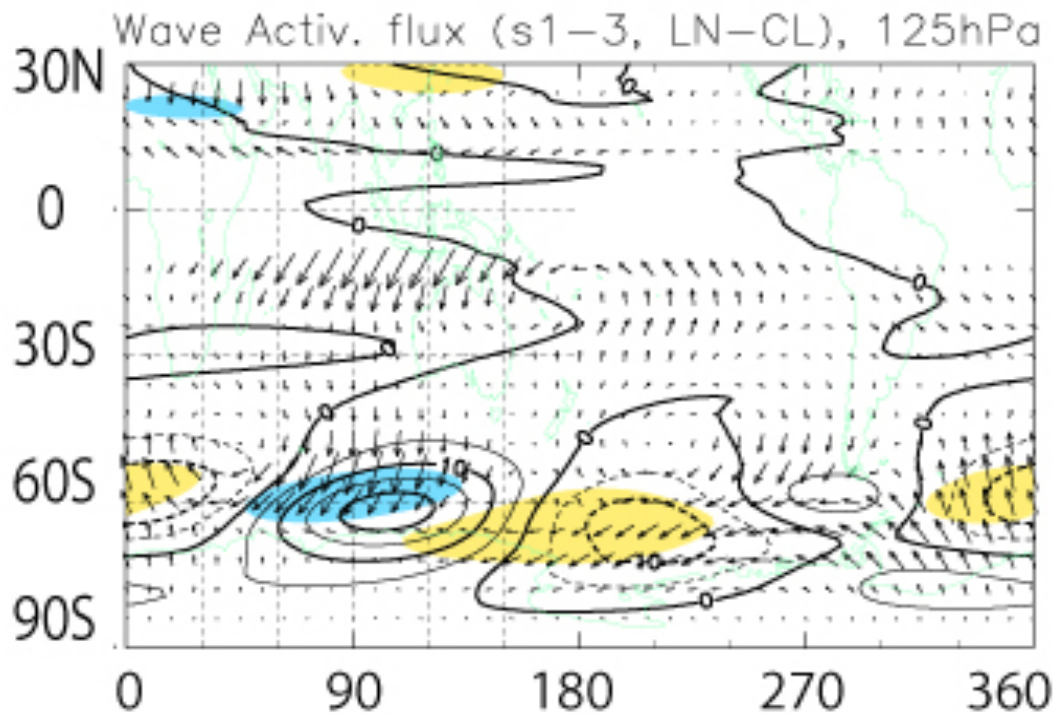




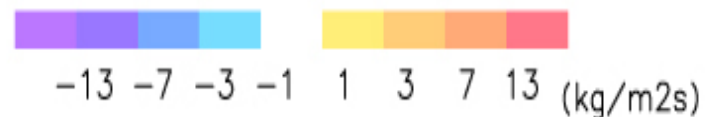
Mid-October

125 hPa

EN - Climatology



LN - Climatology



ENSO effects in MAM, JJA, SON, DJF

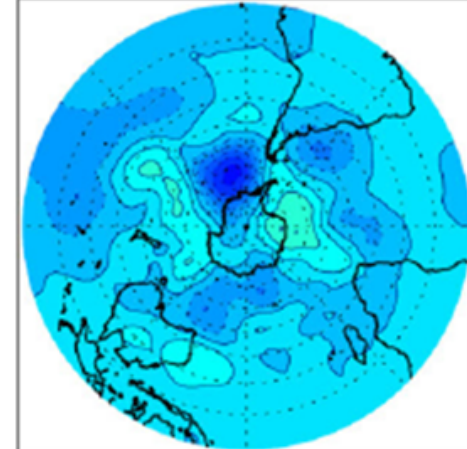
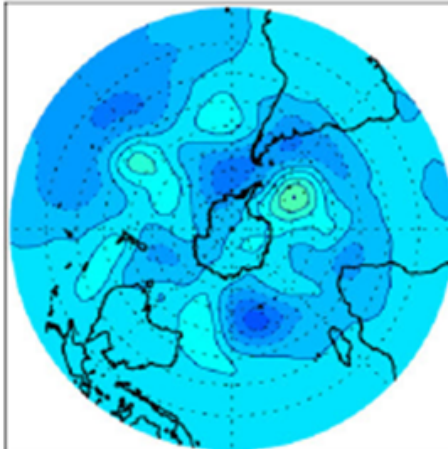
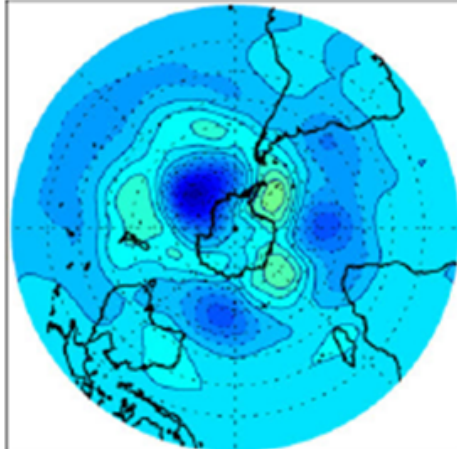
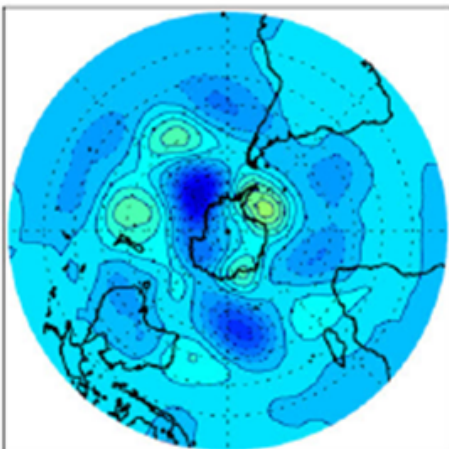
La Nina: stronger polar vortex
stronger planetary wave one
enhanced Ross Sea low

JJA

SON

DJF

MAM



LN – EN

150 hPa Z' (20 m)

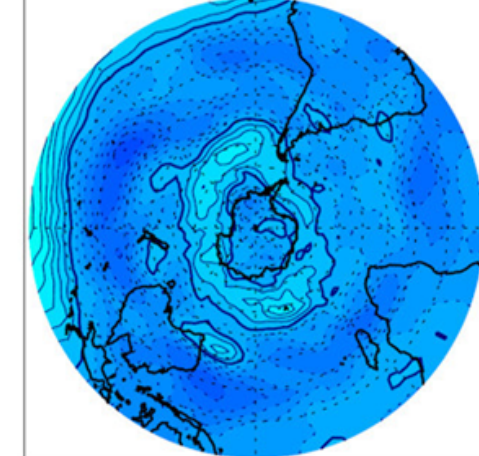
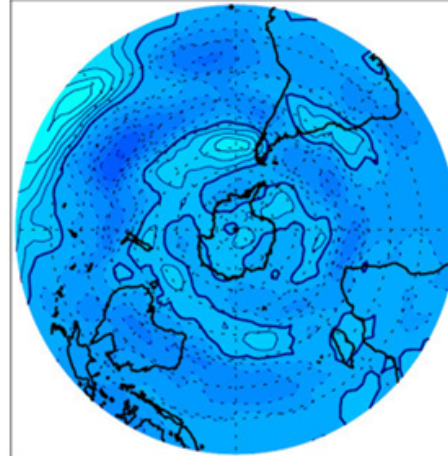
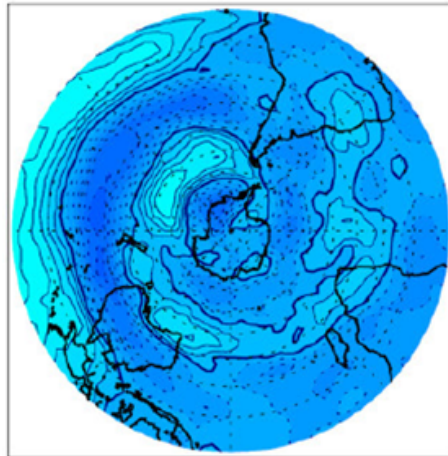
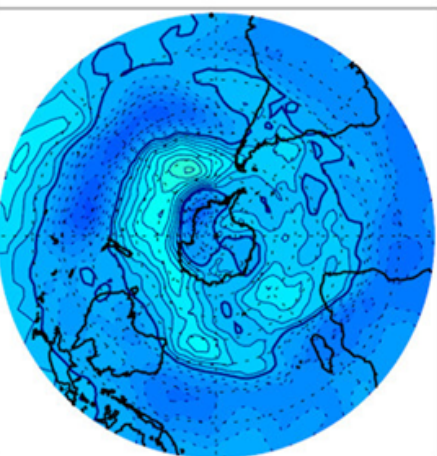
ECMWF 1985-2009

JJA

SON

DJF

MAM



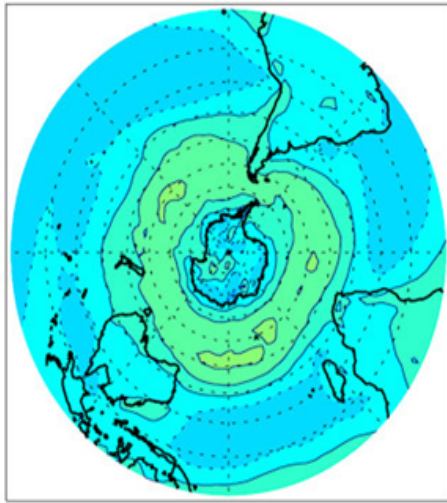
LN – EN

150 hPa zonal wind differences

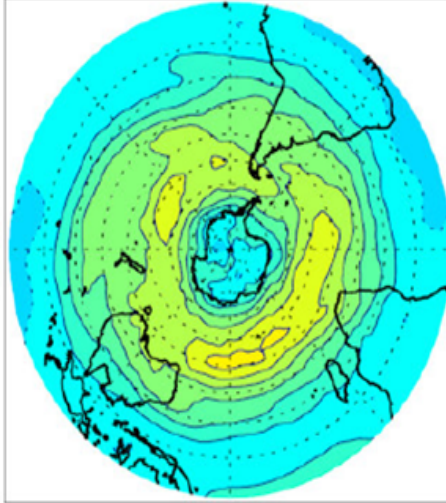
(2 m/s)

Zonal wind speed (contour 5 m/s) during La Nina and SON using daily ECMWF data from 1985-2009.

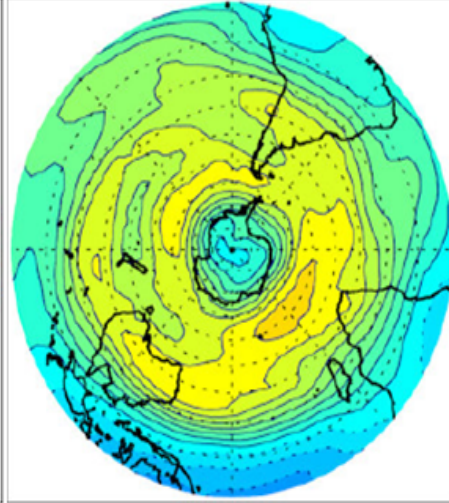
1000 hPa



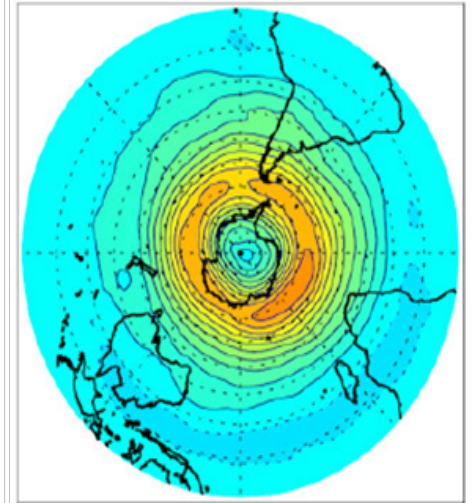
500 hPa



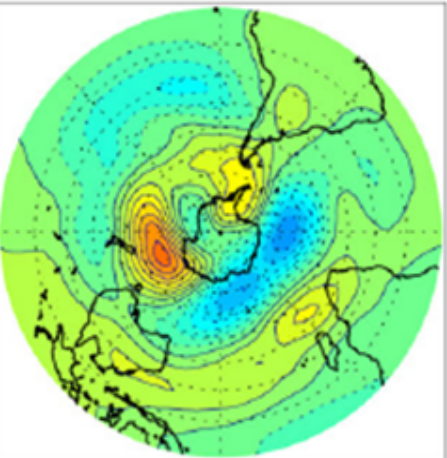
150 hPa



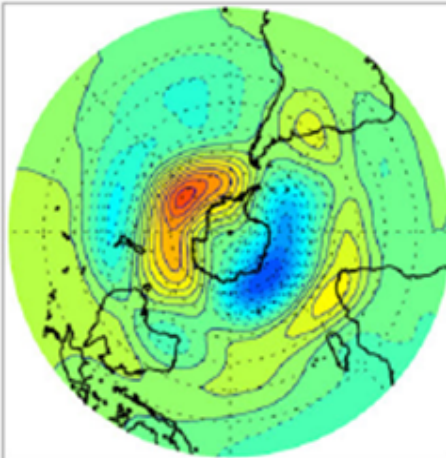
50 hPa



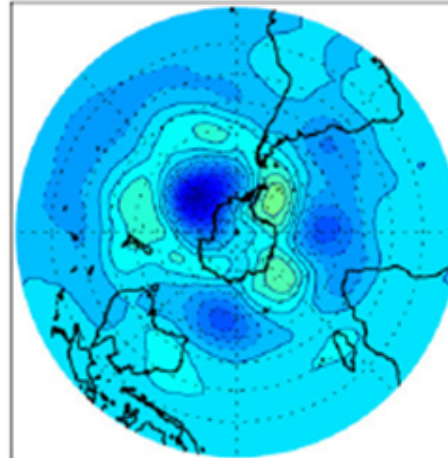
La Nina 150 hPa Z'



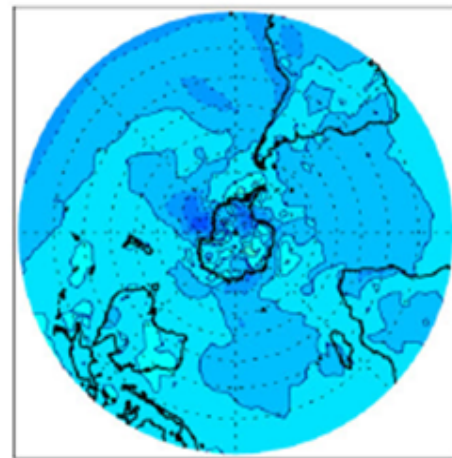
El Nino 150 hPa Z'



LN - EN 150 hPa Z'



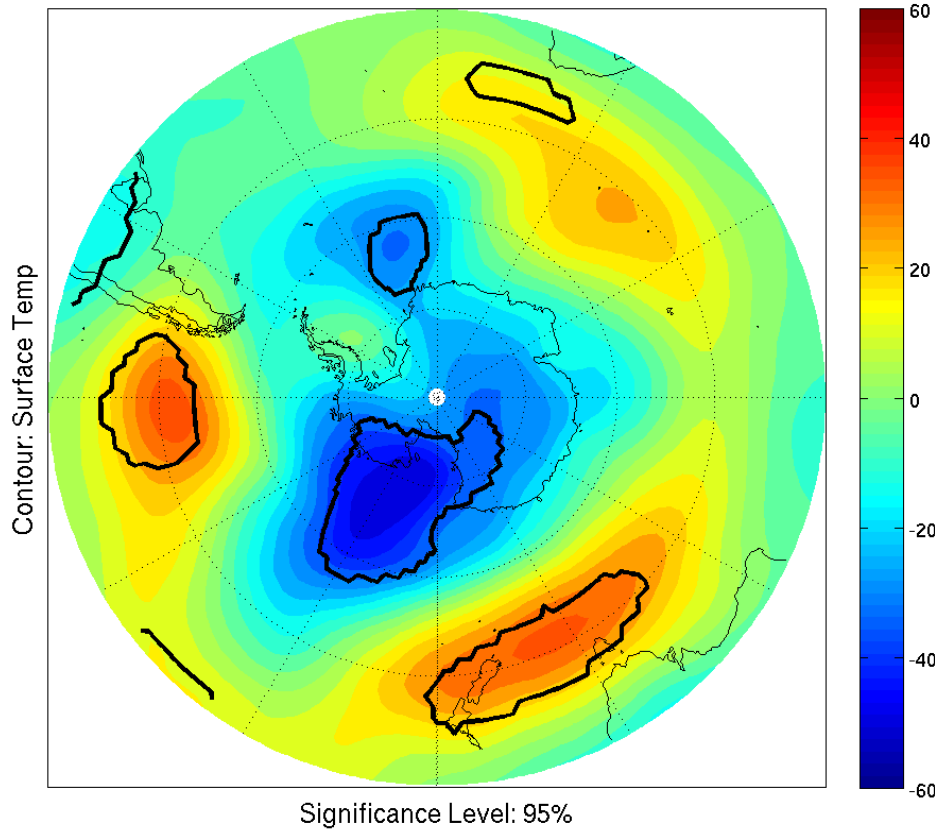
LN - EN 1000 hPa T'



3. Cold East Antarctica during La Nina

* UWNMS simulations Dec. 1-9, 1988

La Nina - Normal



La Nina - Normal

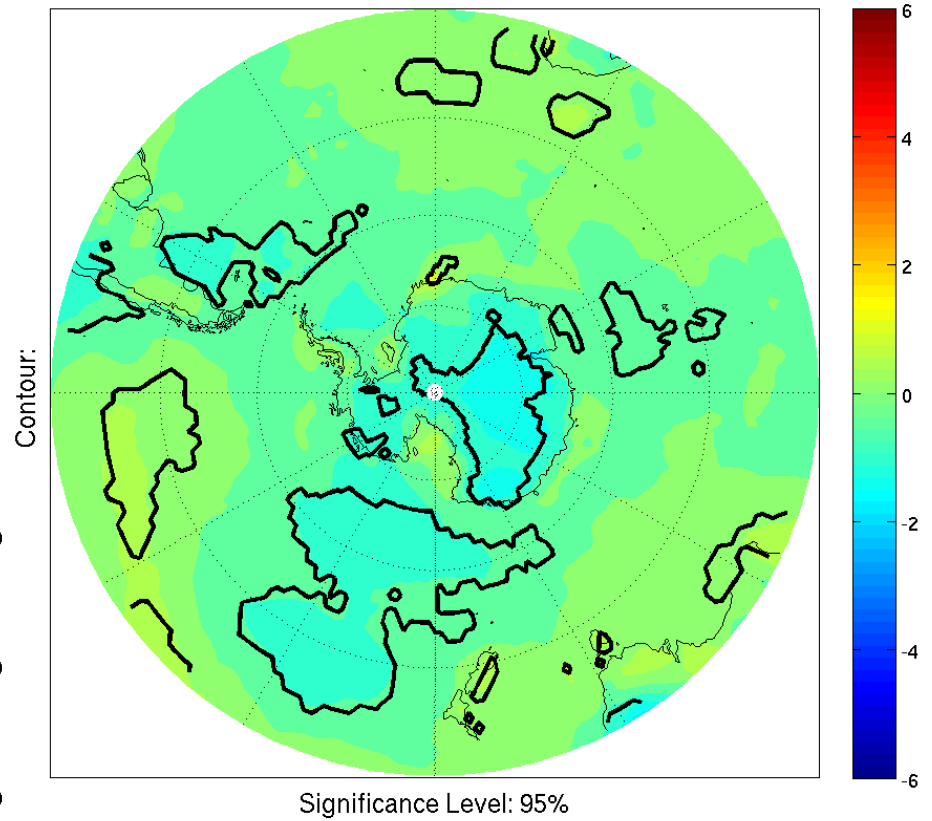
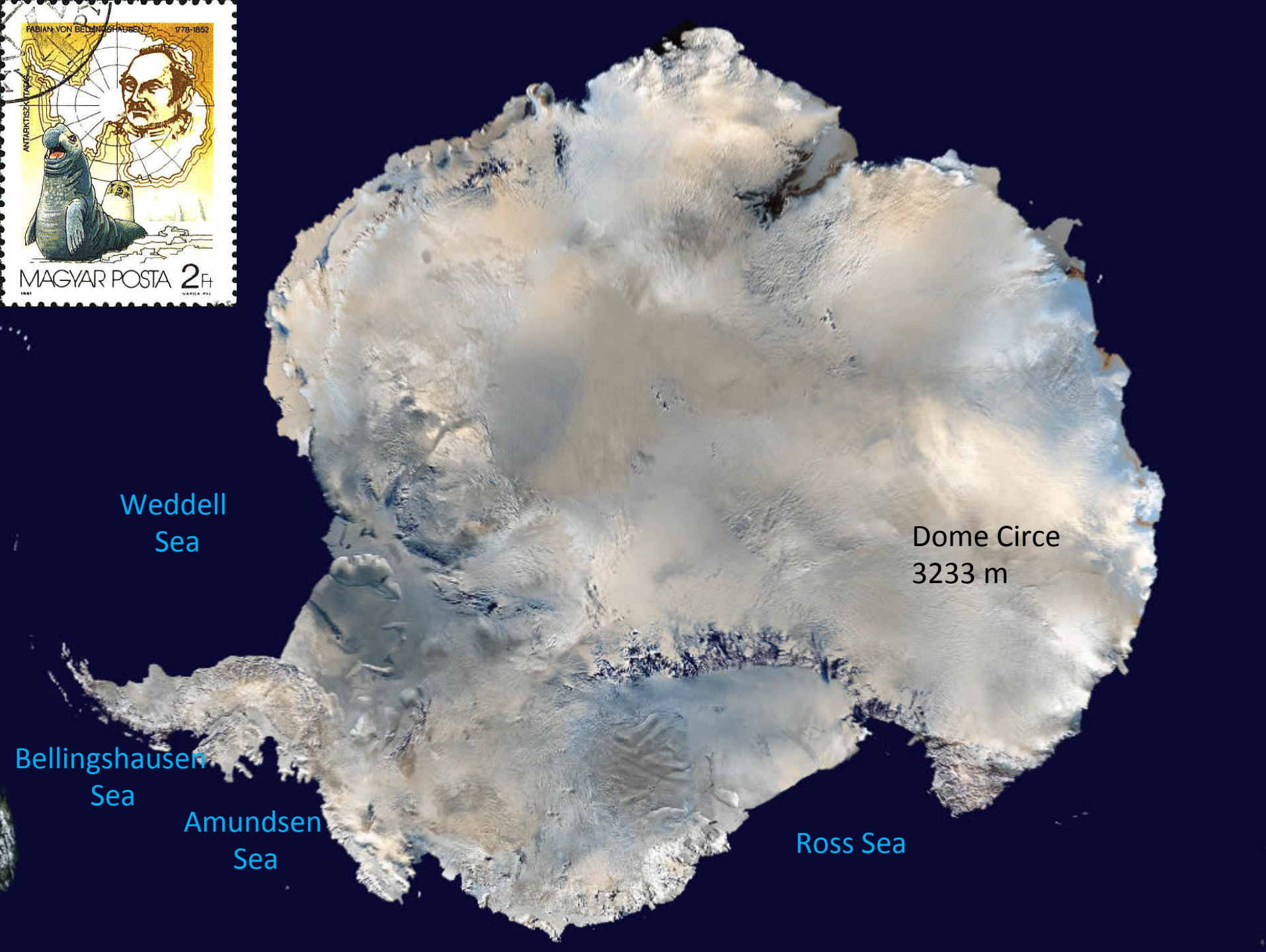


Figure 98: 500 hPa height anomalies for ONI during NDJ of La Niña.

Figure 99: 2 meter temperature anomalies for ONI during NDJ of La Niña.

(Welhouse, 2011)



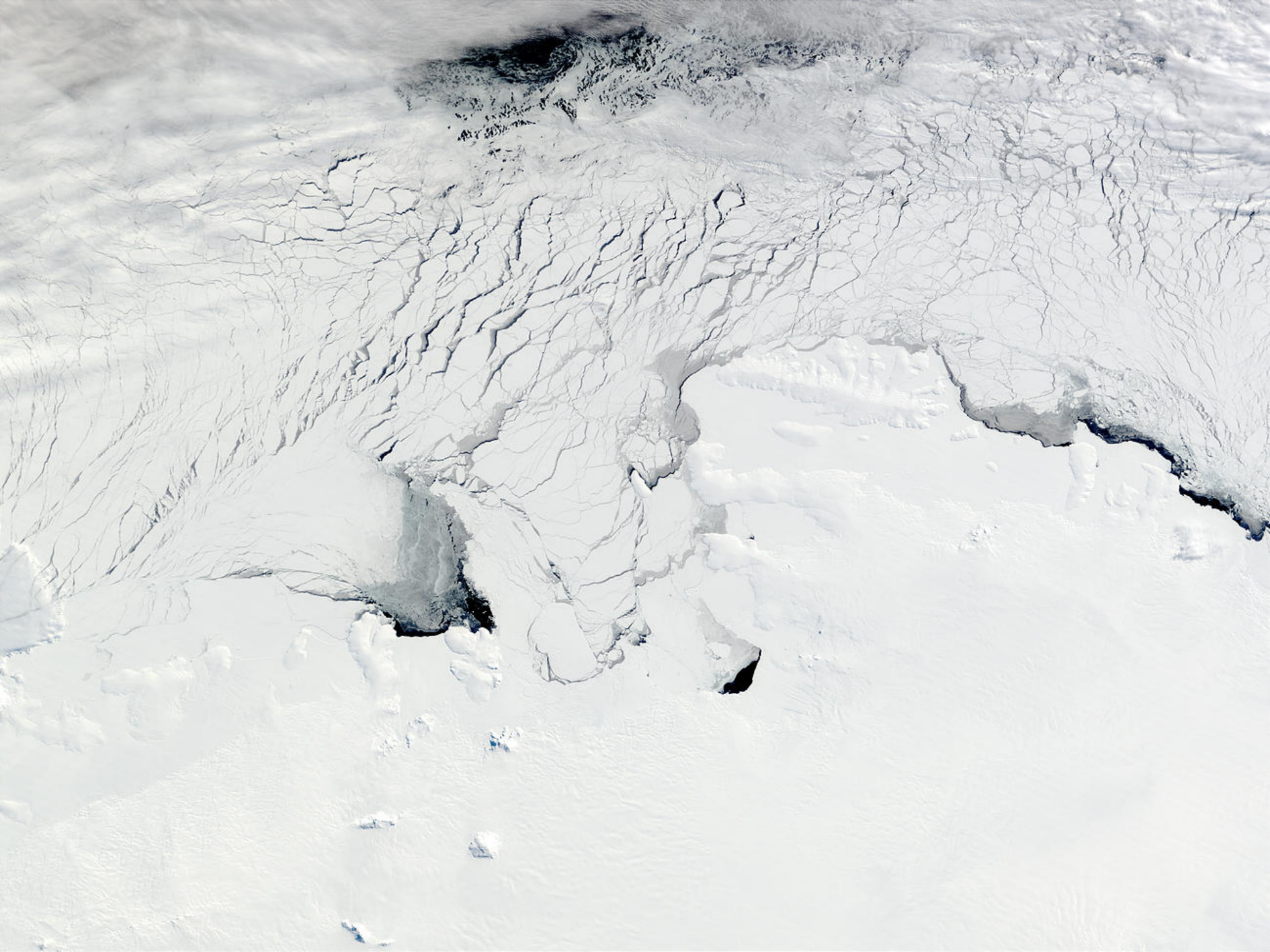
Weddell
Sea

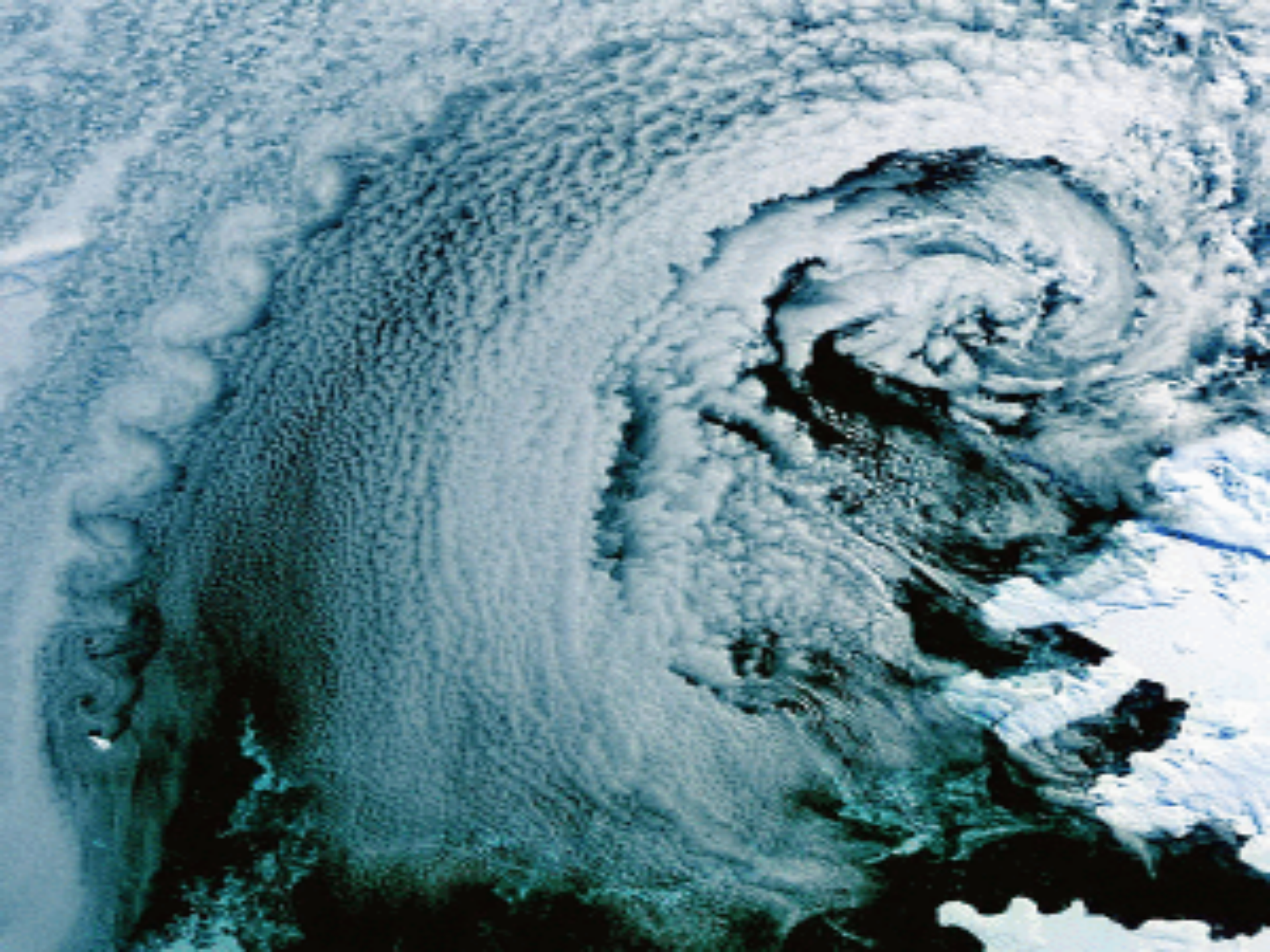
Dome Circe
3233 m

Bellingshausen
Sea

Amundsen
Sea

Ross Sea





University of Wisconsin Nonhydrostatic Modeling System

Synoptic Case Studies December 1 - 3 and 5 - 9, 1988

ECWMF initialization

No inflow forcing

Grid Points: 152 x 152 x 60

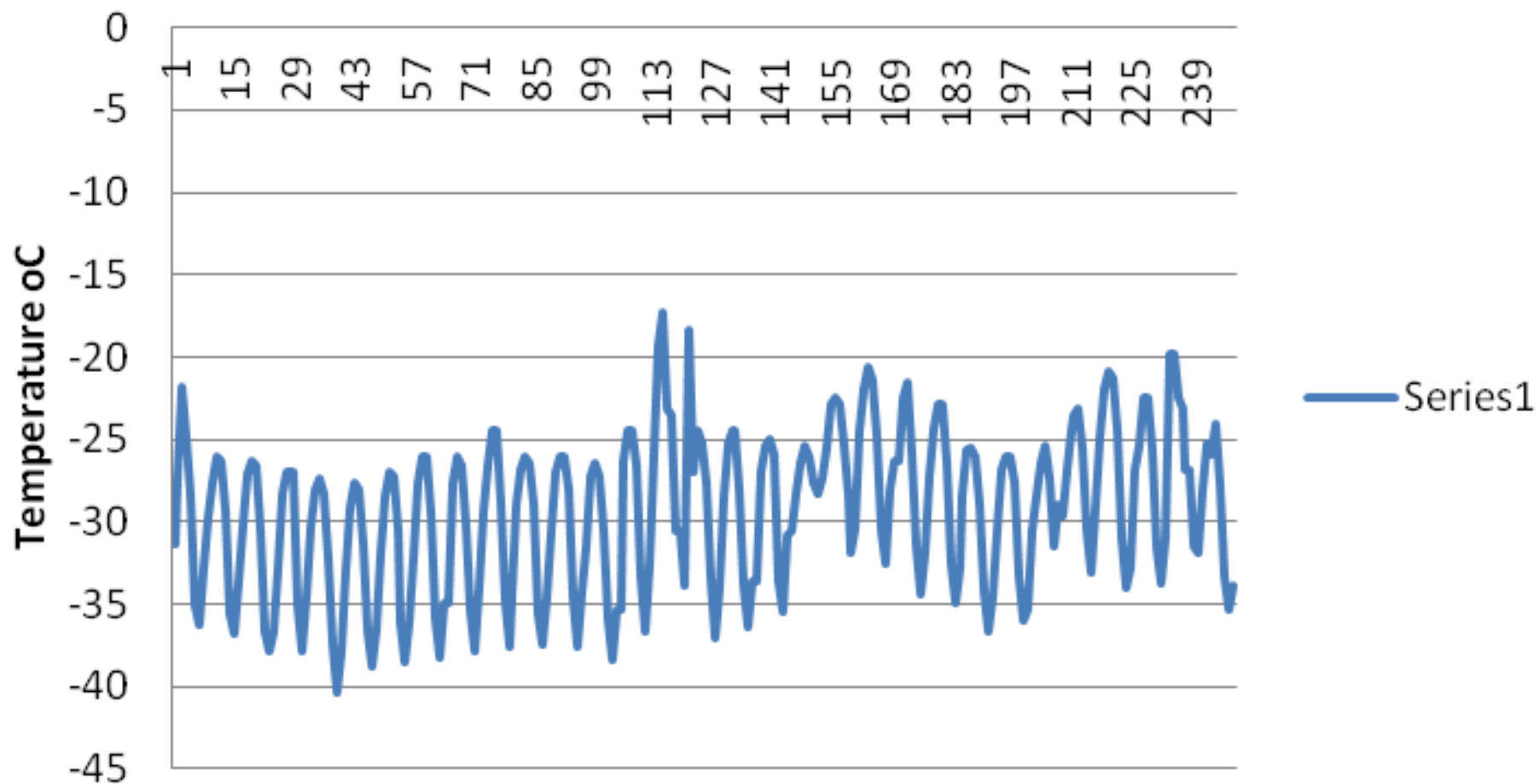
Outer grid: 100 km x 100 km x 300 m

Inner grid: 20 km x 20 km x 300 m

Model top: 18 km

Vis5d movies of output

Dome C Temperature December 1988



40 m/s isosurface

Dec 1-3 1988

5 km pressure (4 hPa contour interval) outer grid

Dec 1-3 1988

10 km pressure (4 hPa contour interval) outer grid

Dec 1-3 1988

PV section

GM – Date Line

Dec 1-3 1988



PV = - 2 isosurface colored by Potential Temperature (yellow ~ 340 K) Dec 1-3 1988

4 km streamfunction

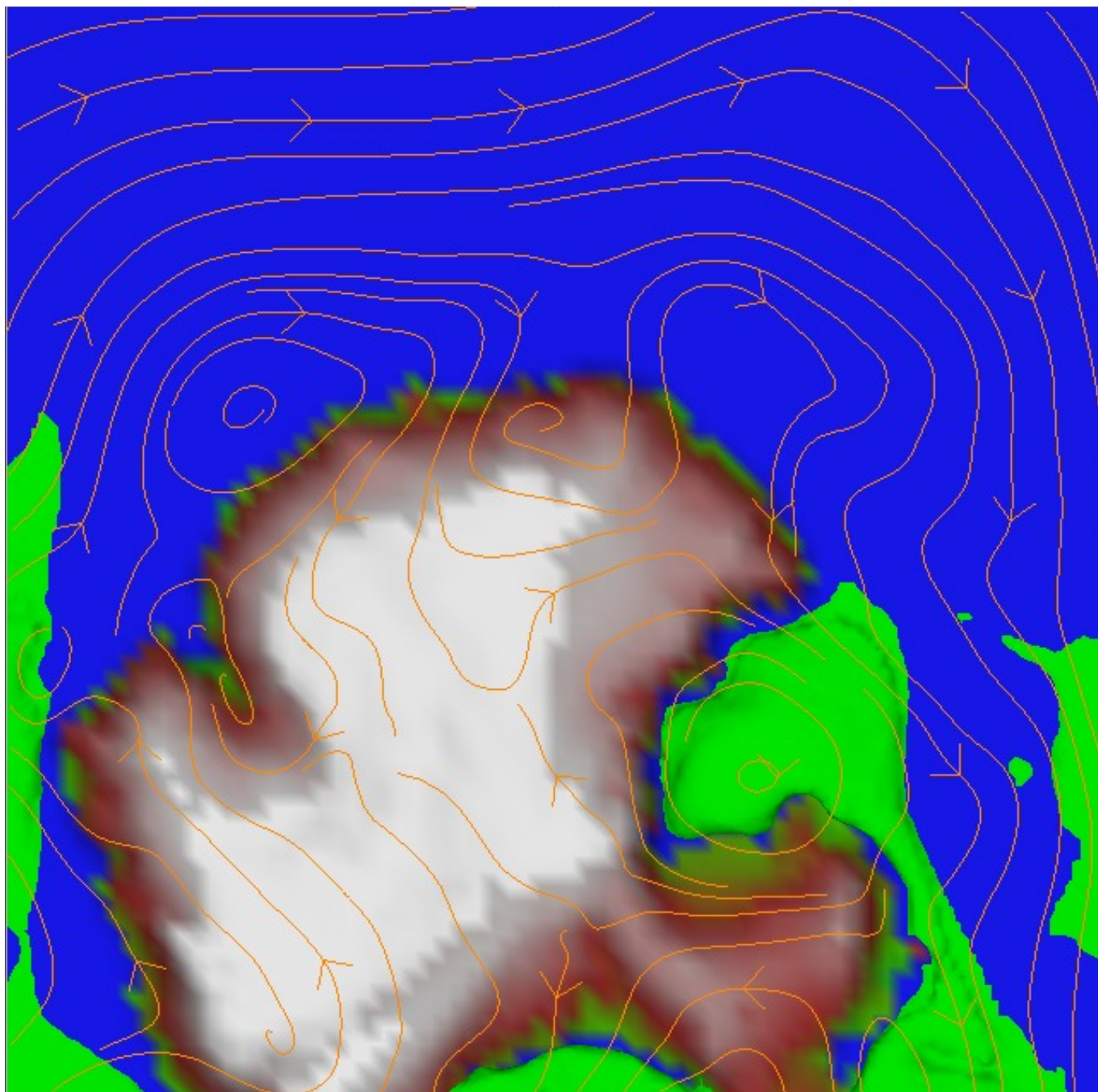
40 km x 40 km x 300 m

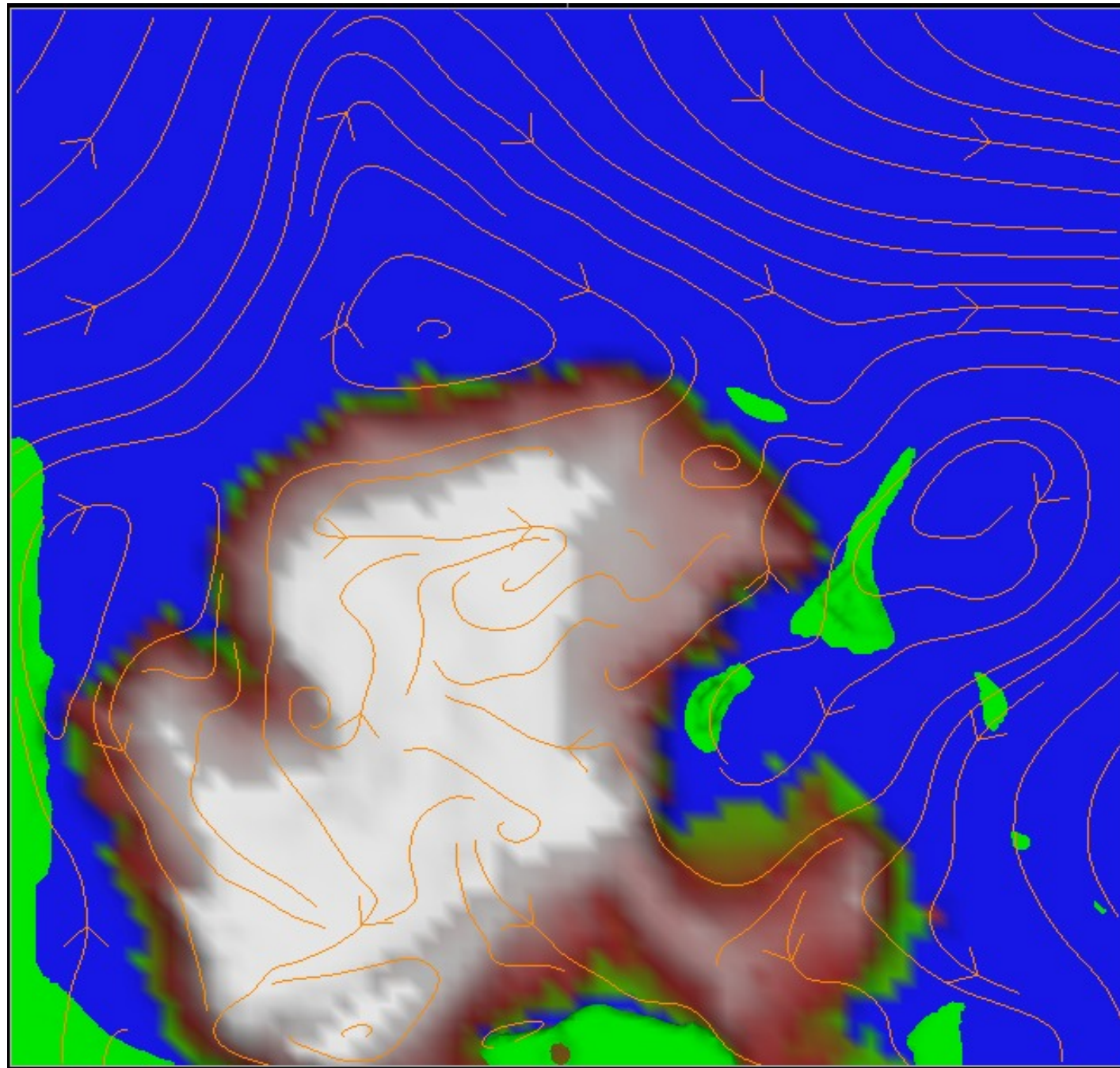
Dec 5-9 1988

4 km streamfunction

Tuesday Dec 6 1988

341





Conclusions

During La Nina:

column ozone maximum shifted westward $\sim 50^\circ$
stronger poleward wave activity flux from Indonesia
stronger planetary wave one
stronger polar vortex
colder East Antarctica
enhanced Ross Sea low